

# Three-Body Forces and the Structure of Exotic Nuclei

Jason D. Holt

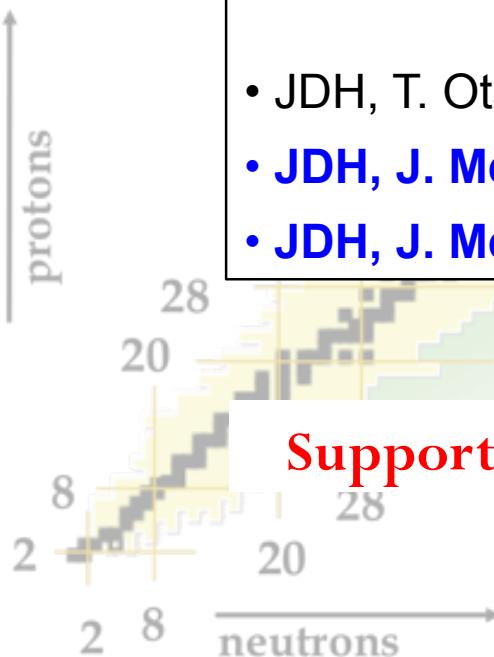


TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

126

Based on

- JDH, T. Otsuka, A. Schwenk, and T. Suzuki, JPG (2012)
- **JDH, J. Menendez, A. Schwenk, PRL (2013)**
- **JDH, J. Menendez, A. Schwenk, in prep.**



# Outline

Goal: understand the role of 3N forces for structure of medium-mass exotic nuclei

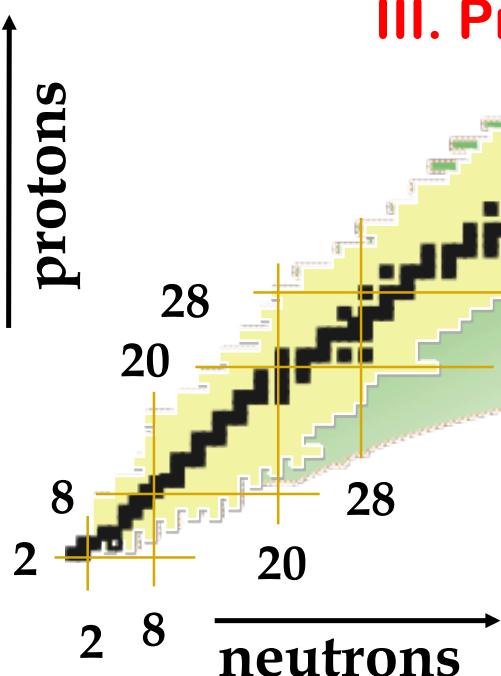
- What are the limits of nuclear existence?
- How do magic numbers form and evolve?

I. Microscopic approach: theoretical inputs consistently from NN+3N

II. Shell evolution in calcium

- a) Pairing gaps
- b) Spectra

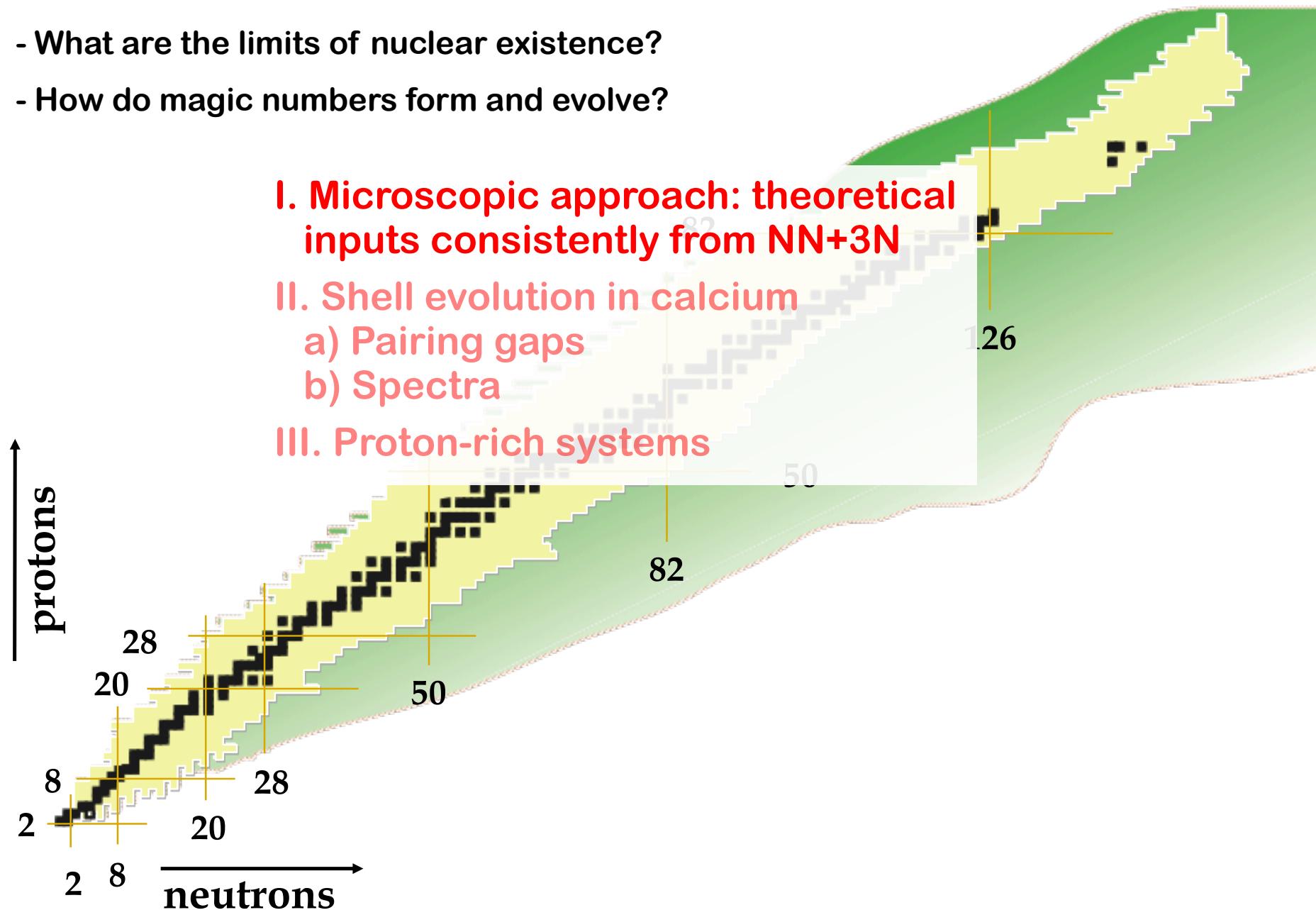
III. Proton-rich systems



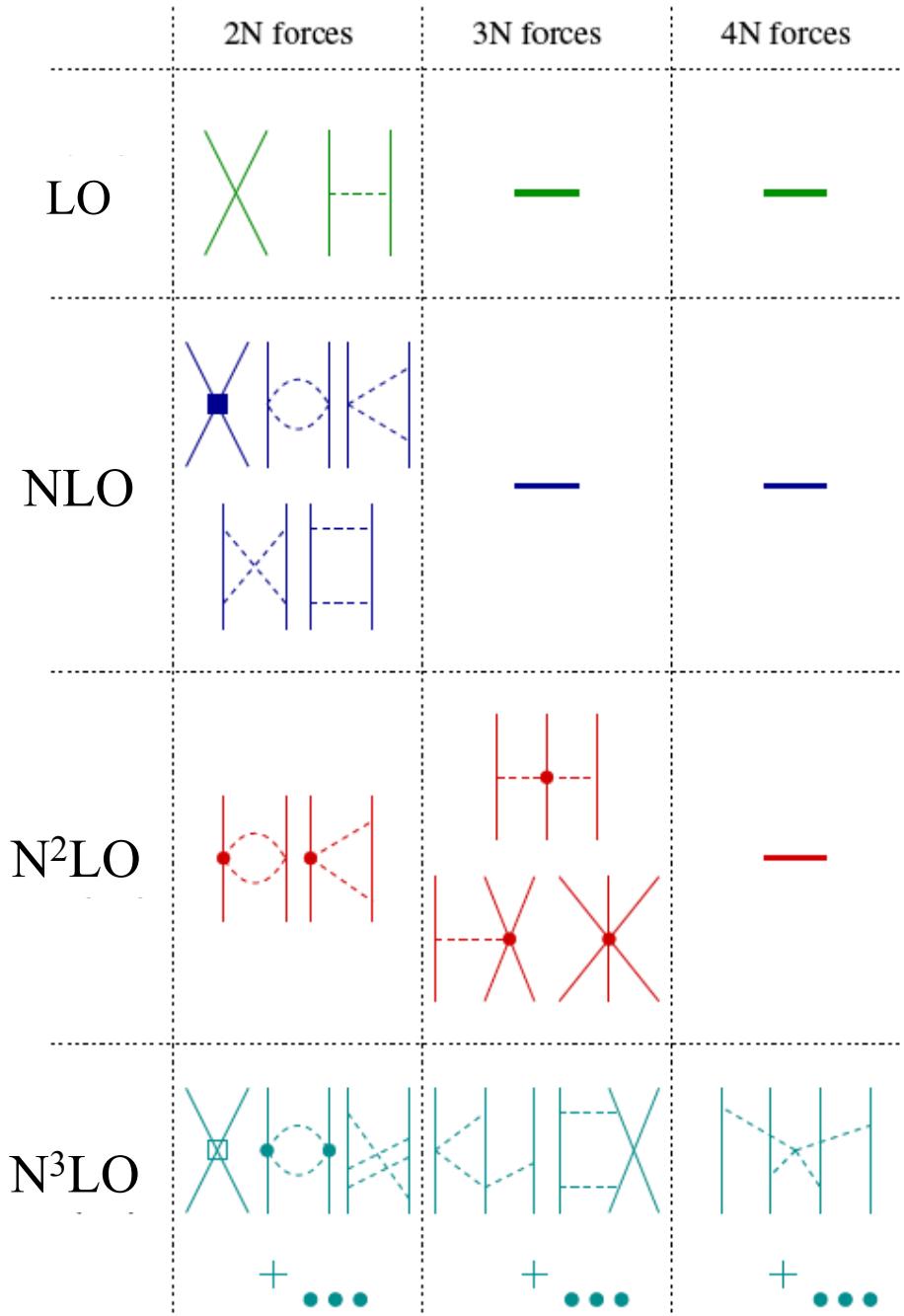
# Outline

Goal: understand the role of 3N forces for structure of medium-mass exotic nuclei

- What are the limits of nuclear existence?
- How do magic numbers form and evolve?



