

Efficient calculation of 3N forces at N^3LO for ab initio studies

Kai Hebler

Vancouver, February 17, 2015

**TRIUMF workshop on
“Progress in ab initio Techniques in Nuclear Physics”**



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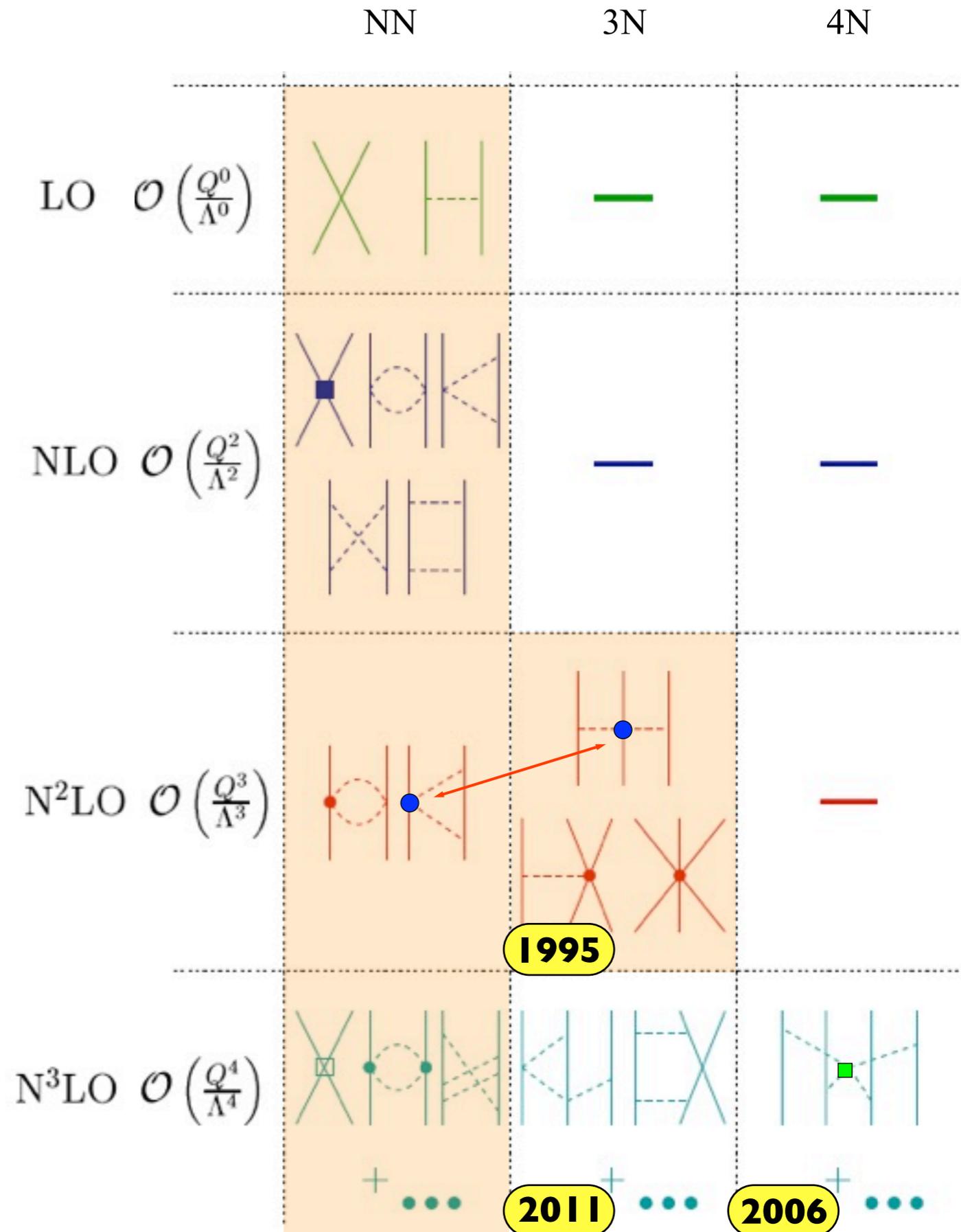
Hermann Krebs, Evgeny Epelbaum, Roman Skibinski, Jacek Golak

[arXiv:1502.02977](https://arxiv.org/abs/1502.02977)

Chiral effective field theory for nuclear forces

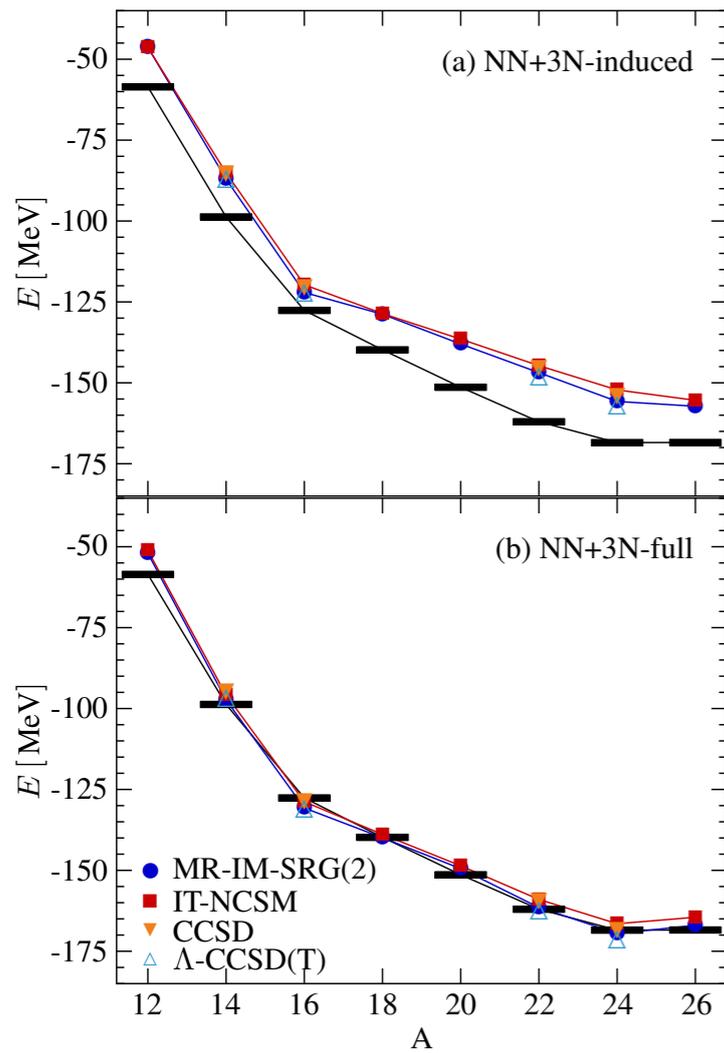
- choose relevant degrees of freedom: here nucleons and pions
- operators constrained by symmetries of QCD
- short-range physics captured in few short-range couplings
- separation of scales: $Q \ll \Lambda_b$, breakdown scale $\Lambda_b \sim 500$ MeV
- power-counting: expand in powers Q/Λ_b
- systematic: work to desired accuracy, obtain error estimates

treatment of NN and 3N forces
not consistent in present
ab initio calculations



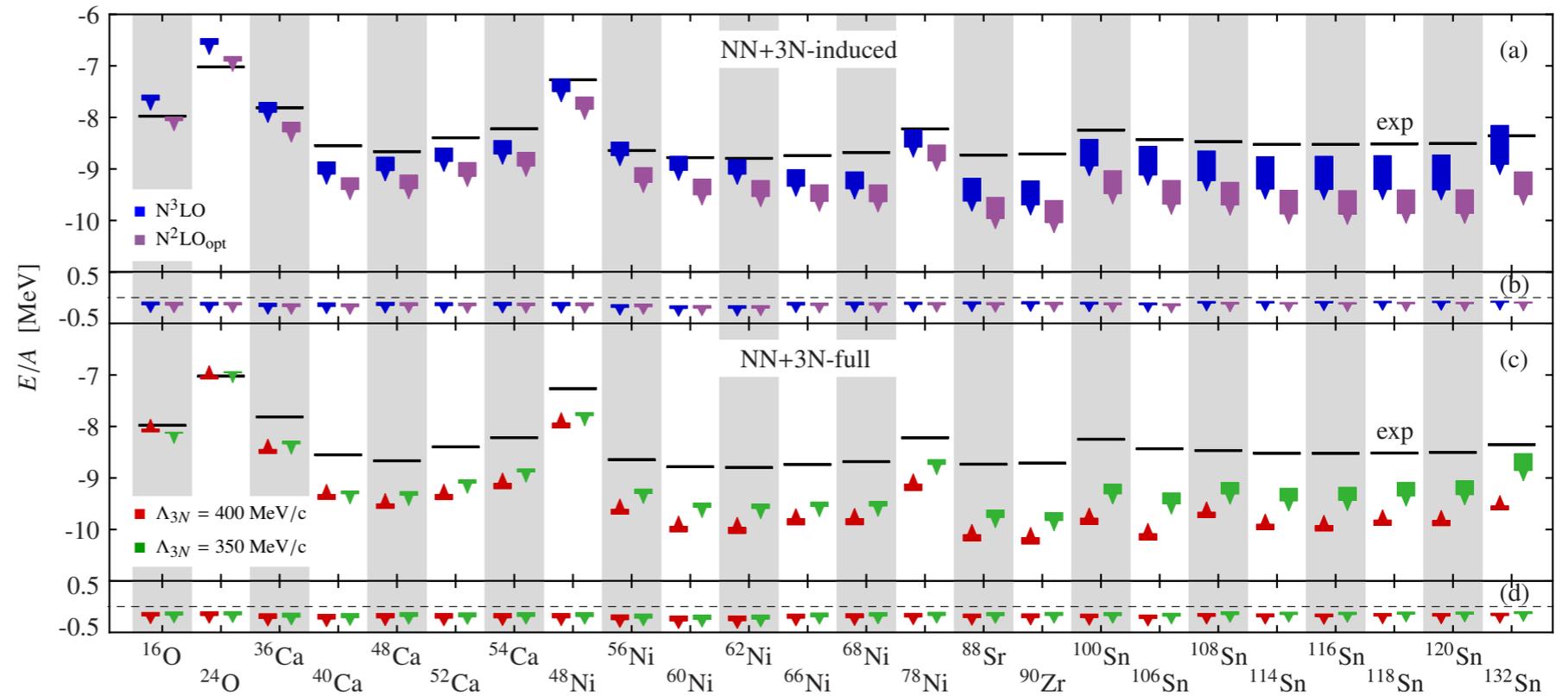
Open issues in nuclear interactions

oxygen chain



Hergert et al.,
 PRL 110, 242501 (2013)

