# Infrared properties of the harmonic oscillator basis

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R. J. Furnstahl, G. Hagen, TP, Phys. Rev. C 86, 031301(R) (2012); arXiv:1207.6100 Sushant N. More, A. Ekström, R. J. Furnstahl, G. Hagen, TP, arXiv:1302.3815





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## Convergence in finite oscillator spaces

Calculations are performed in finite oscillator spaces. How can one reliably extrapolate to infinity?

Convergence in momentum space (UV) and in position space (IR) needed



• cutoff of interaction  $\Lambda < \Lambda_{UV}$ 

### "What is the infrared cutoff in the HO basis?" [U. van Kolck at INT workshop in spring 2009]

Very precise answer [More, Ekström, Furnstahl, Hagen, TP, 2013] based on length scale

$$L_2 = \sqrt{2(N+3/2+2)}b$$

- 1. At low energies, the HO basis looks like a "box" of radius  $L_2$ .
- 2.  $\pi/L_2$  is the infrared cutoff.
- Knowledge can be used for theoretically founded extrapolations in HO basis, computations of phase shifts in HO basis ...

## Spectrum of the operator p<sup>2</sup>



- At low momentum, number of states increases linearly with increasing momentum
- Spectrum looks like that of the momentum operator in a box

Eigenfunctions of p<sup>2</sup> with lowest eigenvalues in oscillator basis



Eigenfunctions looks like those from a box of size  $L_2$ .

## Squared infrared cutoff is the lowest eigenvalue of $p^2$

The lowest eigenvalue  $\kappa_{min}$  can be computed analytically for N>>1. **Result:**  $\pi/L_2$ 

"N>>1" does not imply impractically large model spaces

N	$\kappa_{ m min}$	$\pi/L_2$	$\pi/L_0$	$L_i \equiv \sqrt{2(N+3/2+i)b}$
0	1.2247	1.1874	1.8138	
2	0.9586	0.9472	1.1874	1% deviation at N>2
4	0.8163	0.8112	0.9472	
6	0.7236	0.7207	0.8112	
8	0.6568	0.6551	0.7207	
10	0.6058	0.6046	0.6551	
12	0.5651	0.5642	0.6046	
14	0.5316	0.5310	0.5642	0.1% deviation at N>14
16	0.5035	0.5031	0.5310	
18	0.4795	0.4791	0.5031	
20	0.4585	0.4582	0.4791	$\pi/L_2$ is very precise value of the IR cutoff

#### IR corrections to bound-state energies

Simple view: A node in the wave function

$$u_E(r) \xrightarrow{r \gg R} A_E(e^{-k_E r} + \alpha_E e^{+k_E r})$$

at  $r=L_2$  requires  $\alpha_e = exp(-2k_eL_2)$ . This yields a (kinetic) energy correction

$$E_L = E_\infty + a_0 e^{-2k_\infty L}$$

**Model-independent approach** based on [D. Djajaputra & B. R. Cooper, Eur. J. Phys. 21, 261 (2000)].

$$\Delta E_L \approx -u_{\infty}(L) \left( \frac{du_E(L)}{dE} \bigg|_{E_{\infty}} \right)$$

Final results: ANC<sup>2</sup> Binding momentum  

$$\Delta E_L = \frac{\hbar^2 k_{\infty} \gamma_{\infty}^2}{\mu} e^{-2k_{\infty}L} + \mathcal{O}(e^{-4k_{\infty}L}) \qquad \text{Only observables enter}$$

$$\langle r^2 \rangle_L \approx \langle r^2 \rangle_\infty [1 - (c_0 \beta^3 + c_1 \beta) e^{-\beta}] \quad (\text{with } \beta \equiv 2k_\infty L)$$

Energy extrapolation explains findings by Coon et al, Phys. Rev. C 86, 054002 (2012)

## Triton binding energy from SRG interactions: only observables enter into the IR extrapolation



## Phase shifts

- 1. Compute states in channel I with positive energies  ${\rm E}_{\rm i}$  and momentum  ${\rm p}_{\rm i}$  in HO basis at fixed N
- 2. In a box, the i<sup>th</sup> state determines the box size  $L_i = L(p_i)$  at that energy via  $j_l(p_i L_i/\hbar) = 0$   $i_l(k_i L(\hbar k_i))$
- 3. Compute phase shift from usual formula:  $\tan \delta_l(k_i) = \frac{j_l(k_i L(\hbar k_i))}{\eta_l(k_i L(\hbar k_i))}$
- 4. Repeat for several  $\hbar\Omega$



#### Phase shifts



Alternative approaches based on [Busch et al 1998] employ a harmonic potential and use  $\hbar\Omega \rightarrow 0$  for finite-range interactions:

T. Luu, M. J. Savage, A. Schwenk, and J. P. Vary, Phys. Rev. C 82, 034003 (2010). I. Stetcu, J. Rotureau, B. R. Barrett, and U. van Kolck, J. Phys. G 37, 064033 (2010).

## How well can one distinguish $L_2$ in practice?





Deuteron (N<sup>3</sup>LO E&M)

## How well can one distinguish $L_2$ in practice?



Deuteron (N<sup>3</sup>LO E&M)

#### How well can one distinguish the exponential law in practice?

Gaussian well

Deuteron (N<sup>3</sup>LO E&M)



## **Corrections for shallow bound states**



#### **Corrections due to finite Hilbert spaces**

- UV practically converged (because  $\lambda < \Lambda_{UV}$ )
- IR convergence is slower due to exponential decay of wave function
- Dirichlet boundary condition at x=L in position space





# Recipe

- 1. Perform calculations at sufficiently large values of  $\hbar\Omega$  (these have small or no UV corrections)
- 2. Plot results (energies, radii) vs.  $L_2$  (UV converged results are expected to fall onto a single line)
- 3. Perform fit to extrapolation formulas and read off asymptotic value

# Summary

- Much improved understanding of IR properties of HO basis
- At low momenta, HO basis behaves as a box of size L<sub>2</sub>
- $\pi/L_2$  is the IR cutoff
- Computation of phase shifts directly from the positive energy states in HO basis
- Energy extrapolation law expressed solely in terms of observables
- Corrections for shallow bound states worked out

Outlook: IR properties in *any localized* basis

- Diagonalize operator  $p^2$  in a given model space  $\rightarrow$  L for this model space
- Be in the UV-converged regime
- Plot energies and radii as a function of L, and extrapolate