# Chiral Effective Field Theory: Nuclear Forces



Nucleons interact via pion exchanges and contact interactions

Hierarchy:  $V_{\text{NN}} > V_{\text{3N}} > \dots$

Consistent treatment of NN, 3N, ... electroweak operators

Couplings fit to experiment once

Evolve to **low-momentum**  $V_{\text{low } k}$

3N constants fit to properties of light nuclei at low momentum

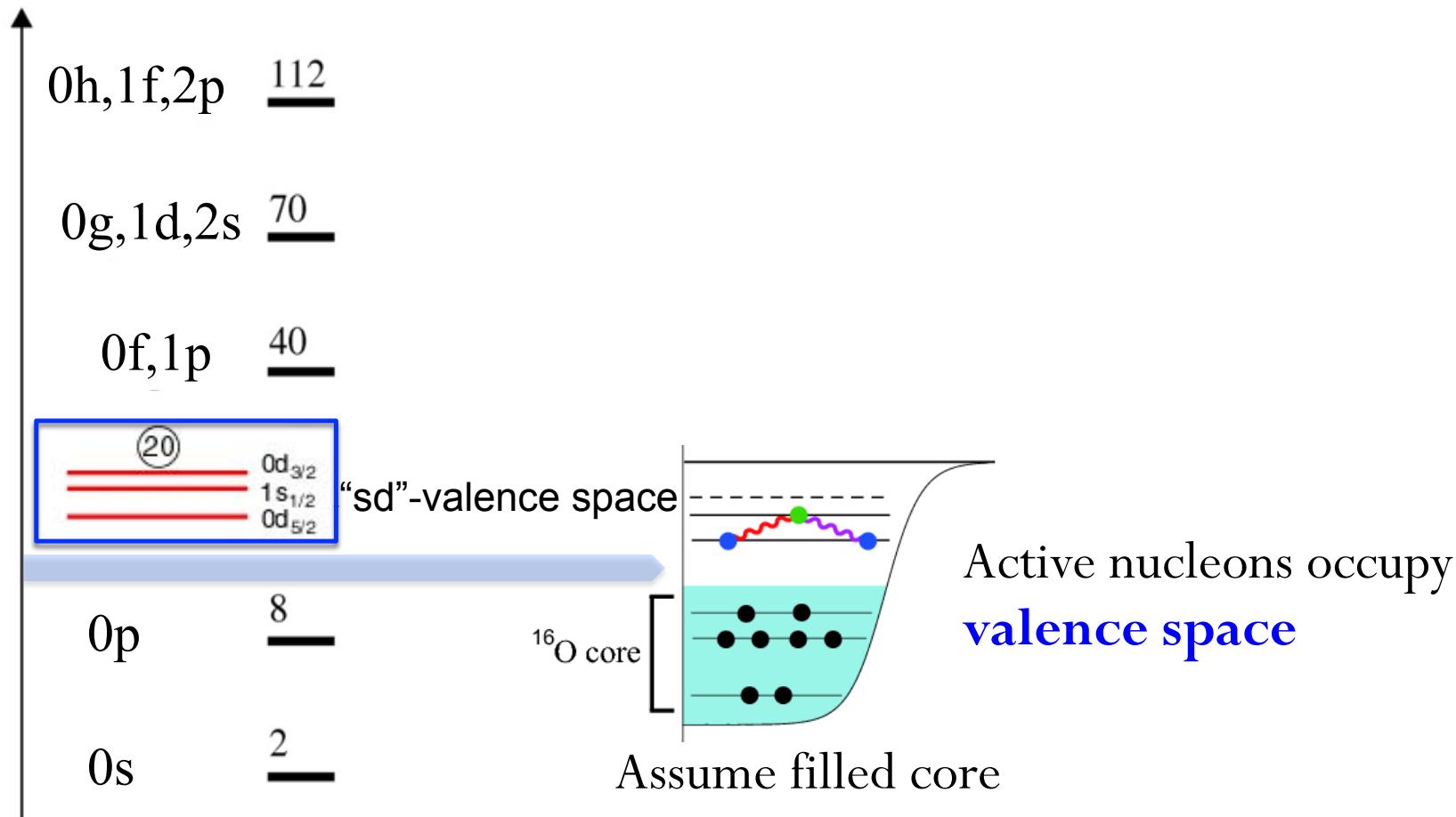
Weinberg, van Kolck, Kaplan, Savage, Wise, Epelbaum, Kaiser, Meissner,...

# Solving the Nuclear Many-Body Problem

Nuclei understood as many-body system starting from closed shell, add nucleons

Interaction and energies of valence space orbitals from  $V_{\text{low } k}$

**Does not reproduce experimental data**



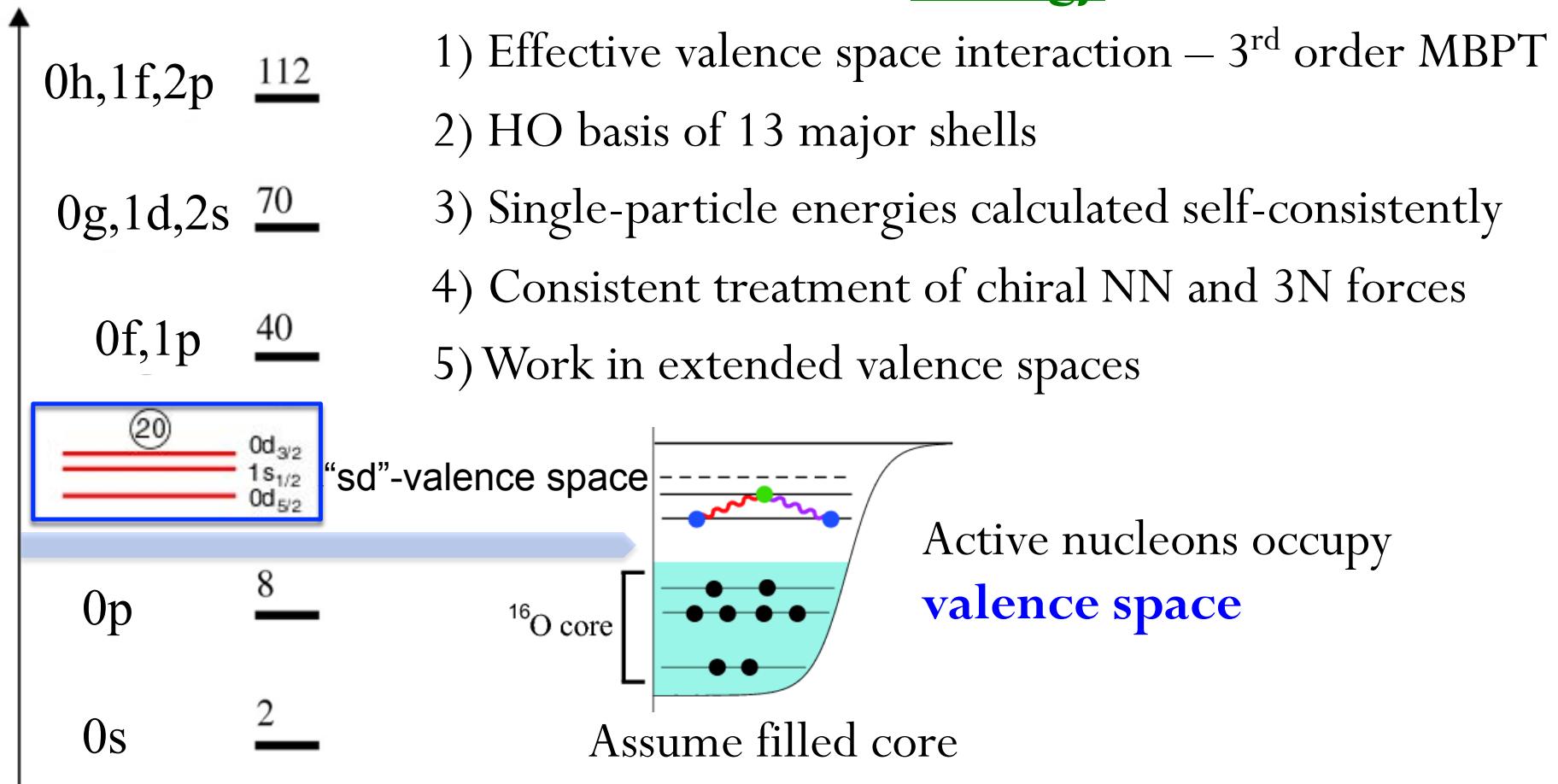
# Solving the Nuclear Many-Body Problem

Nuclei understood as many-body system starting from closed shell, add nucleons

Interaction and energies of valence space orbitals from  $V_{\text{low } k}$

Does not reproduce experimental data – **allow explicit breaking of core**

## Strategy



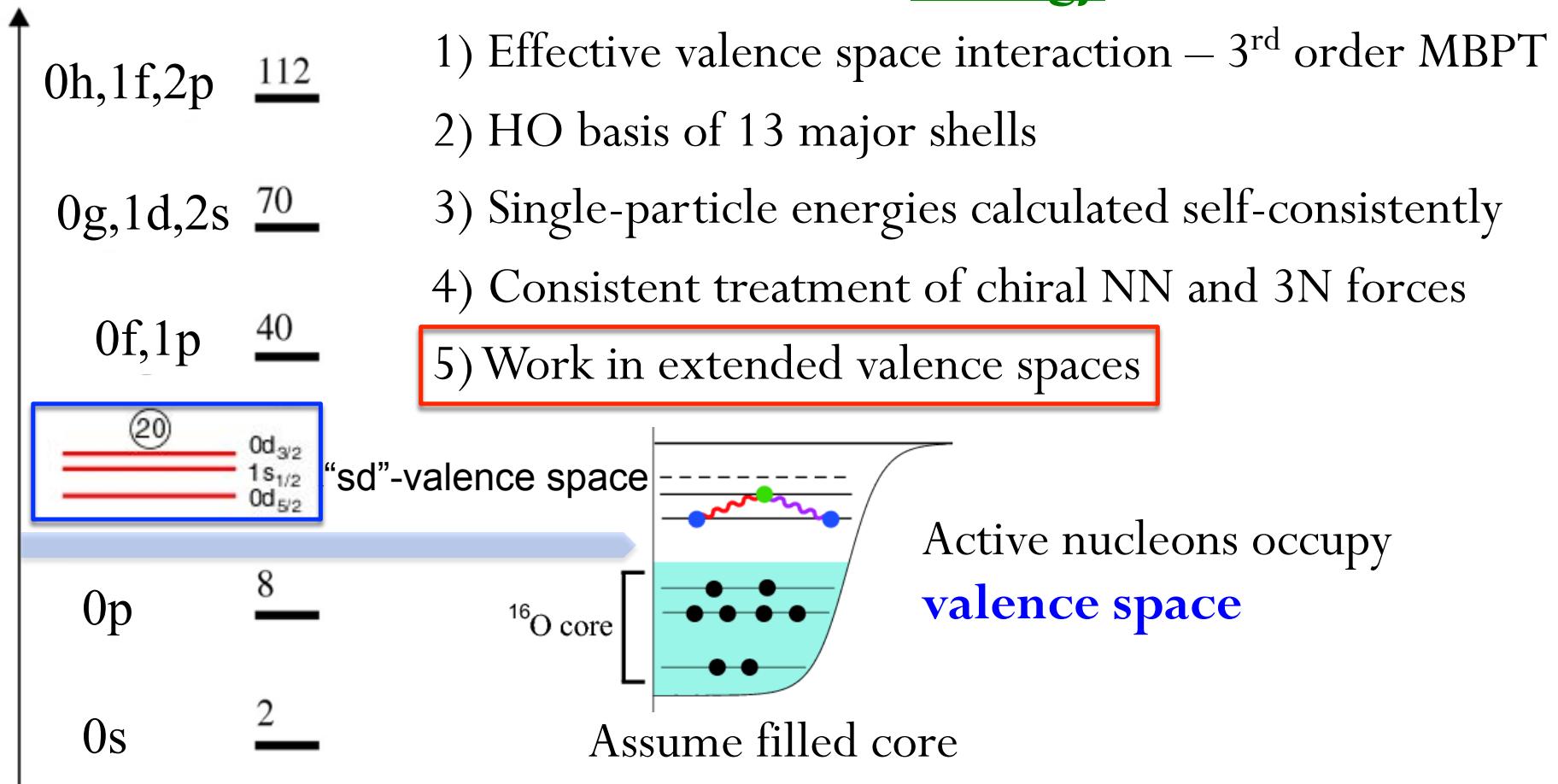
# Solving the Nuclear Many-Body Problem

Nuclei understood as many-body system starting from closed shell, add nucleons

Interaction and energies of valence space orbitals from  $V_{\text{low } k}$

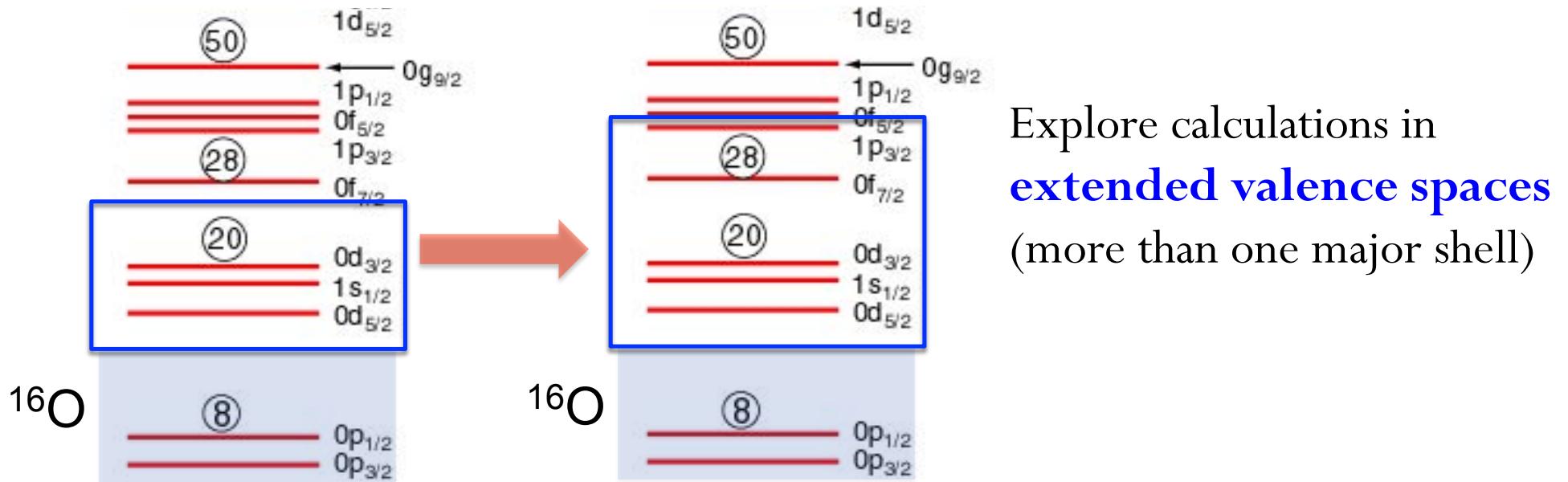
Does not reproduce experimental data – **allow explicit breaking of core**