heavy nuclei

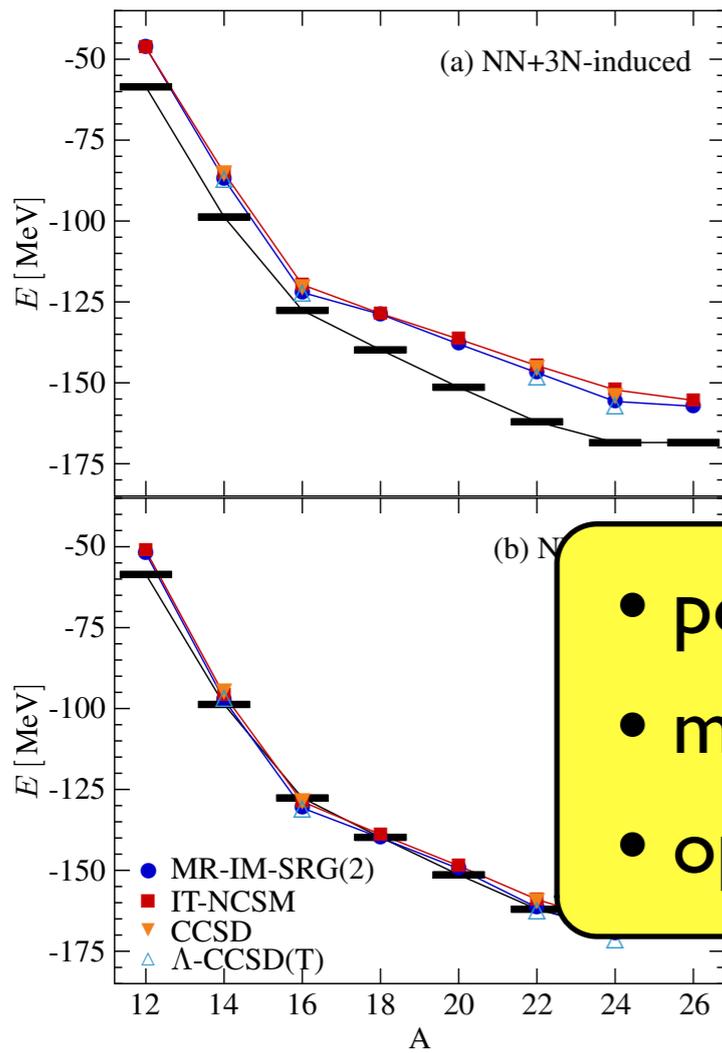


Binder et al., Phys. Lett B 736, 119 (2014)

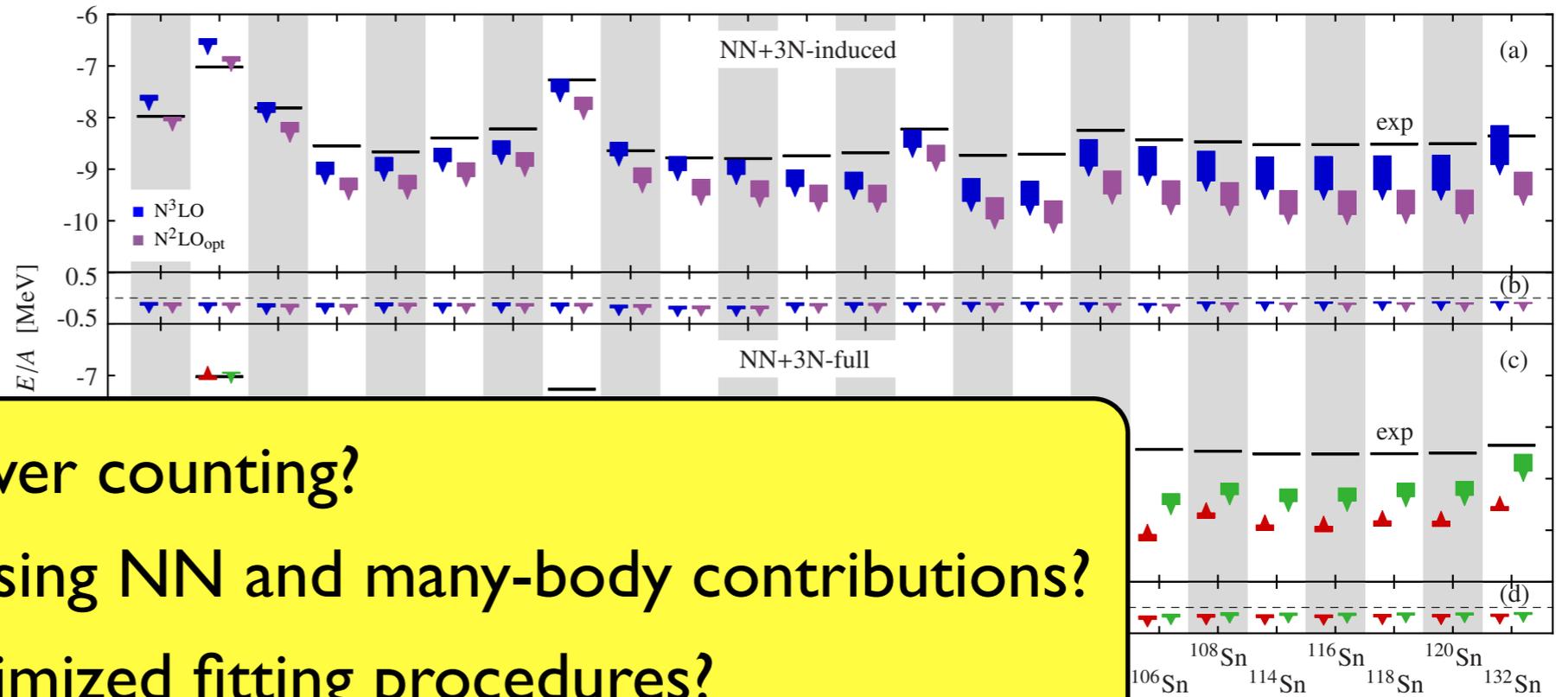
- remarkable agreement between different many-body frameworks
- significant overbinding in heavy nuclei

Open issues in nuclear interactions

oxygen chain



heavy nuclei



- power counting?
- missing NN and many-body contributions?
- optimized fitting procedures?

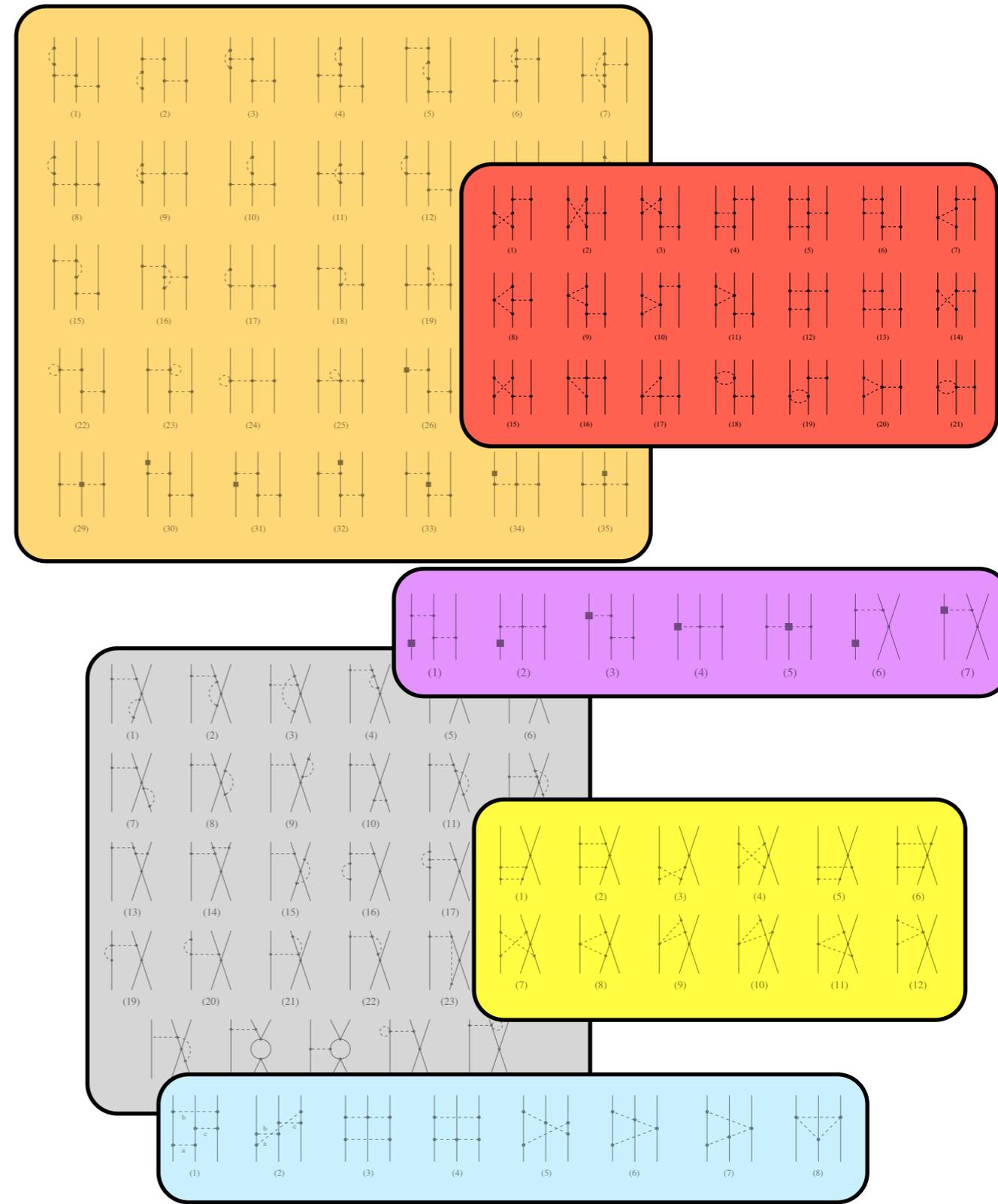
Hergert et al.,
PRL 110, 242501 (2013)