## Strategy



# Extended Valence Spaces

**Philosophy:** diagonalize in largest possible valence space (where orbits relevant)



When do extended-space orbits impact exotic nuclei

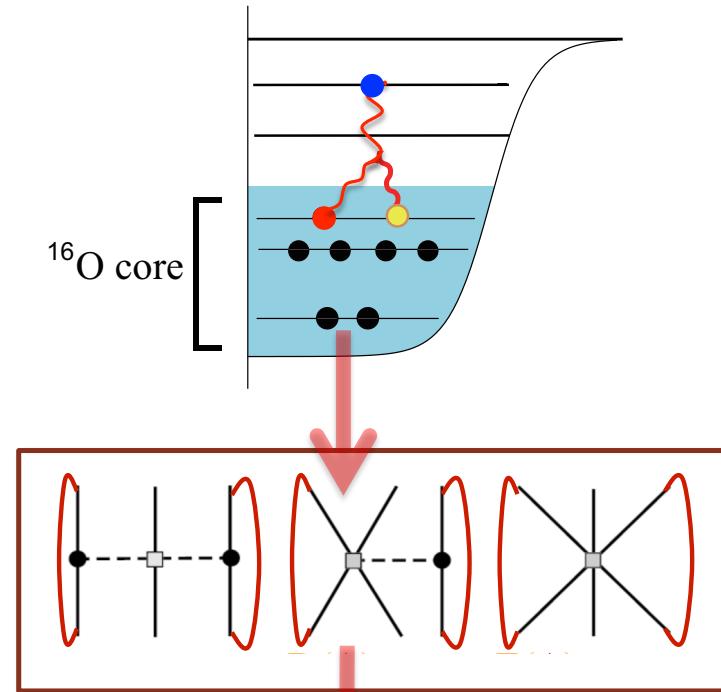
**Caution:** possible center-of-mass contamination

Explore calculations in  
**extended valence spaces**  
(more than one major shell)

# 3N Forces for Valence-Shell Theories

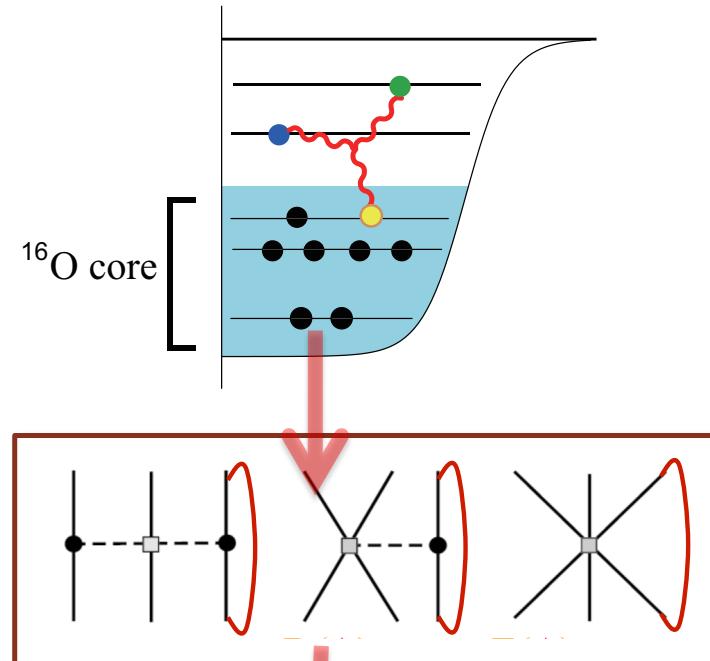
Normal-ordered 3N: contribution to valence nucleon interactions

## Effective one-body



$$\langle a | V_{3N,\text{eff}} | a' \rangle = \frac{1}{2} \sum_{\alpha\beta=\text{core}} \langle \alpha\beta a | V_{3N} | \alpha\beta a' \rangle$$

## Effective two-body



$$\langle ab | V_{3N,\text{eff}} | a'b' \rangle = \sum_{\alpha=\text{core}} \langle \alpha ab | V_{3N} | \alpha a'b' \rangle$$

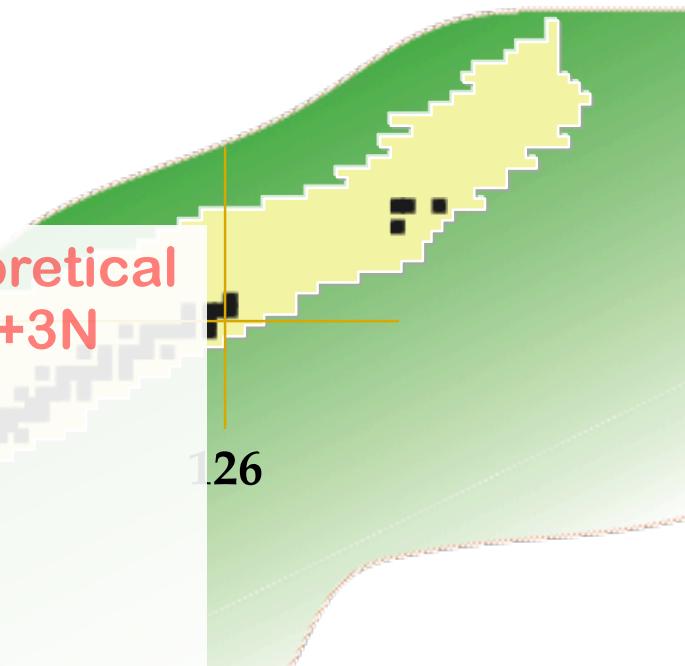
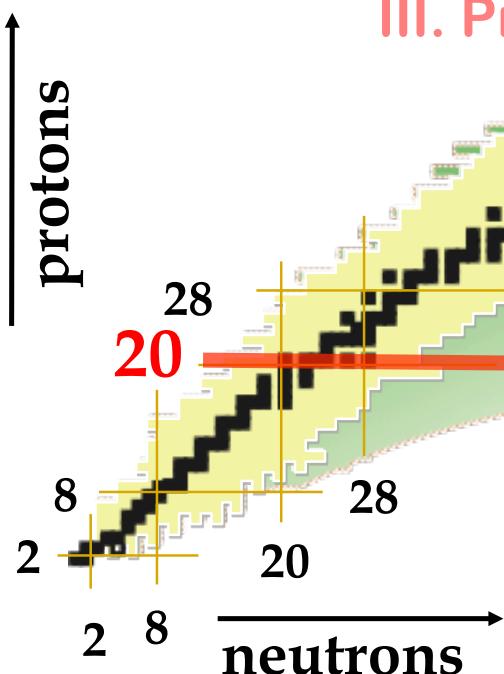
Combine with microscopic NN (**Third Order**): no empirical adjustments

# Shell Formation/Evolution in Calcium Isotopes

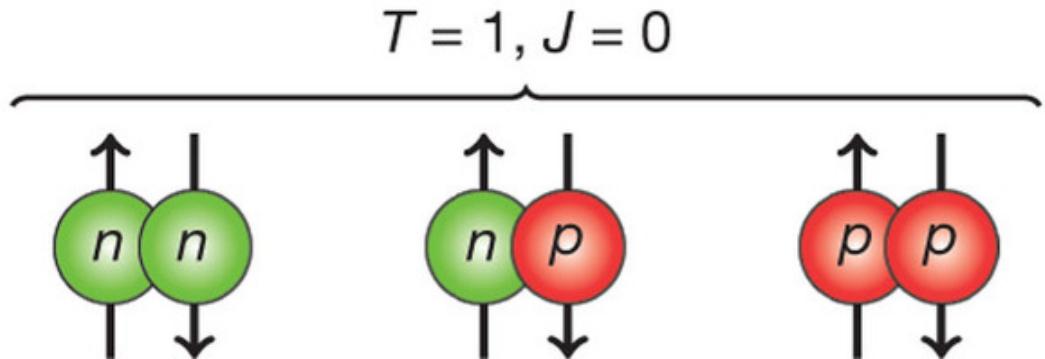
Goal: understand the role of 3N forces for structure of medium-mass exotic nuclei

- What are the limits of nuclear existence?
- How do magic numbers form and evolve?

- I. Microscopic approach: theoretical inputs consistently from NN+3N
- II. Shell evolution in calcium
  - a) Pairing gaps
  - b) Spectra
- III. Proton-rich systems



# Nuclear Pairing



Pairing of even number of nucleons – even/odd staggering

Pairing gaps deduced from **3-point mass difference**:

$$\Delta_n^{(3)} = \frac{(-1)^N}{2} [BE(N+1,Z) + BE(N-1,Z) - 2BE(N,Z)]$$

Allows comparison with experiment

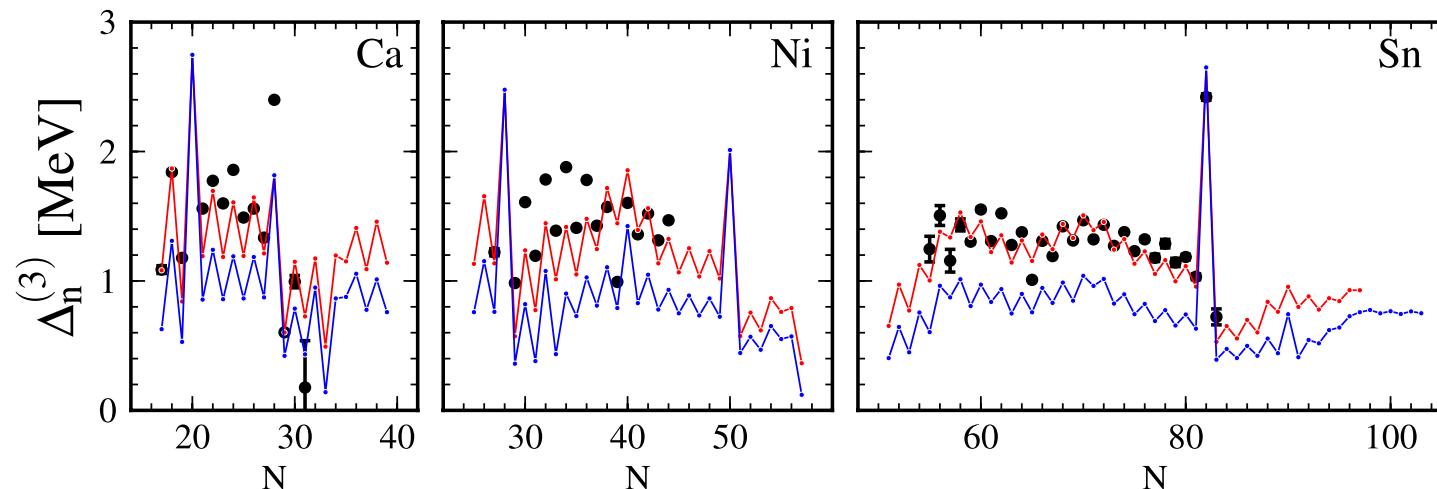
Relative peak in  $\Delta_n^{(3)}$  indicates **shell closure**

- **additional tool to evaluate shell evolution**

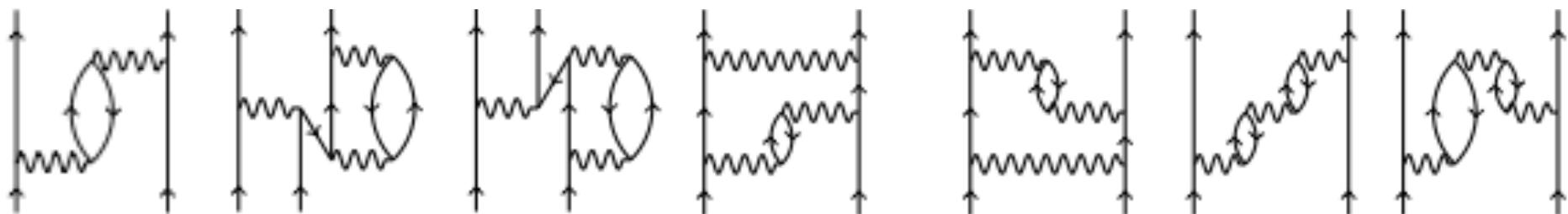
# Pairing in EDF with 3N Forces

Energy Density Functional calculations: 3N lowers gaps systematically  $\sim 30\%$

Lesinski, Hebeler, Duguet, Schwenk, JPG (2012)

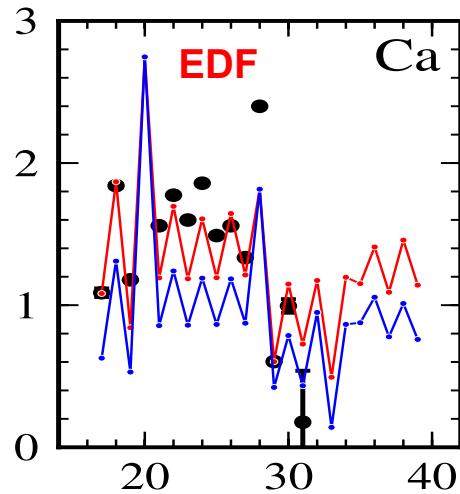


What are the contributions from neglected many-body effects?  
(Core polarization)