- remarkable agreement between different many-body frameworks
- significant overbinding in heavy nuclei

Chiral 3N forces at subleading order (N^3LO)

	2N forces	3N forces	4N forces
$\mathcal{O}\left(\frac{Q^0}{\Lambda^0}\right)$			
$\mathcal{O}\left(\frac{Q^2}{\Lambda^2}\right)$			
$\mathcal{O}\left(\frac{Q^3}{\Lambda^3}\right)$			
$\mathcal{O}\left(\frac{Q^4}{\Lambda^4}\right)$			

1995
2011
2006



Bernard et al., PRC 77, 064004 (2008)
 Bernard et al., PRC 84, 054001 (2011)
 Krebs et al., PRC 85, 054006 (2012)
 Krebs et al., PRC 87, 054007 (2013)

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2011

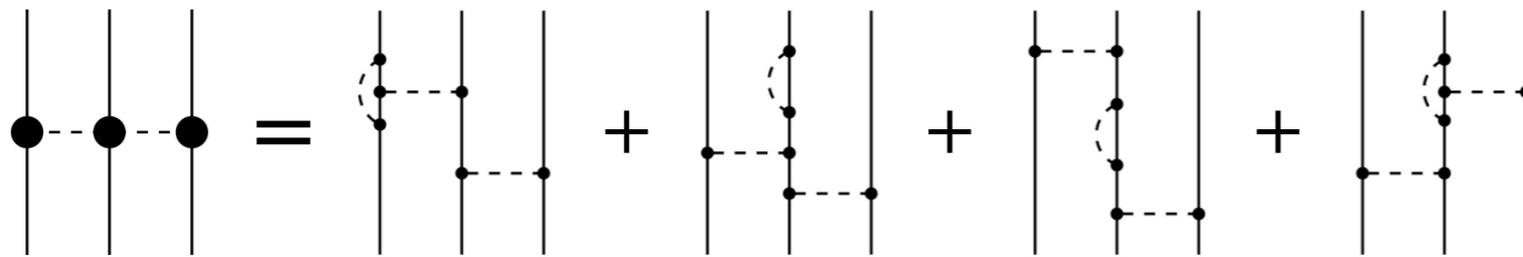
ALL TERMS PREDICTED

key for

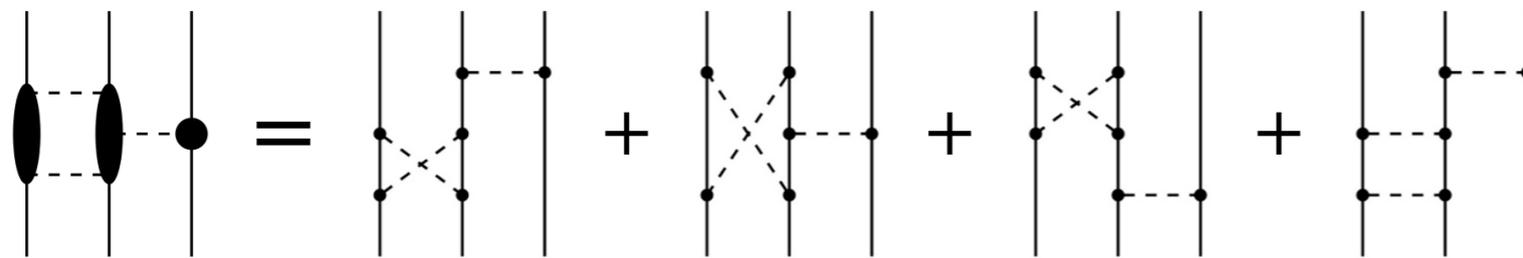
- consistency
- tests
- improved precision
- uncertainty estimates of the theory

Bernard et al., PRC 77, 064004 (2008)
 Bernard et al., PRC 84, 054001 (2011)
 Krebs et al., PRC 85, 054006 (2012)
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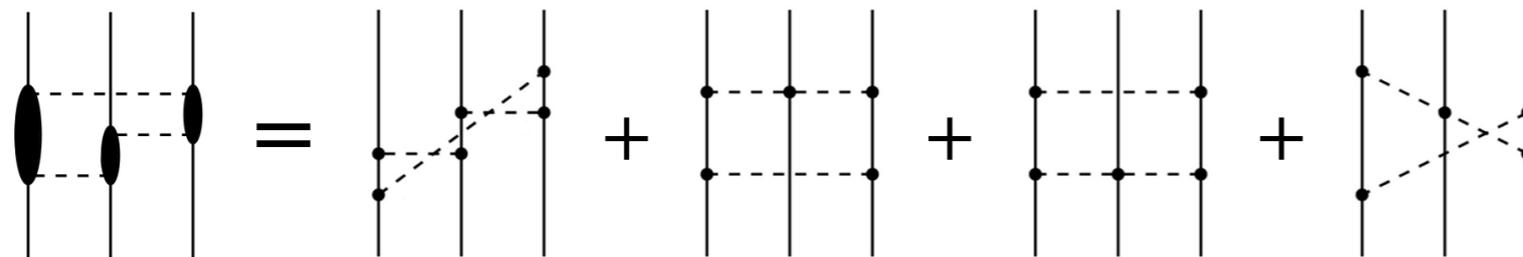
Three-nucleon force contributions at N³LO



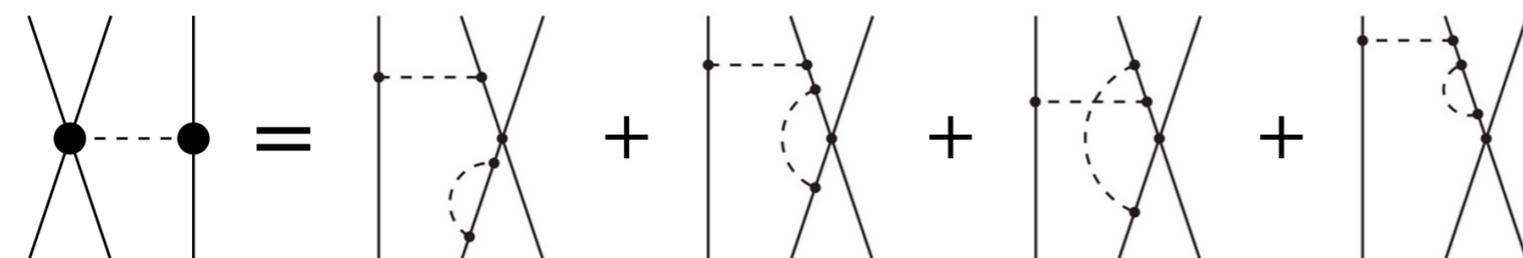
2pi exchange



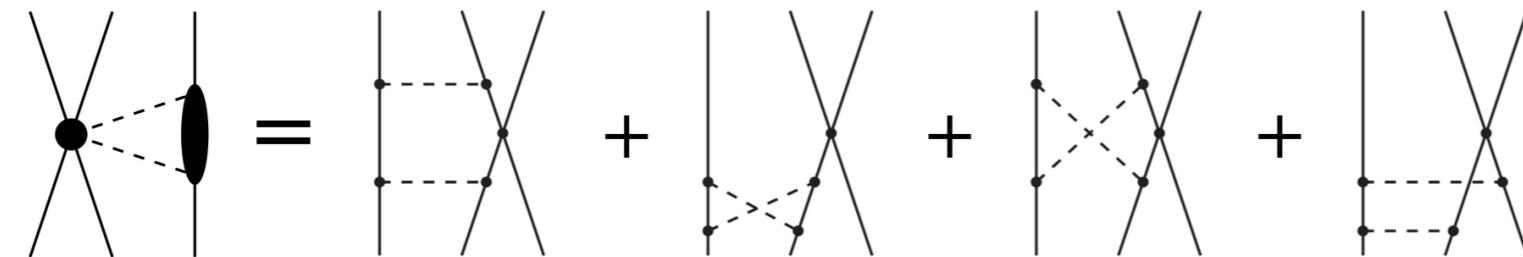
2pi-1pi exchange



pion rings



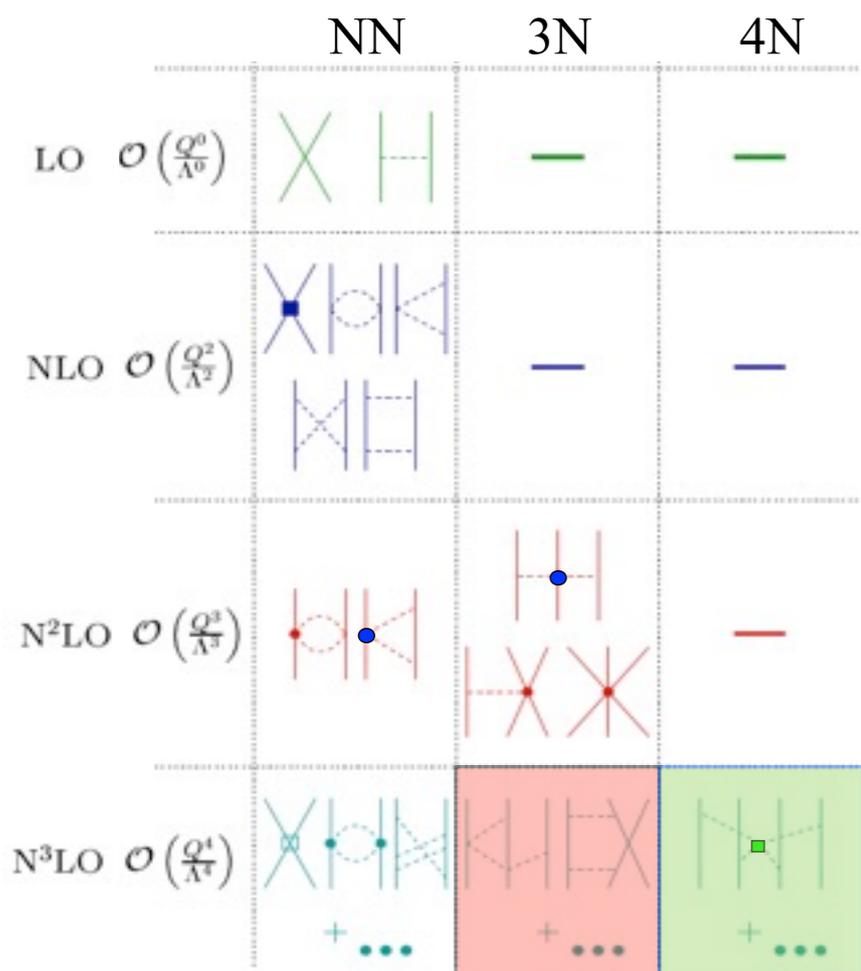
1pi-contact



2pi-contact

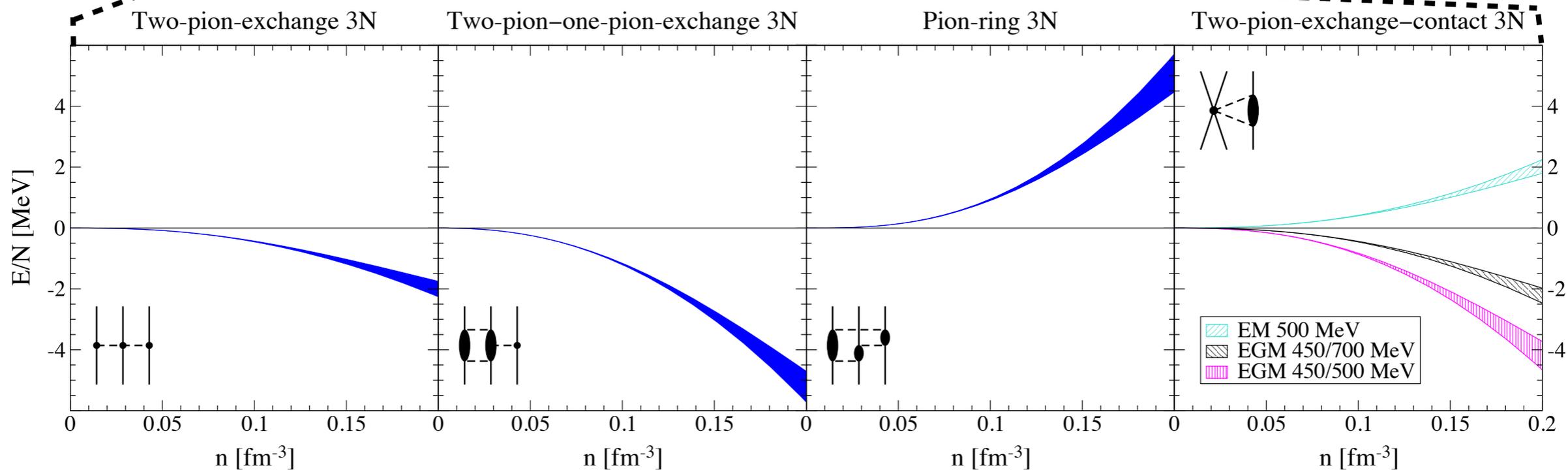
rel. corrections

Contributions of many-body forces at N³LO in neutron matter

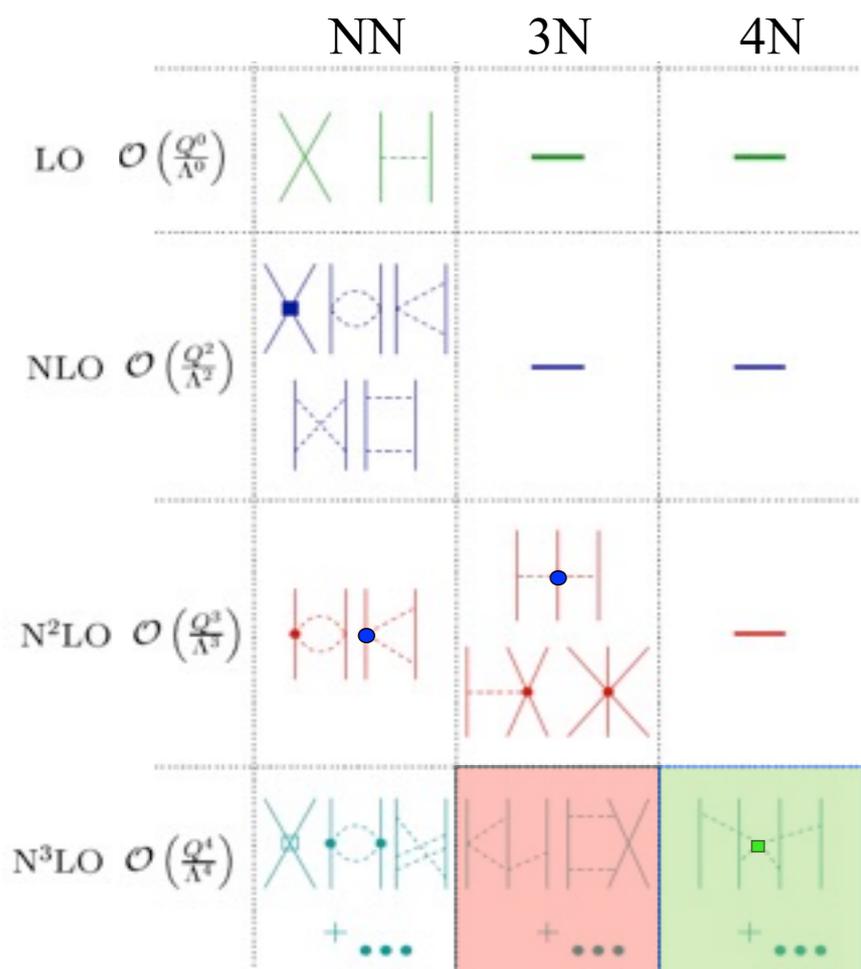


- first calculations of N³LO 3NF and 4NF contributions to EOS of neutron matter
- found **large contributions** in Hartree Fock appr., comparable to size of N²LO contributions

Tews, Krüger, KH, Schwenk
PRL 110, 032504 (2013)