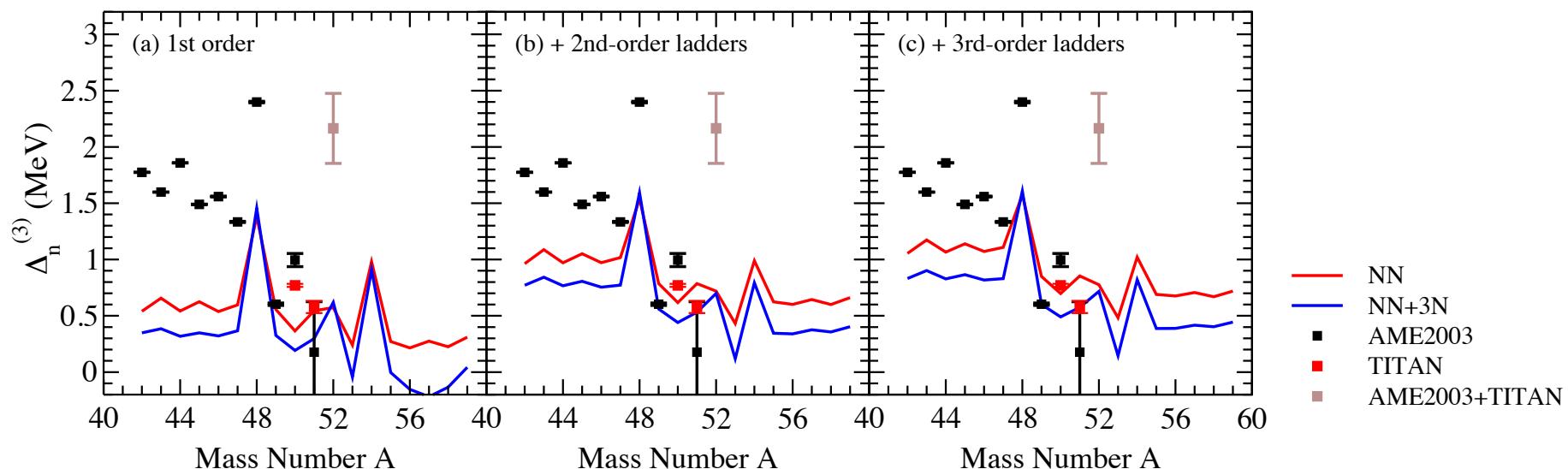
# Pairing in Calcium Isotopes: Ladders

Compare with  $\Delta_n^{(3)}$  calculated from microscopic NN+3N in calcium



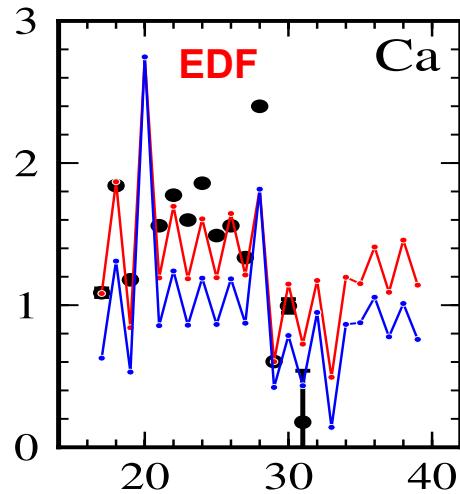
- HFB iterates ladders microscopically in pairing channel
- Compare with *pp, hh ladders to 3<sup>rd</sup> order*
- Improved agreement with experiment
- Convergence in order-by-order ladders
- Suppression from 3N forces as in EDF
- Incorrect odd/even staggering

JDH, Menendez, Schwenk, in prep



# Pairing in Calcium Isotopes: Full 3<sup>rd</sup> order

Compare with  $\Delta_n^{(3)}$  calculated from microscopic NN+3N in calcium



## Full 3<sup>rd</sup>-order MBPT

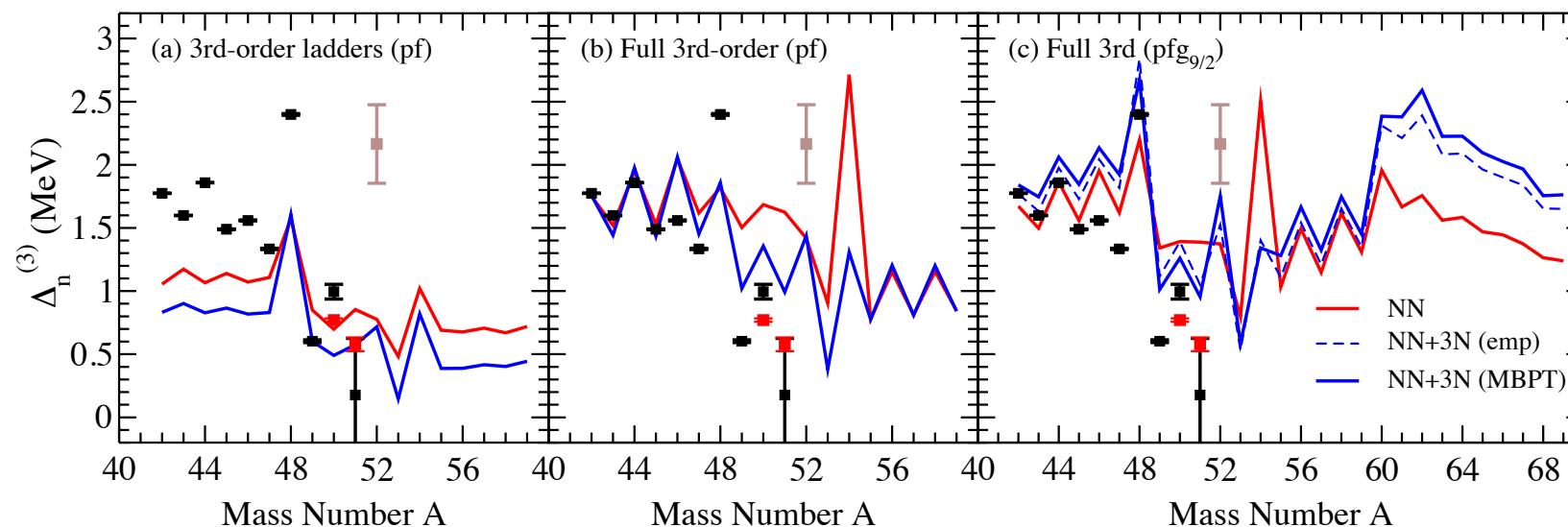
Further increases gaps

Correct odd/even staggering; more pronounced

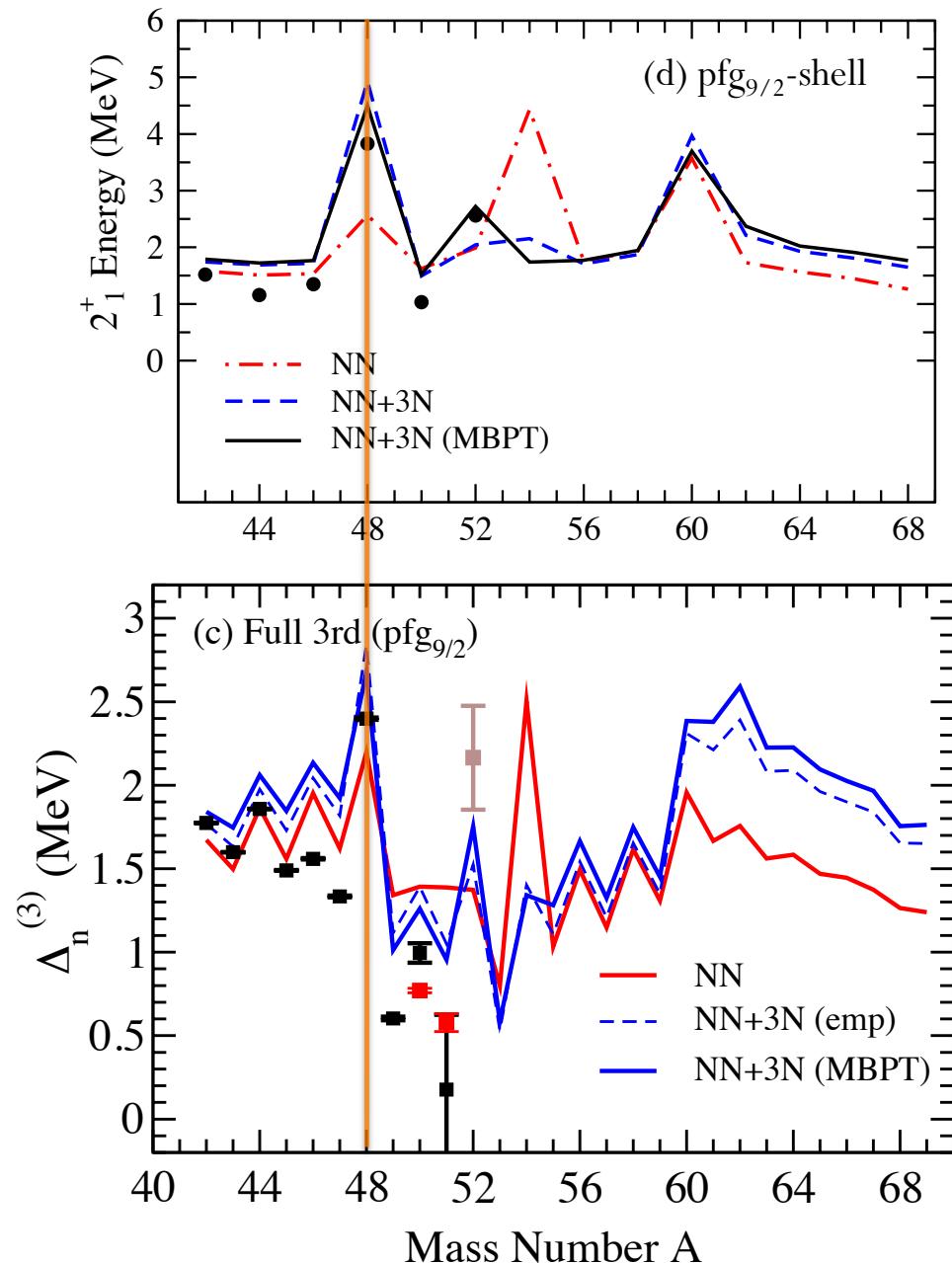
Good experimental reproduction with 3<sup>rd</sup>-order NN+3N

Can account for missing physics in EDF calculations

JDH, Menendez, Schwenk, in prep



# Pairing for Shell Evolution N=28



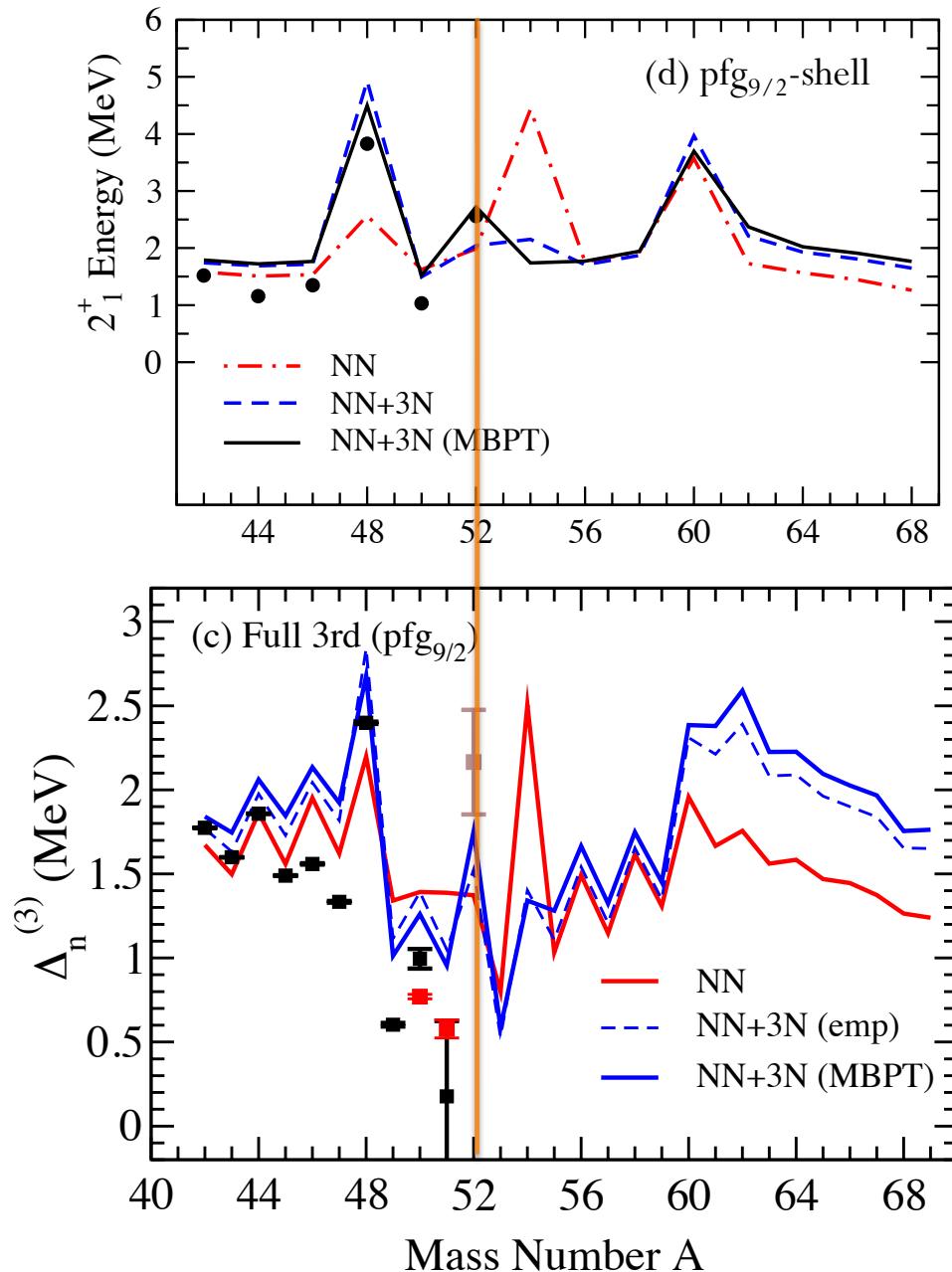
Peak in pairing gaps: complementary signature for shell closure

Compare with  $2^+$  energies for Ca

General agreement with CC predictions [Hagen et al PRL \(2012\)](#)

**N=28:** strong peak, strength overpredicted in both cases

# Pairing for Shell Evolution N=32



Peak in pairing gaps: complementary signature for shell closure

Compare with  $2^+$  energies for Ca

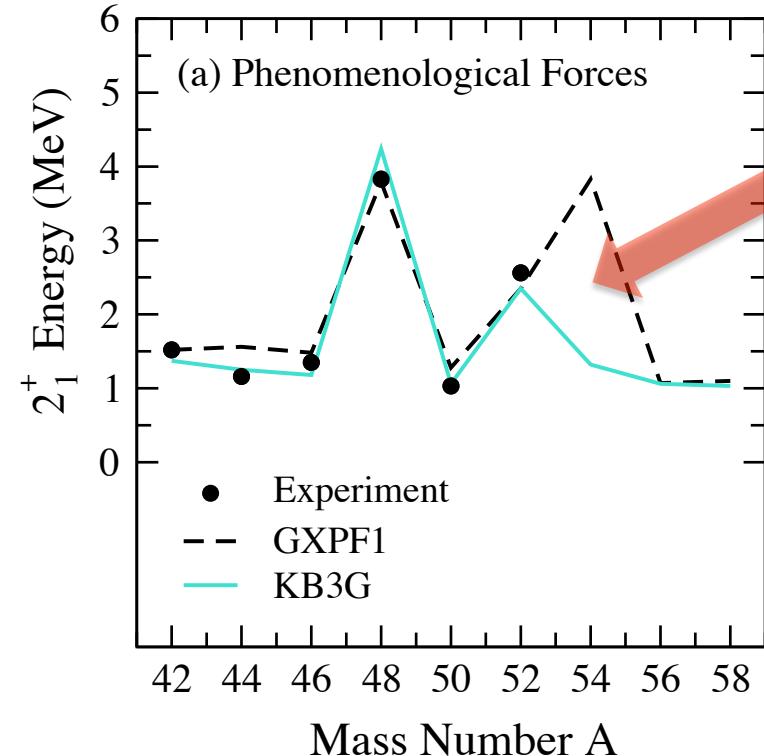
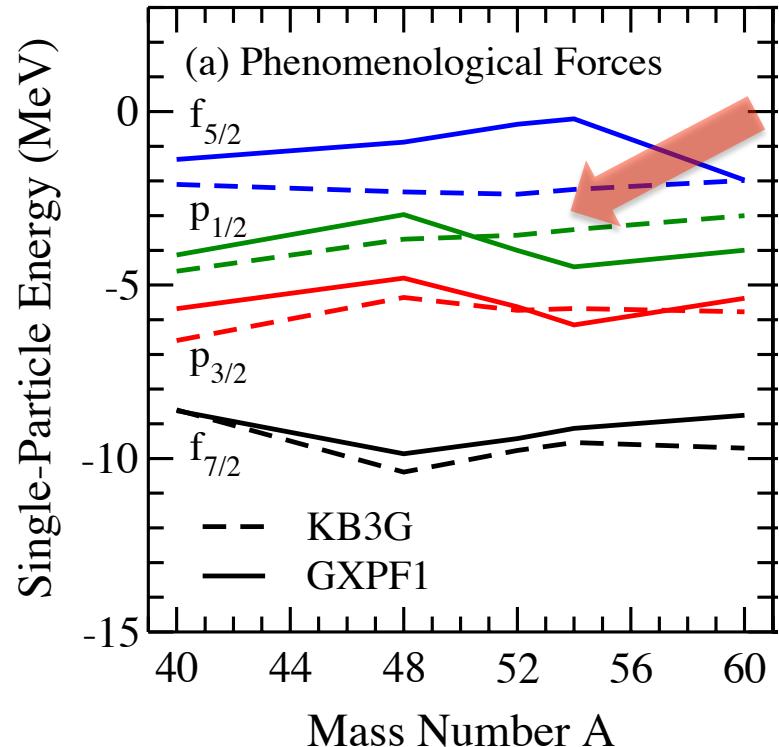
**N=32:** moderate peak

Close to experimental value with new TITAN data

Experimental measurement of  $^{53}\text{Ca}$  mass needed to reduce uncertainty

# Evolution of Magic Numbers: N=34

N=34 magic number in calcium?

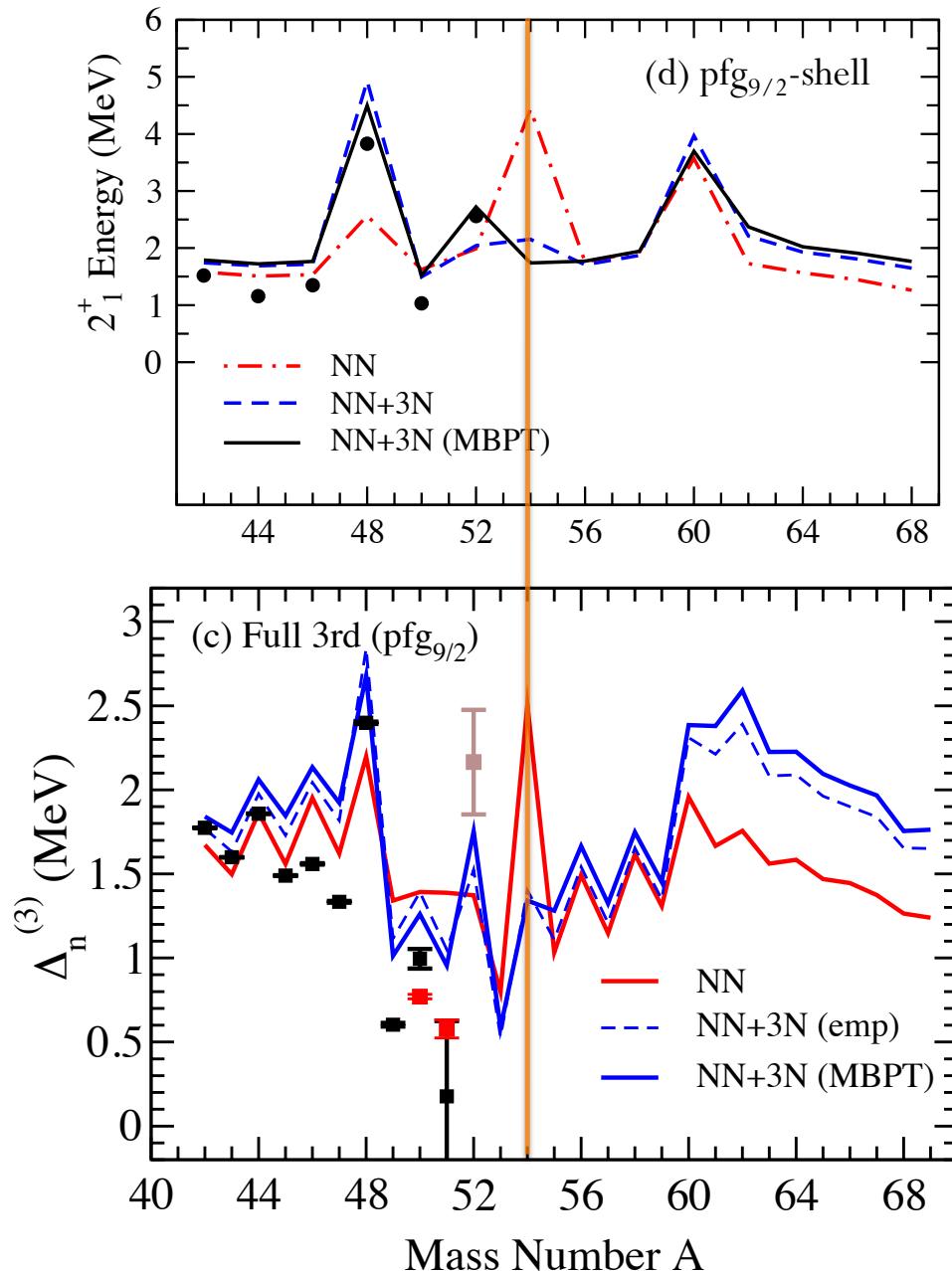


GXPF1: Honma, Otsuka, Brown, Mizusaki (2004)

KB3G: Poves, Sanchez-Solano, Caurier, Nowacki (2001)

Significant phenomenological disagreement for neutron-rich calcium

# Pairing for Shell Evolution N=34



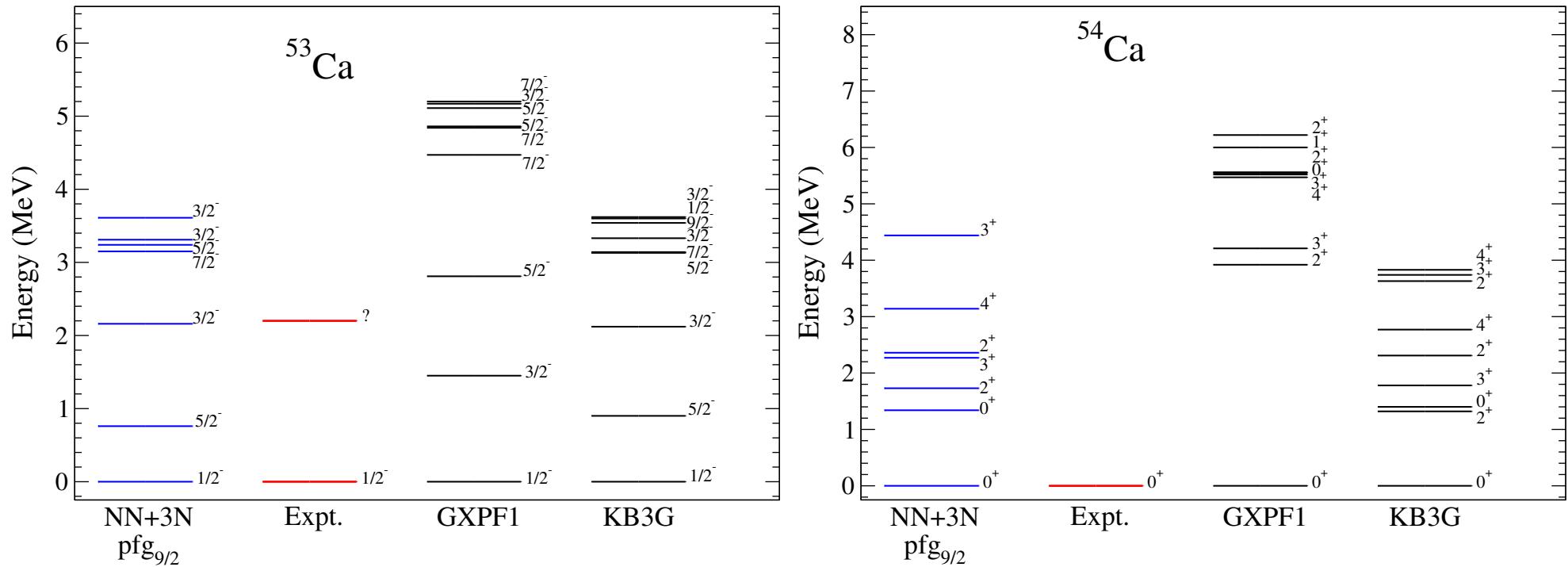
Peak in pairing gaps: complementary signature for shell closure