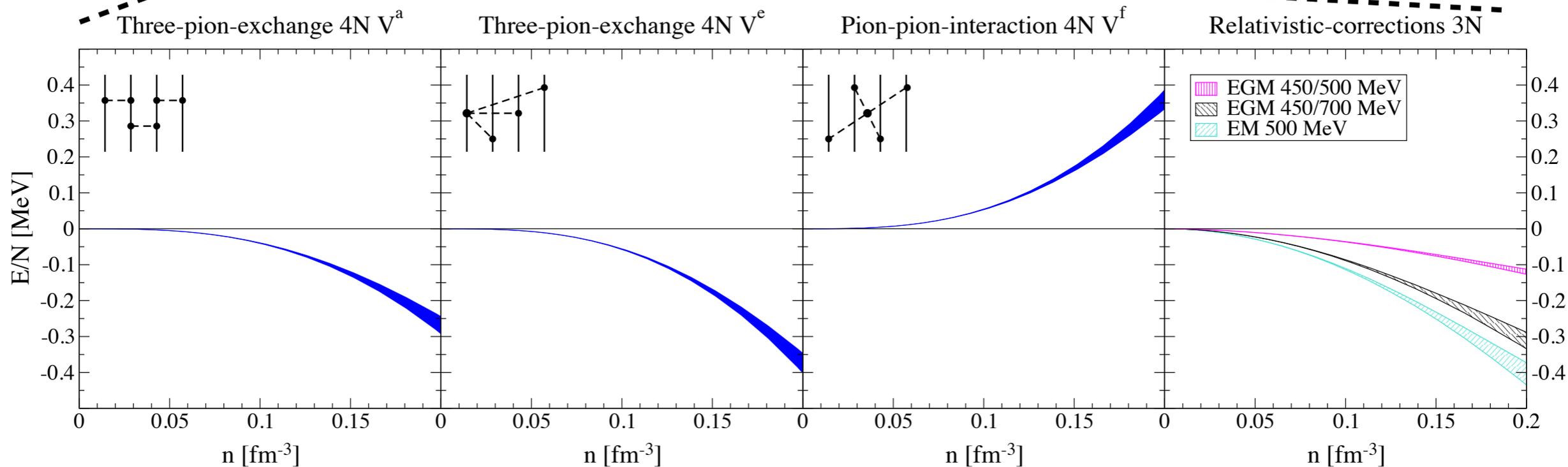


Contributions of many-body forces at N³LO in neutron matter

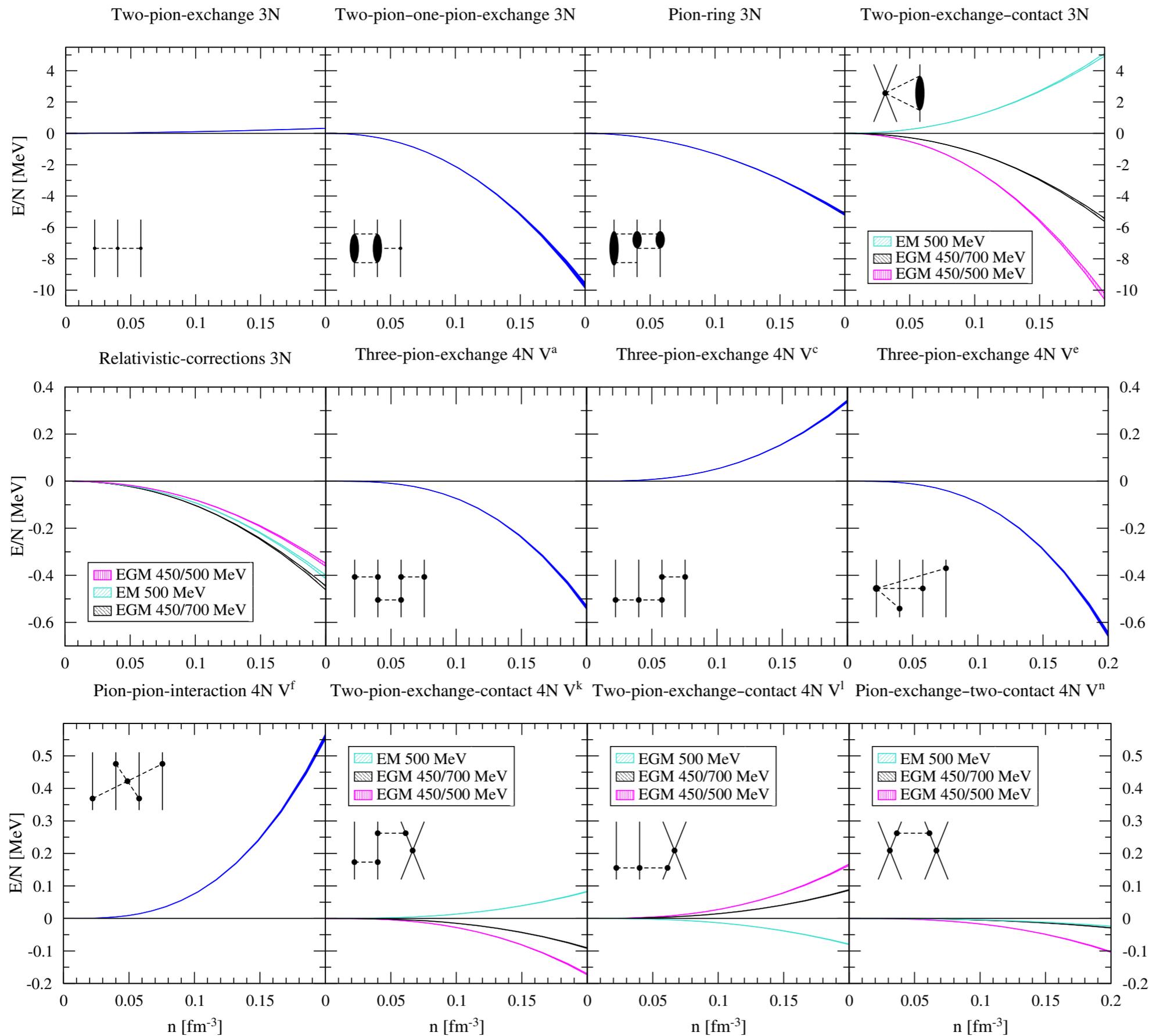


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- found **large contributions** in Hartree Fock appr., comparable to size of N²LO contributions
- 4NF contributions **small**

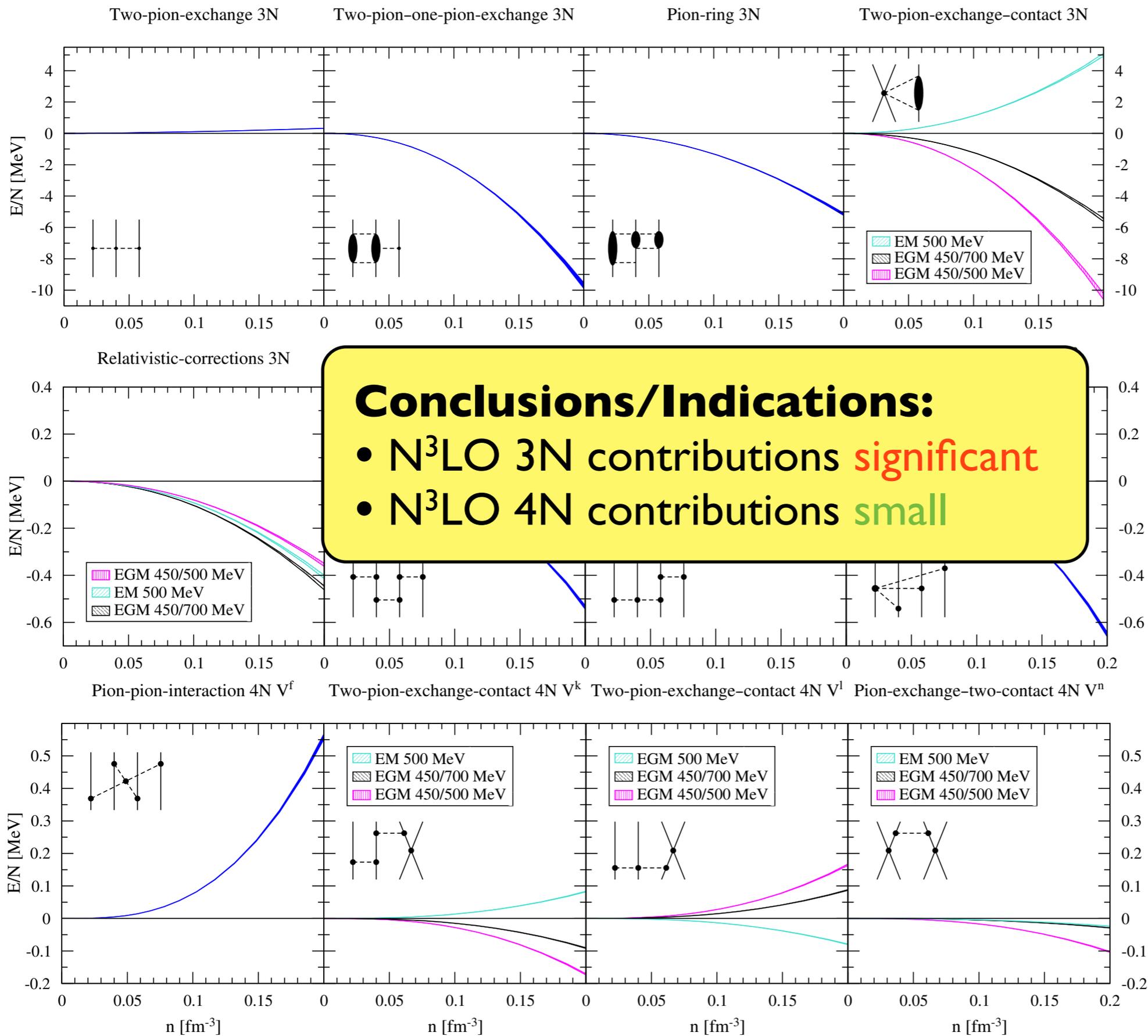
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N³LO contributions in nuclear matter (Hartree Fock)