Compare with  $2^+$  energies for Ca

**N=34:** weak signature – suppression from 3N forces

# Neutron-Rich Ca Spectra Near N=34

Neutron-rich calcium spectra with NN+3N



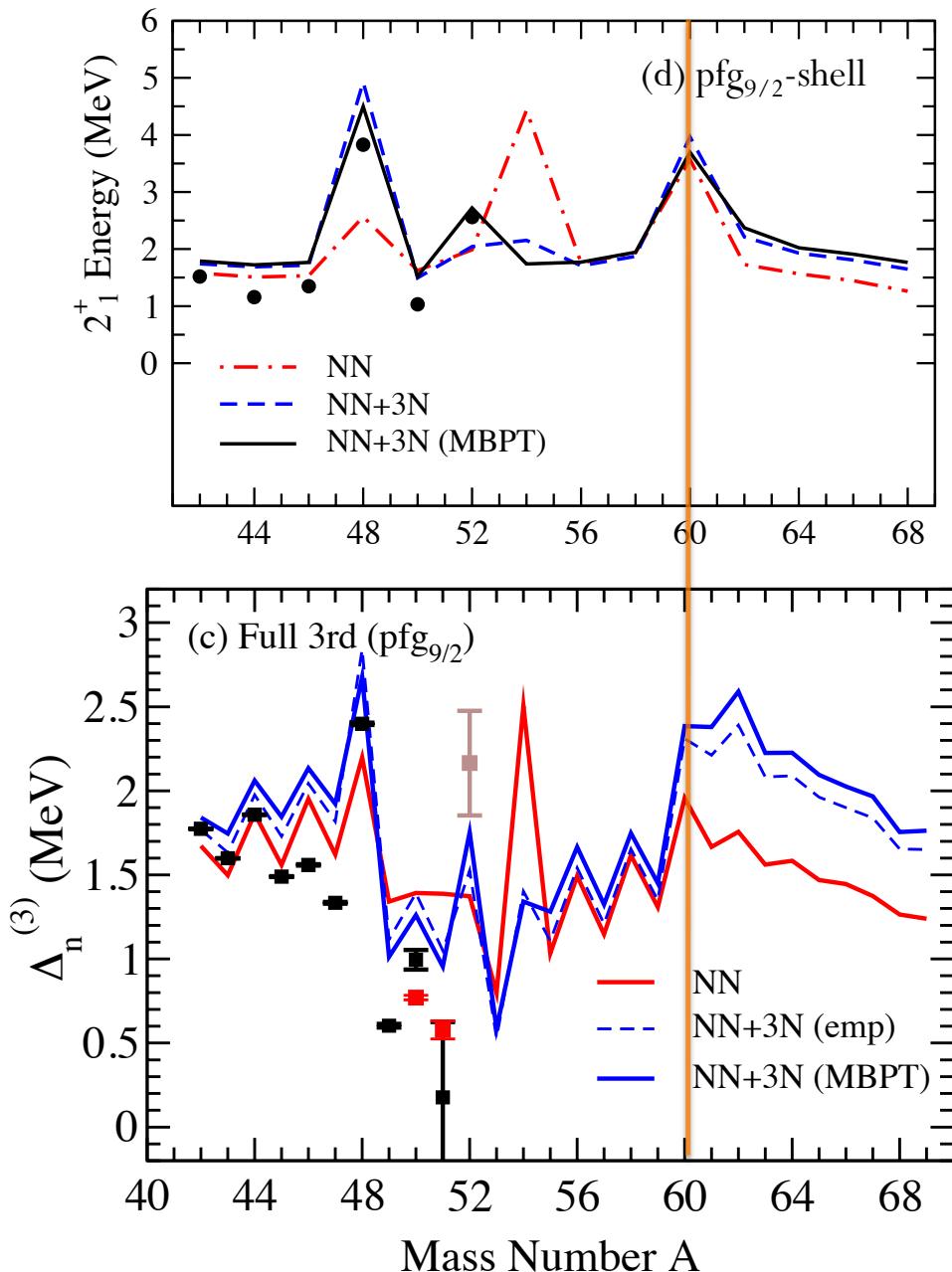
JDH, Menendez, Schwenk, in prep.

Different predictions from phenomenology

NN+3N similar to KB3G – no indication of  $N=34$  magic number

**Consistent with predictions from Coupled-Cluster theory**

# Pairing for Shell Evolution N=40



Peak in pairing gaps: complementary signature for shell closure

Compare with  $2^+$  energies for Ca

**N=40:** robust signature of shell closure

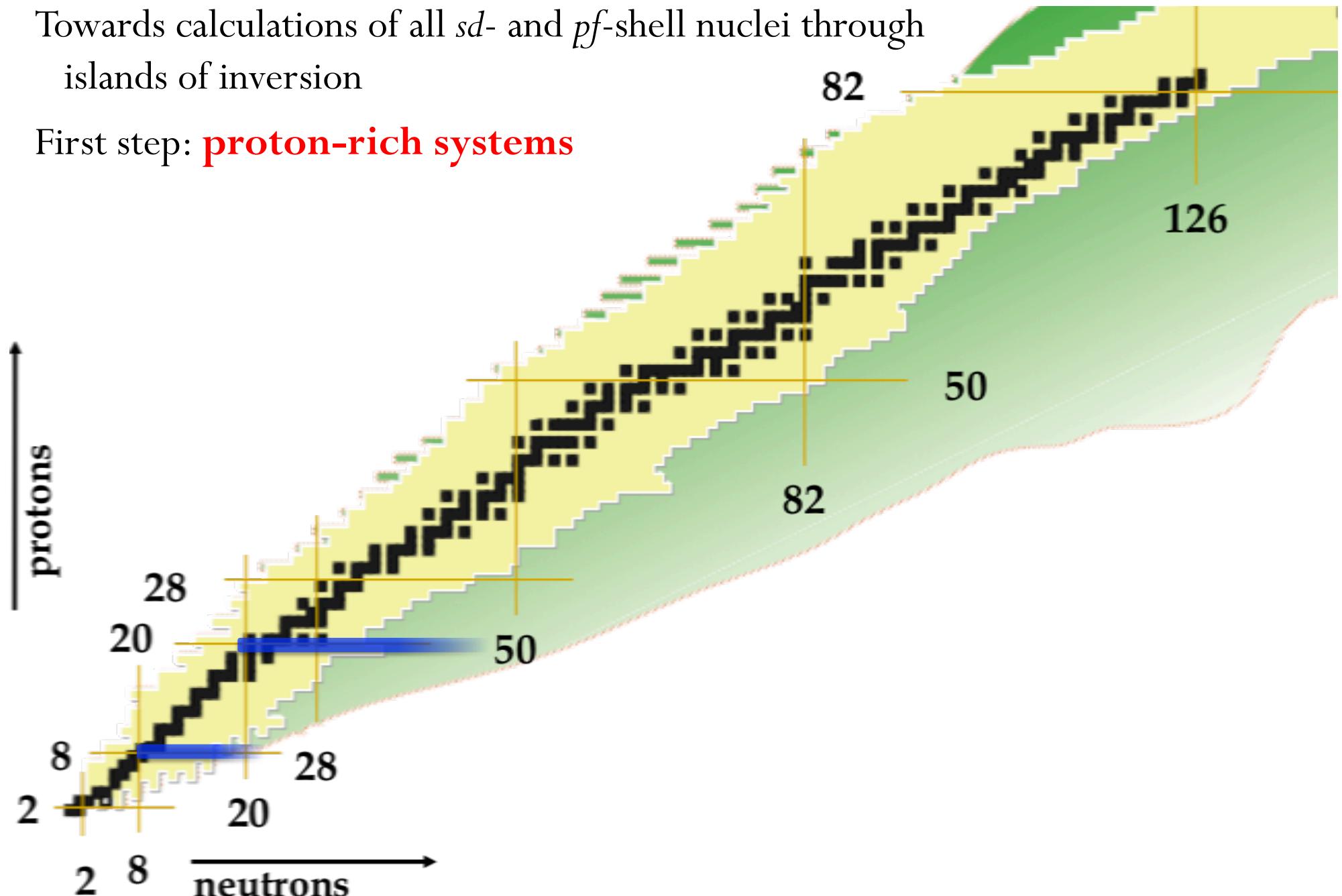
CC: continuum lowers  $2s_{1/2}$ ,  $1d_{5/2}$

Inclusion will affect N=40 prediction

# Proton-Rich Systems

Towards calculations of all *sd*- and *pf*-shell nuclei through islands of inversion

First step: **proton-rich systems**

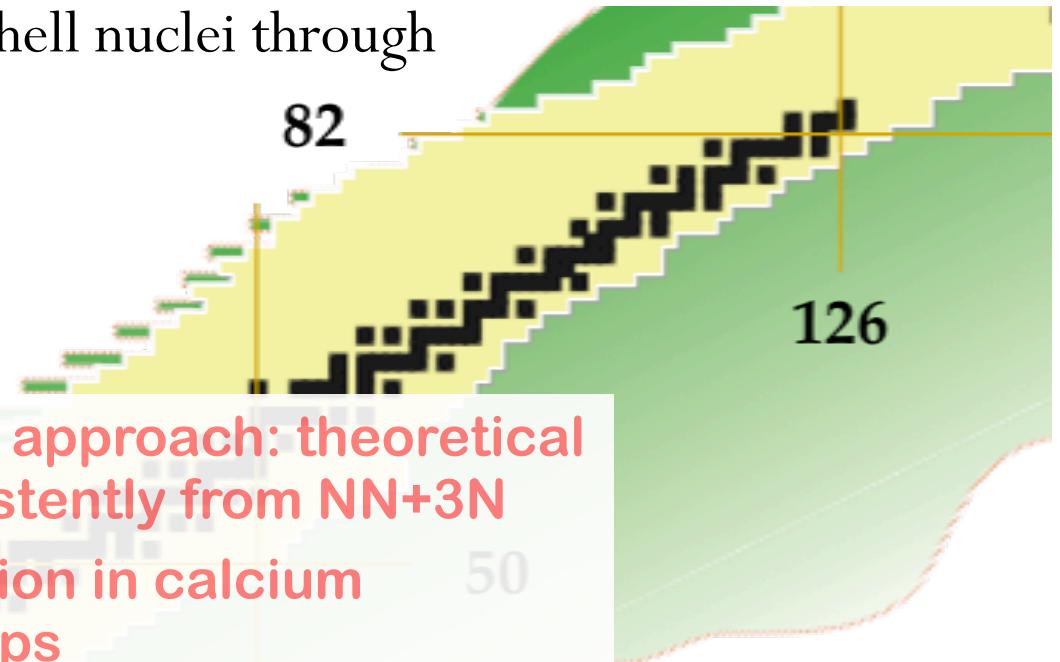
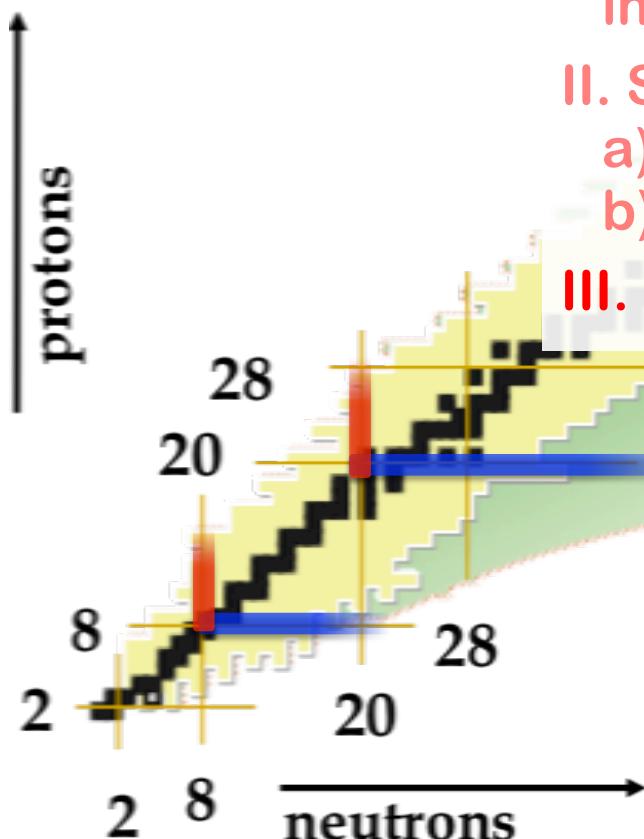


# Proton-Rich Systems

Towards calculations of all *sd*- and *pf*-shell nuclei through islands of inversion

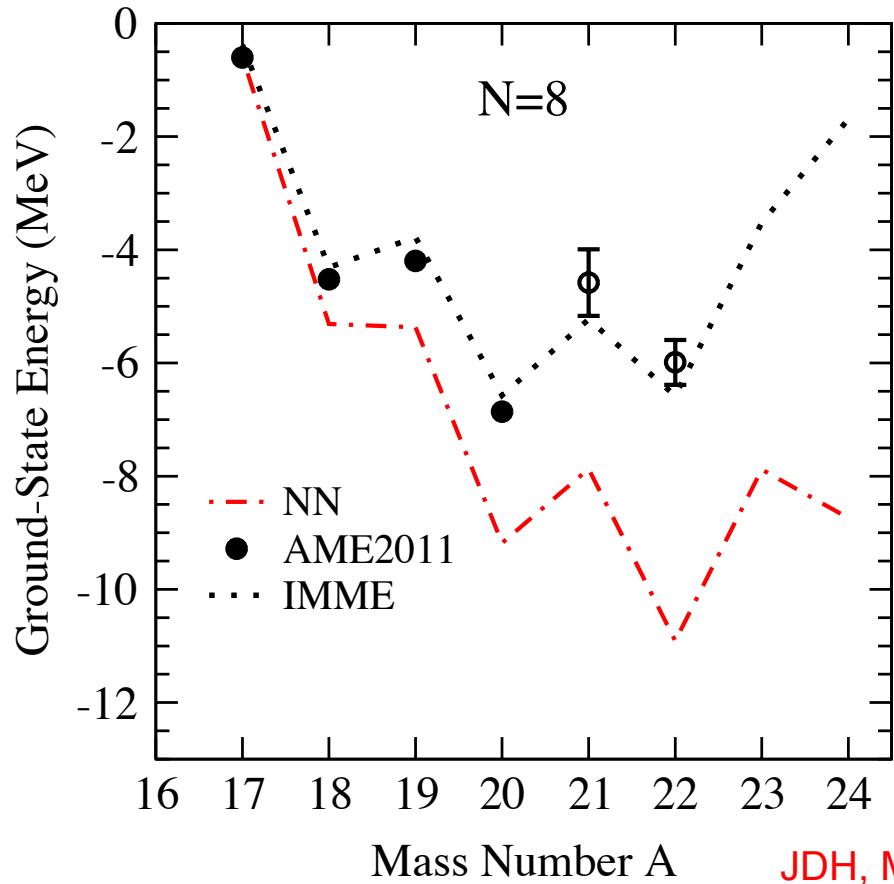
First step: proton-rich systems

**Study N=8, 20 isotone chains**



- I. Microscopic approach: theoretical inputs consistently from NN+3N
- II. Shell evolution in calcium
  - a) Pairing gaps
  - b) Spectra
- III. Proton-rich systems

# Ground-State Energies of N=8 Isotones



Data limited – use phenomenological isobaric multiplet mass equation (IMME)

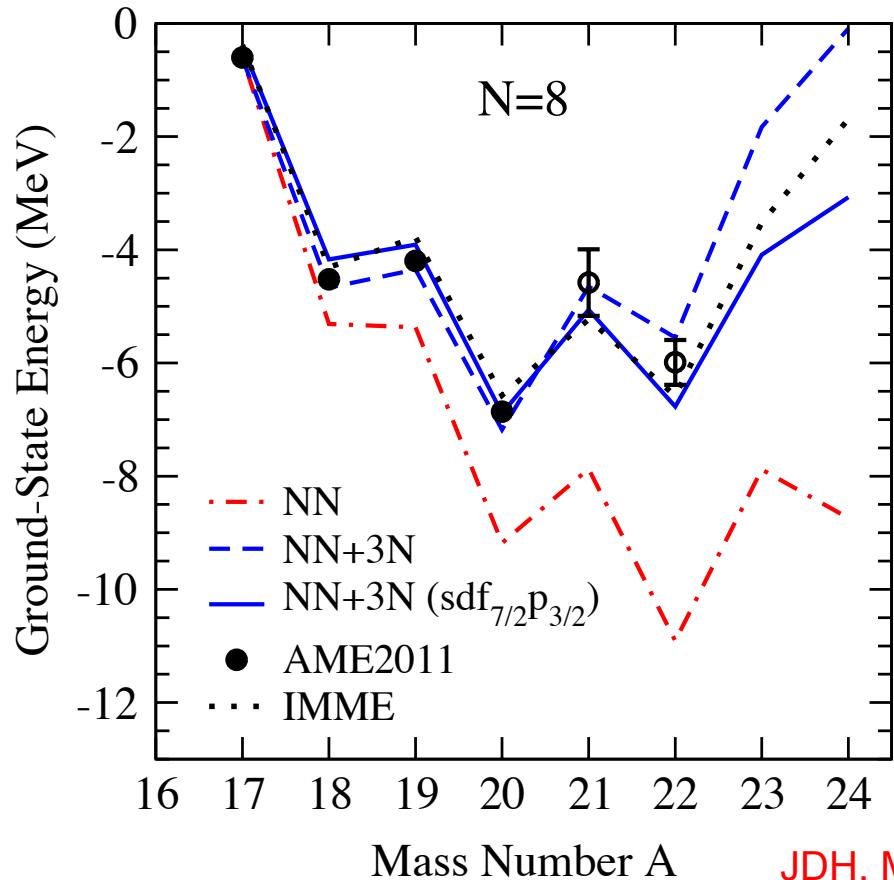
$$E(A, T, T_z) = E(A, T, -T_z) + 2b(A, T)T_z$$

$$b = 0.7068A^{2/3} - 0.9133$$

**NN-only:** overbound

JDH, Menendez, Schwenk, PRL (2013)

# Ground-State Energies of N=8 Isotones



Data limited – use phenomenological isobaric multiplet mass equation (IMME)

$$E(A, T, T_z) = E(A, T, -T_z) + 2b(A, T)T_z$$

$$b = 0.7068A^{2/3} - 0.9133$$

**NN-only**: overbound

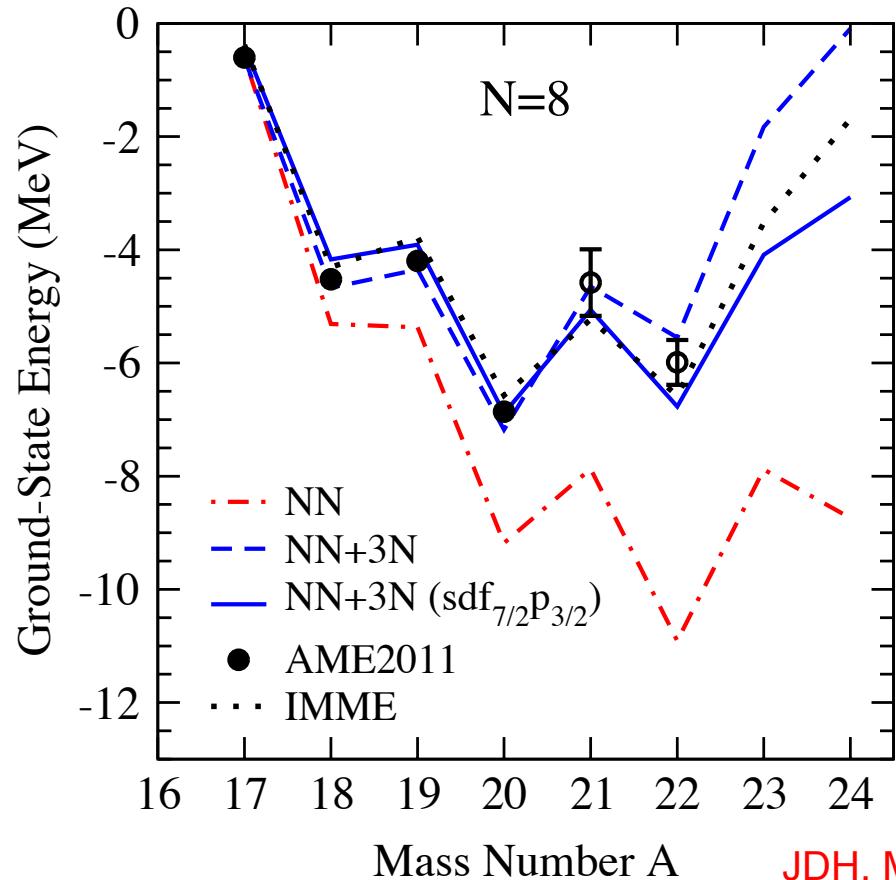
**NN+3N**: improved agreement with experiment/IMME

Extended space important  $\sim A = 21$

JDH, Menendez, Schwenk, PRL (2013)

**Dripline unclear**: <sup>22</sup>Si unbound in AME, NN+3N; bound in IMME

# Ground-State Energies of N=8 Isotones



Data limited – use phenomenological isobaric multiplet mass equation (IMME)

$$E(A, T, T_z) = E(A, T, -T_z) + 2b(A, T)T_z$$

$$b = 0.7068A^{2/3} - 0.9133$$

**NN-only**: overbound

**NN+3N**: improved agreement with experiment/IMME

Extended space important  $\sim A = 21$

JDH, Menendez, Schwenk, PRL (2013)

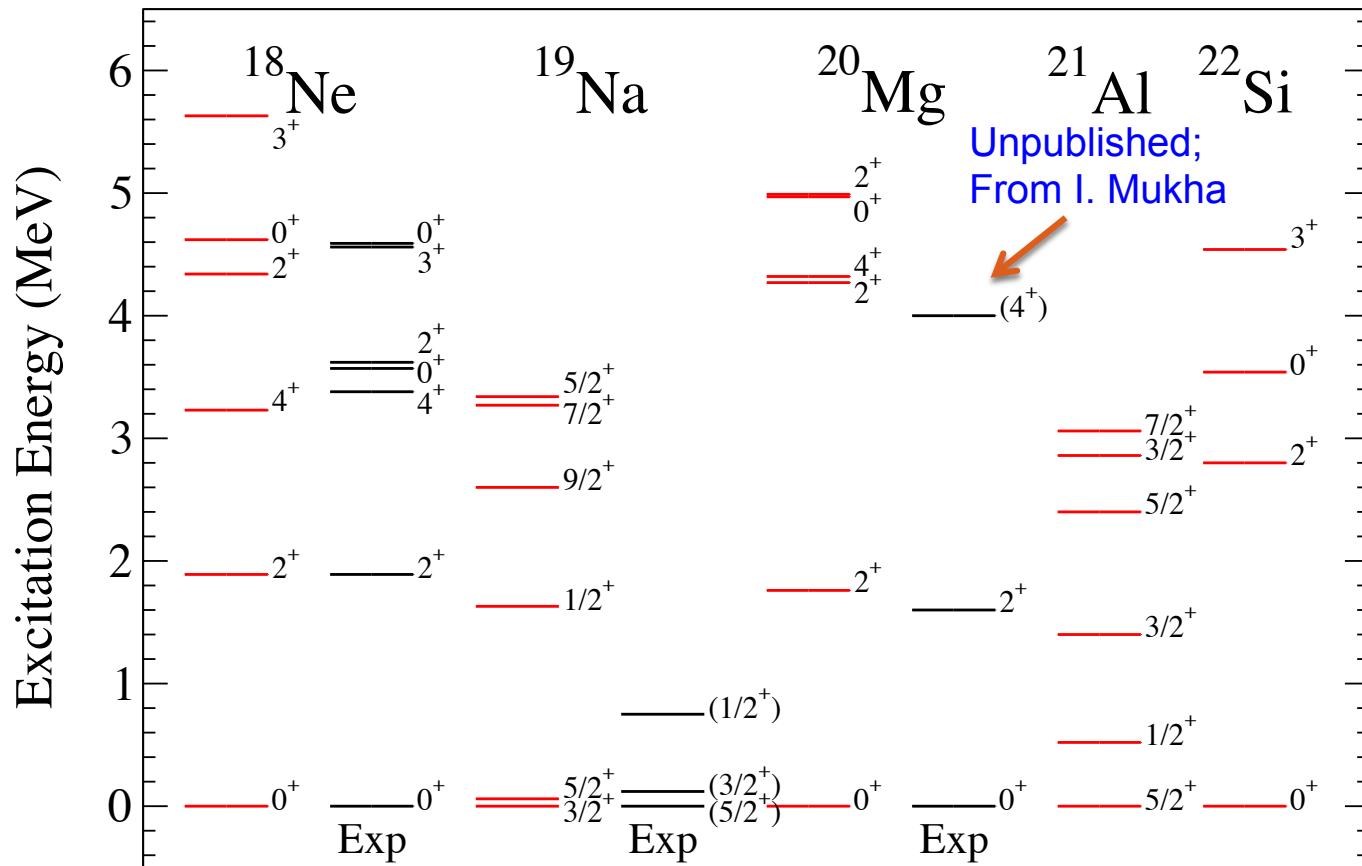
**Dripline unclear**:  $^{22}\text{Si}$  unbound in AME, NN+3N; bound in IMME

$^{22}\text{Si}$  possible two-proton emitter

Measurement needed

$S_{2p}$	IMME	NN+3N ( $sd$ )	NN+3N ( $sdf_{7/2}p_{3/2}$ )
	0.01 MeV	-1.63 MeV	-0.12 MeV

# Spectra of N=8 Isotones



JDH, Menendez, Schwenk, PRL (2012)