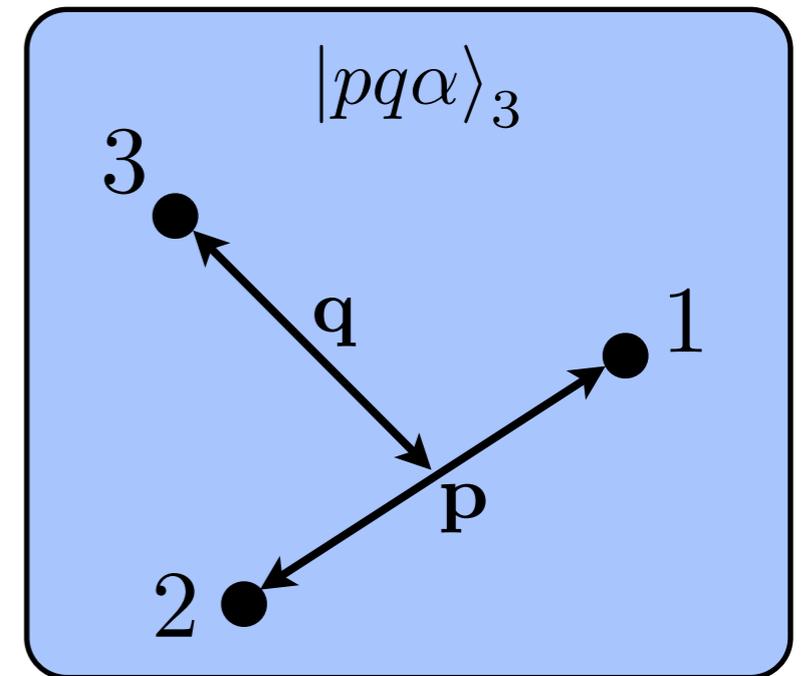
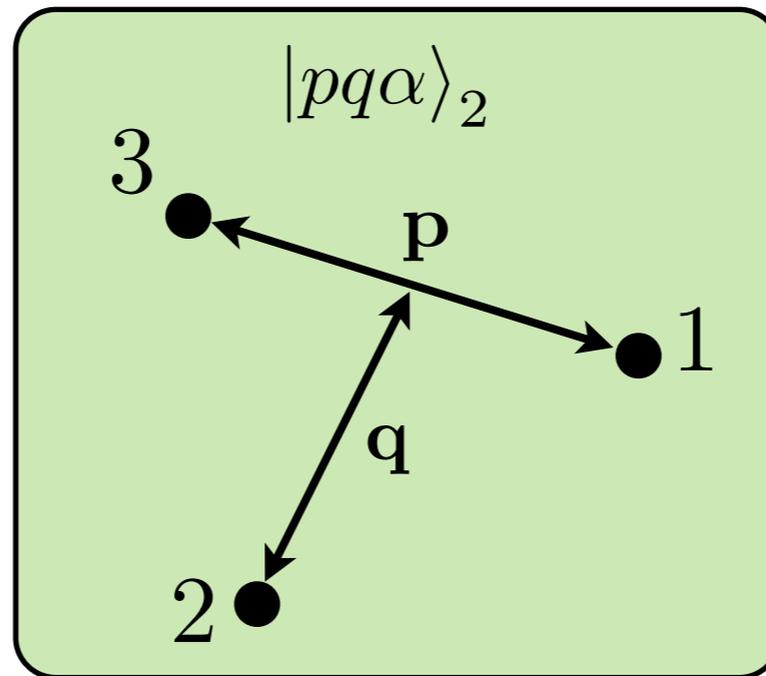
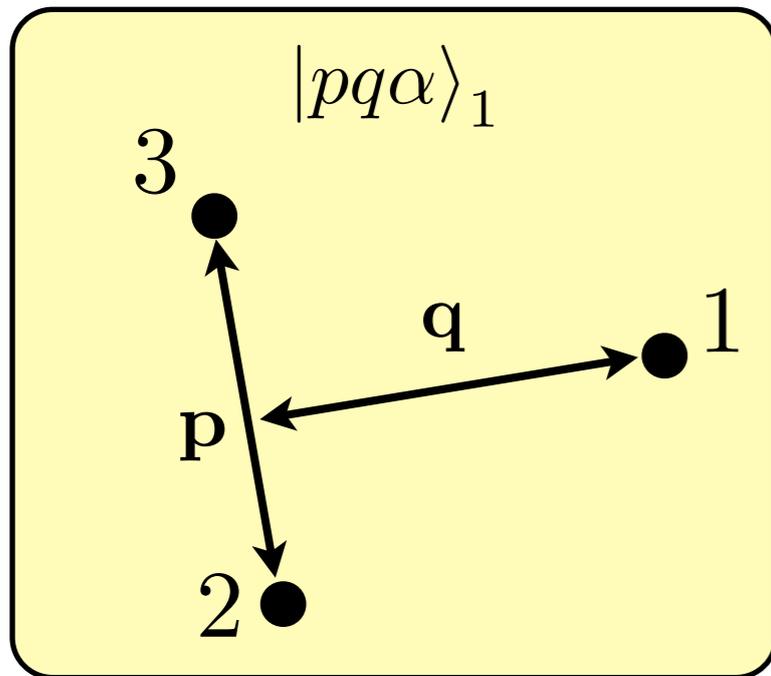


N³LO contributions in nuclear matter (Hartree Fock)



Representation of 3N interactions in momentum space

$$|pq\alpha\rangle_i \equiv |p_i q_i; [(LS)J(l s_i)j] \mathcal{J} \mathcal{J}_z (T t_i) \mathcal{T} \mathcal{T}_z\rangle$$



Due to the large number of matrix elements, the traditional way of computing matrix elements requires extreme amounts of computer resources.

$$\begin{array}{l} N_p \simeq N_q \simeq 15 \\ N_\alpha \simeq 30 - 180 \end{array} \longrightarrow \dim[\langle pq\alpha | V_{123} | p' q' \alpha' \rangle] \simeq 10^7 - 10^{10}$$

Number of matrix elements was so far not sufficient for studies of $A \geq 4$ systems.

Calculation of 3N forces in momentum partial-wave representation

$$\langle pq\alpha|V_{123}|p'q'\alpha'\rangle \sim \sum_{m_i} \int d\hat{\mathbf{p}} d\hat{\mathbf{q}} d\hat{\mathbf{p}}' d\hat{\mathbf{q}}' Y_l^m(\hat{\mathbf{p}}) Y_{\bar{l}}^{\bar{m}}(\hat{\mathbf{q}}) \langle \mathbf{p}\mathbf{q}ST|V_{123}|\mathbf{p}'\mathbf{q}'S'T'\rangle Y_{l'}^{m'}(\hat{\mathbf{p}}') Y_{\bar{l}'}^{\bar{m}'}(\hat{\mathbf{q}}')$$

traditional method:

- reduce dimension of angular integrals from 8 to 5 by using symmetry
- discretize angular integrals and perform all sums numerically

Calculation of 3N forces in momentum partial-wave representation

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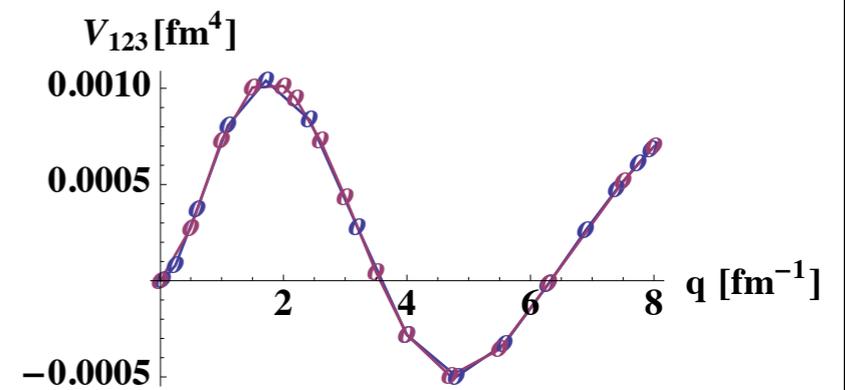
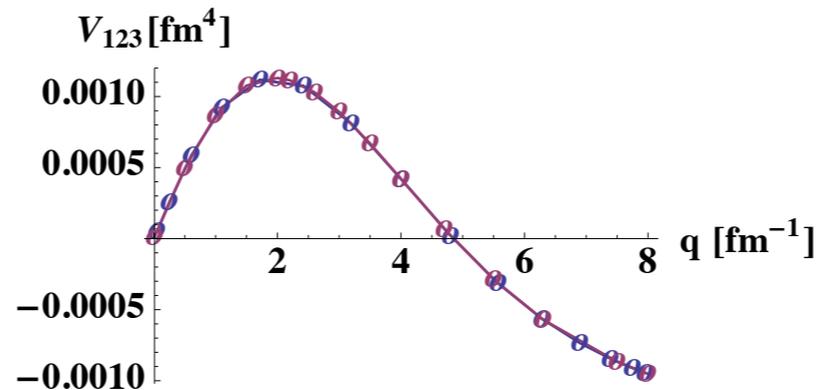
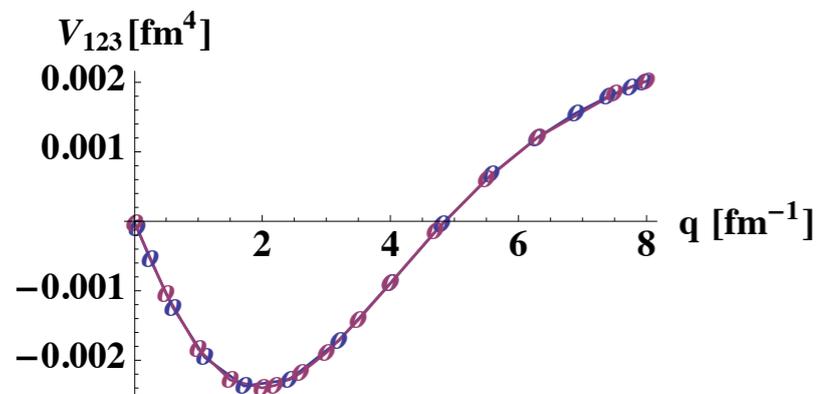
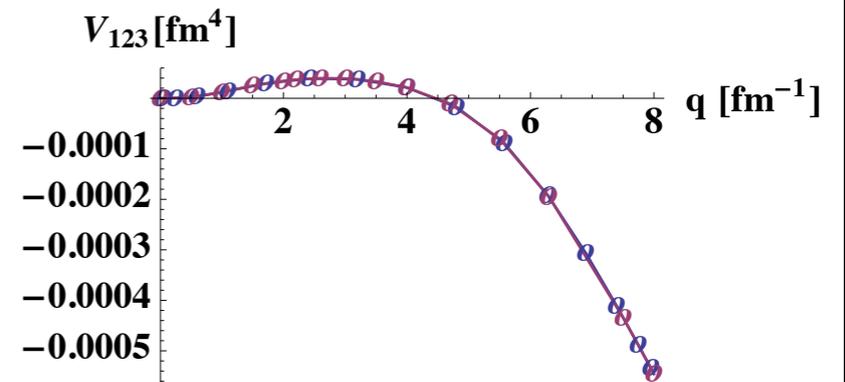
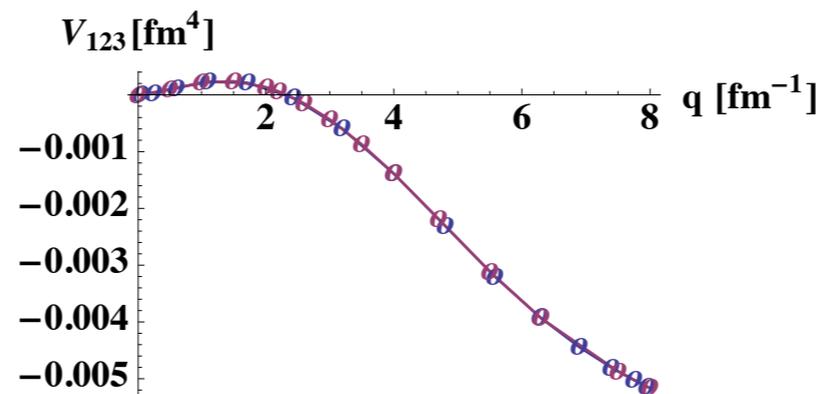
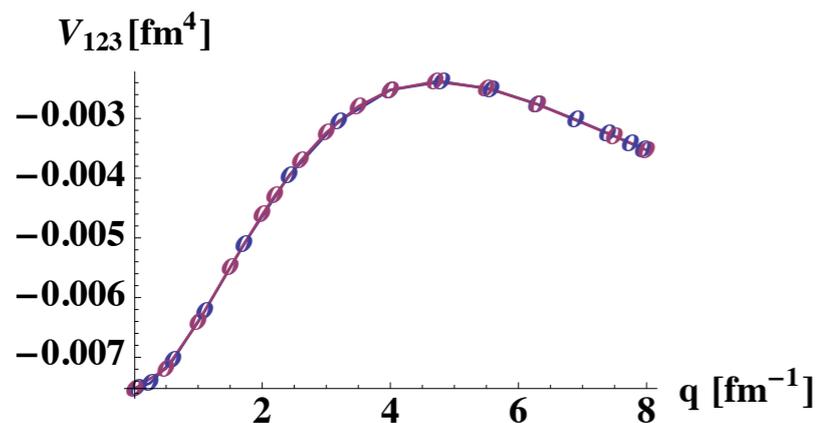
new method:

- use that all interaction contributions (except rel. corr.) are local:

$$\begin{aligned} \langle \mathbf{p}\mathbf{q}|V_{123}|\mathbf{p}'\mathbf{q}'\rangle &= V_{123}(\mathbf{p} - \mathbf{p}', \mathbf{q} - \mathbf{q}') \\ &= V_{123}(p - p', q - q', \cos \theta) \end{aligned}$$

- allows to perform all except 3 integrals analytically
- only a few small discrete internal sums need to be performed for each external momentum and angular momentum

Tests of the new framework



- **perfect agreement** with results based on traditional approach
- **speedup** factors of > 1000
- **very general**, can also be applied to
 - ▶ pion-full EFT
 - ▶ N⁴LO terms
 - ▶ currents?
- **efficient**: allows to study systematically alternative regulators

Current status of calculations

- all 3N topologies are calculated and stored separately, allows to easily adjust values of LECs and the cutoff value and form of non-local regulators

- calculated matrix elements of Faddeev components

$$\langle pq\alpha | V_{123}^i | p'q'\alpha' \rangle$$

as well as antisymmetrized matrix elements

$$\langle pq\alpha | (1 + P_{123} + P_{132}) V_{123}^i (1 + P_{123} + P_{132}) | p'q'\alpha' \rangle$$

- HDF5 file format for efficient I/O



<http://www.hdfgroup.org>

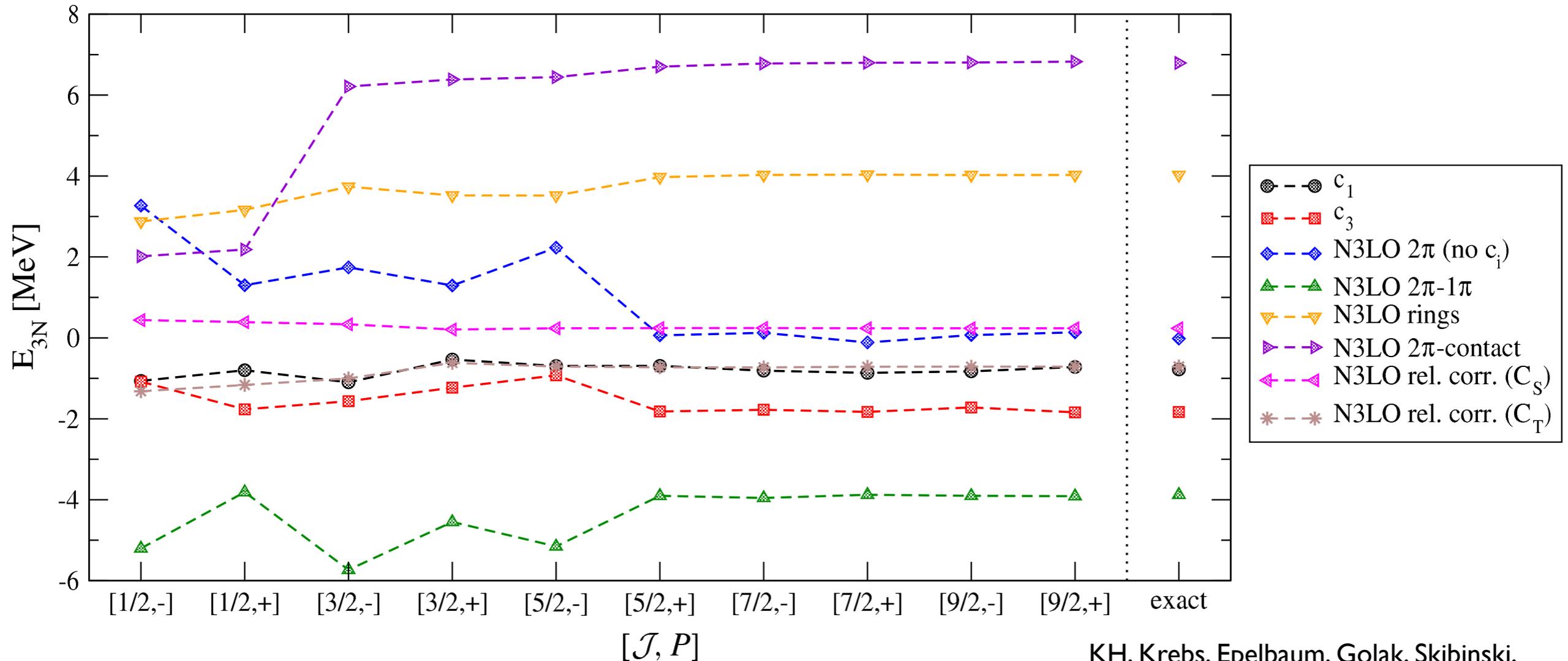
- current model space limits:

\mathcal{J}	\mathcal{T}	J_{\max}^{12}	size [GB]
1/2	1/2	8	1.0
3/2	1/2	8	3.2
5/2	1/2	8	6.2
7/2	1/2	7	6.9
9/2	1/2	6	6.2
1/2	3/2	8	0.3
3/2	3/2	8	0.8
5/2	3/2	8	1.8
7/2	3/2	7	1.8
9/2	3/2	6	1.8

~ 0.5 TB

Partial wave convergence: energy of infinite matter in Hartree-Fock approximation

neutron matter:

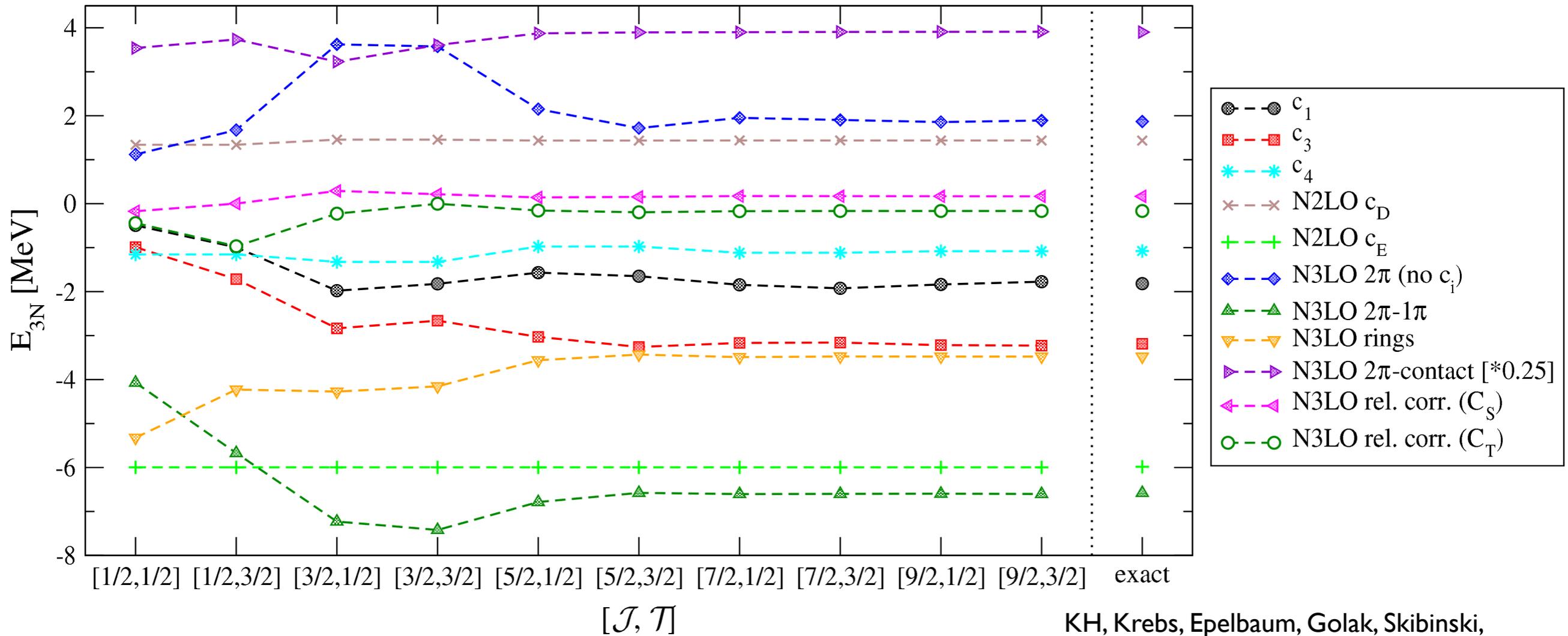


KH, Krebs, Epelbaum, Golak, Skibinski,
arXiv:1502.02977

- in PNM only matrix elements with $\mathcal{T} = 3/2$ contribute
- resummation up to $\mathcal{J} = 9/2$ leads to well converged results
- essentially perfect agreement with 'exact' results (cf. PRC88, 025802)