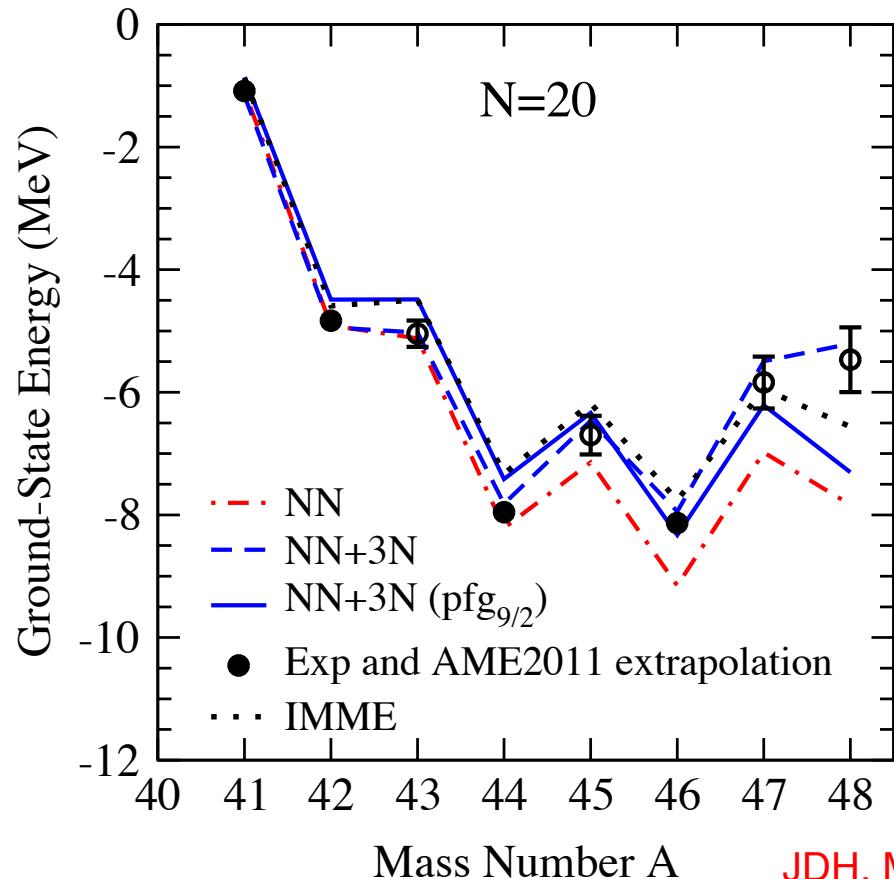
**NN+3N:** reasonable agreement with experiment

**New measurement:** excited state in  $^{20}\text{Mg}$  close to predicted  $4^+-2^+$  doublet

Predictions for proton-rich  $^{21}\text{Al}$ ,  $^{22}\text{Si}$  spectra

Closed sub-shell signature in  $^{22}\text{Si}$

# Ground-State Energies of N=20 Isotones



**NN-only:** overbound beyond  $^{45}\text{Mn}$   
**NN+3N:** close to experiment/IMME

JDH, Menendez, Schwenk, PRL (2013)

**Dripline:** Predicted to be  $^{46}\text{Fe}$  in all calculations

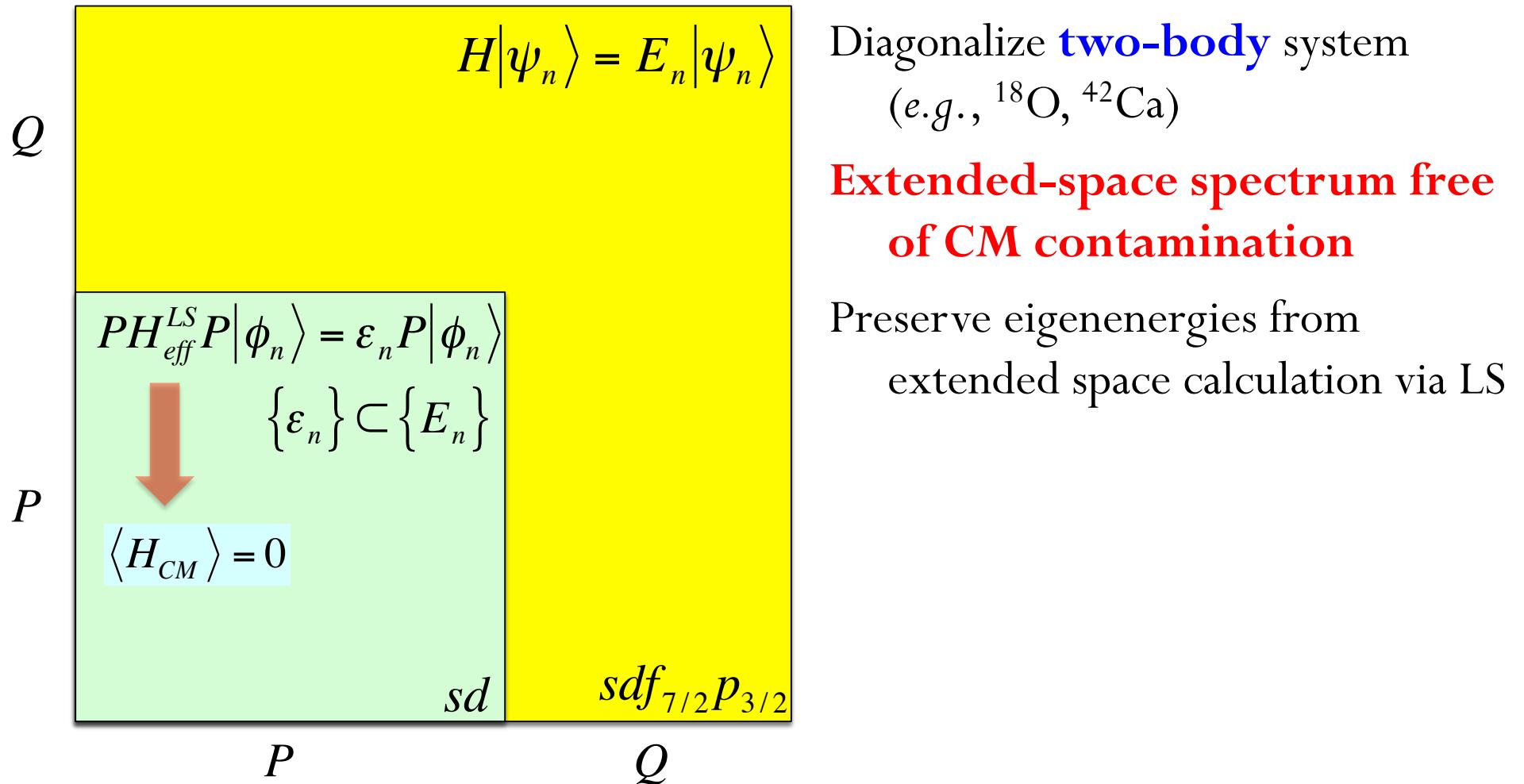
$S_{2p}$	Expt.	NN+3N ( $pf$ )	NN+3N ( $pfg_{9/2}$ )
	-1.28(6) MeV	-2.73 MeV	-1.02 MeV

Prediction for  $^{48}\text{Ni}$  within 300keV of experiment

Dossat et al (2005); Pomorski et al (2012)

# Evaluating Center-of-Mass Contamination

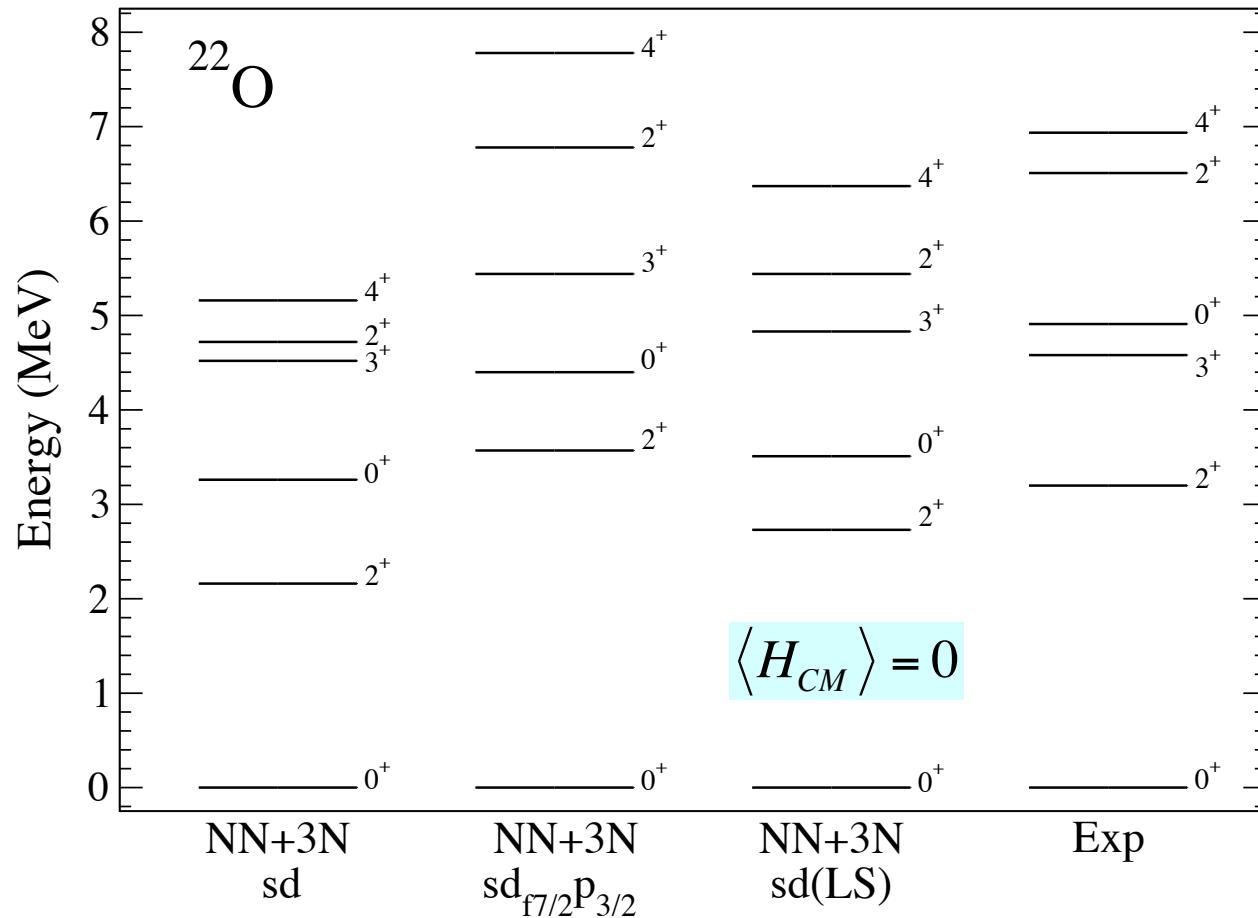
Nonperturbative Lee-Suzuki (LS) transformation from extended space



Use  $H_{eff}^{LS}$  as new two-body Hamiltonian in *sd*-shell valence-space calculations

# Evaluating Center-of-Mass Contamination

Apply new  $H_{eff}^{LS}$  to calculate spectra in neutron-rich oxygen



Improvements from standard  $sd$ -shell – not due to center of mass

**Work in progress:** involving  $N > 2$  neutrons in extended space

# Conclusion

- Nuclear structure theory of medium-mass nuclei with 3N forces, extended spaces
- Robust repulsive 3N mechanism for  $T=1$  neutron/proton-rich nuclei
- **Oxygen isotopes**
  - Cures NN-only failings: dripline, shell evolution, spectra
- **Calcium isotopes**
  - Shell evolution towards the dripline from  $2^+$  energies and  $\Delta_n^{(3)}$
  - Weak  $N=34$  closure predicted
- **Proton-rich N=8, 20 isotones**: similar improvements in g.s. energies/spectra

# Acknowledgments

## Collaborators



TECHNISCHE  
UNIVERSITÄT  
DARMSTADT

J. Menendez, J. Simonis,  
A. Schwenk

T. Otsuka

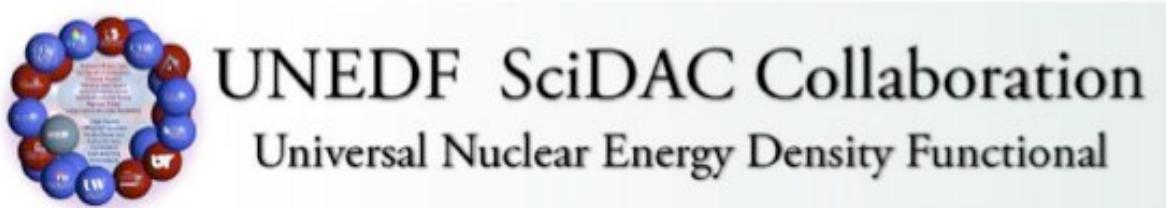


T. Suzuki (Nihon U.)



S. Bogner

Supported by BMBF under 06DA7047I (NuSTAR.DA)



Computing support