Partial wave convergence: energy of infinite matter in Hartree-Fock approximation

symmetric nuclear matter:

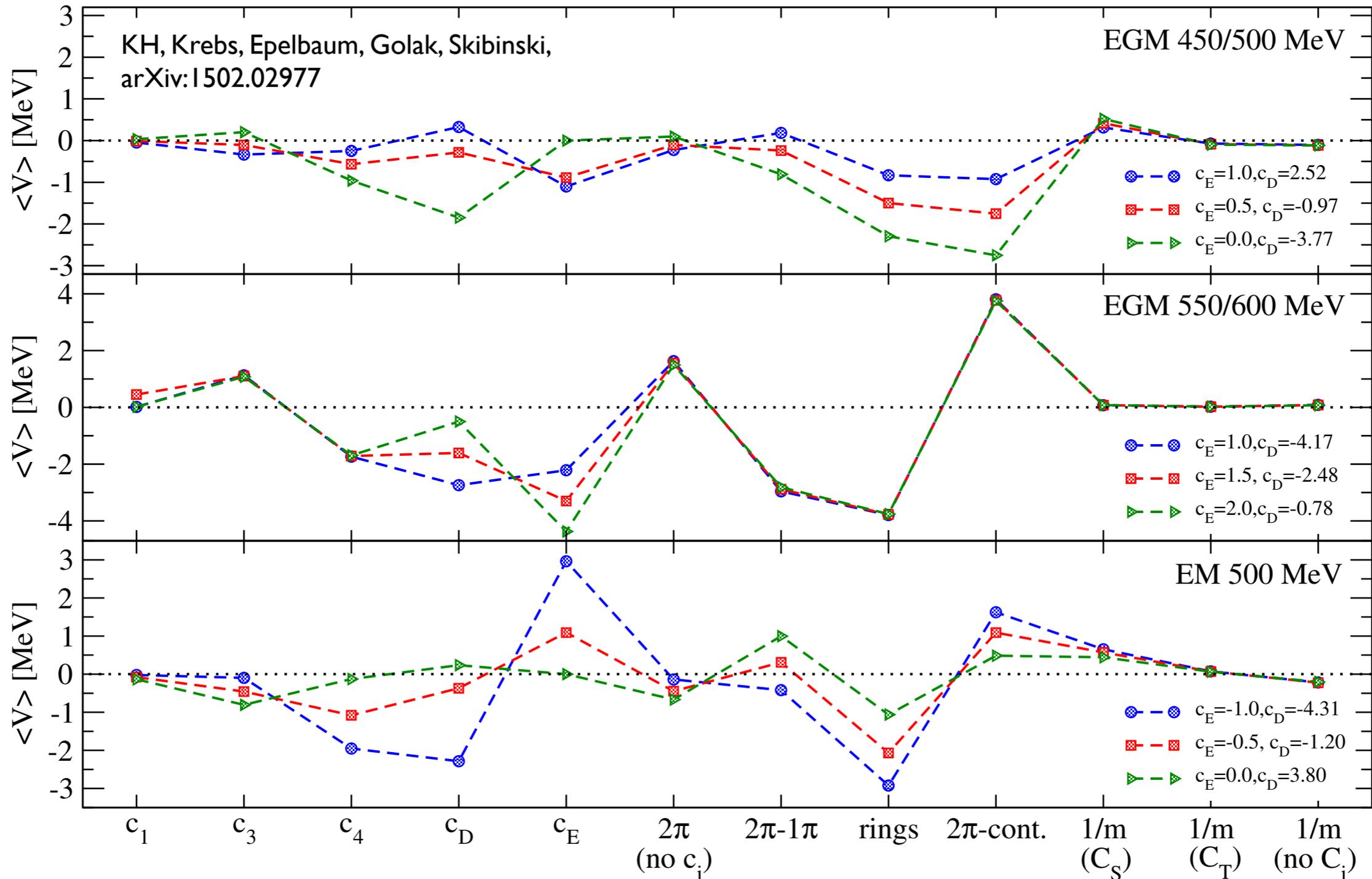


KH, Krebs, Epelbaum, Golak, Skibinski,
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Contributions of individual topologies in ${}^3\text{H}$

for specific choices of NN interactions and regulator functions!



- contributions of individual contributions depend sensitively on details
- N3LO contributions not suppressed compared to N2LO
- perturbativeness of 3NF strongly depends on NN interaction

Future directions: Incorporation in different many-body frameworks

Hyperspherical harmonics



Faddeev,
Faddeev-Yakubovski

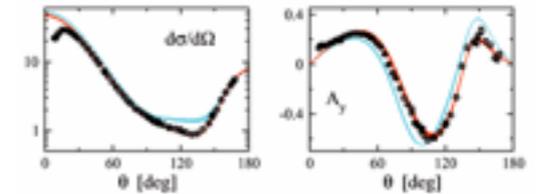


FIG. 4: Nd elastic observables at 65 MeV.

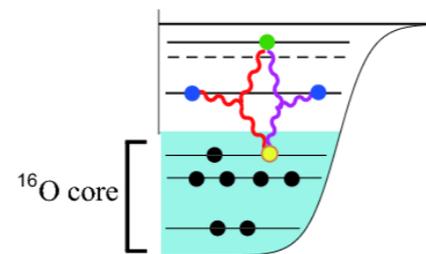
no-core shell model



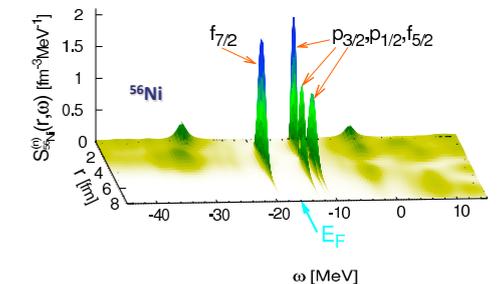
coupled cluster method

$$|\Psi\rangle = e^{\hat{T}}|\Phi_0\rangle = \left(1 + \hat{T} + \frac{1}{2}\hat{T}^2 + \frac{1}{3!}\hat{T}^3 + \dots\right)|\Phi_0\rangle,$$

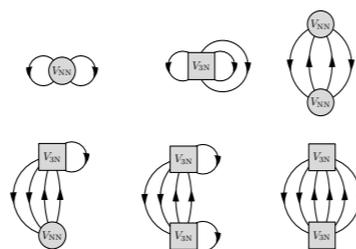
valence shell model



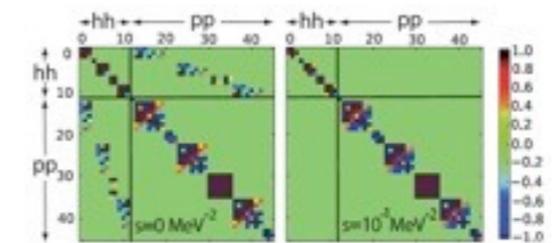
Self-consistent
Greens function



Many-body
perturbation theory



In-medium SRG



Required inputs:

1. **consistent** NN and 3N forces at N³LO in partial-wave-decomposed form
2. **softened** forces for judging approximations and pushing to heavier nuclei

Future directions: Different regularization schemes

Goal of regularization:
Separate long- from short-range physics

1. non-local regularization: $V_{\text{NN}}(p, p') \sim \exp \left[-\frac{p^{2n} + p'^{2n}}{\Lambda^{2n}} \right]$

2. local regularization: $V_{\text{NN}}(r) \sim \left(1 - \exp \left[-\frac{r^n}{R_0^n} \right] \right)$

3. hybrid strategy: regularize long-range parts locally and short-range distance non-locally (see talk by H. Krebs)

-
- different choices regulate short range physics in different ways
 - important to explore various alternatives
 - **need to implement according regularizations in 3NF**

Regularization schemes for 3NF

1. non-local regularization:

$$V_{3N}(p, q, p', q') \sim \exp \left[-\frac{p^2 + 3/4q^2}{\Lambda^2} \right] \exp \left[-\frac{p'^2 + 3/4q'^2}{\Lambda^2} \right]$$

- multiplicative (no partial-wave mixing), trivial to apply
- calculated matrix elements up to N3LO can be used immediately

2. local regularization:

$$V_{3N}(\mathbf{r}_{12}, \mathbf{r}_{23}, \mathbf{r}_{13}) \sim \left(1 - \exp \left[\frac{r_{12}^2}{R_0^2} \right] \right)^n \left(1 - \exp \left[\frac{r_{23}^2}{R_0^2} \right] \right)^n \left(1 - \exp \left[\frac{r_{13}^2}{R_0^2} \right] \right)^n$$

- partial wave mixing, application of regulator non-trivial in partial-wave basis
- different possibilities to calculate 3NF partial wave matrix elements:
 - ★ decompose 3N in coordinate space and then fourier transform
 - ★ perform folding integrals in momentum space partial wave basis

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Work in progress. Stay tuned!

Summary

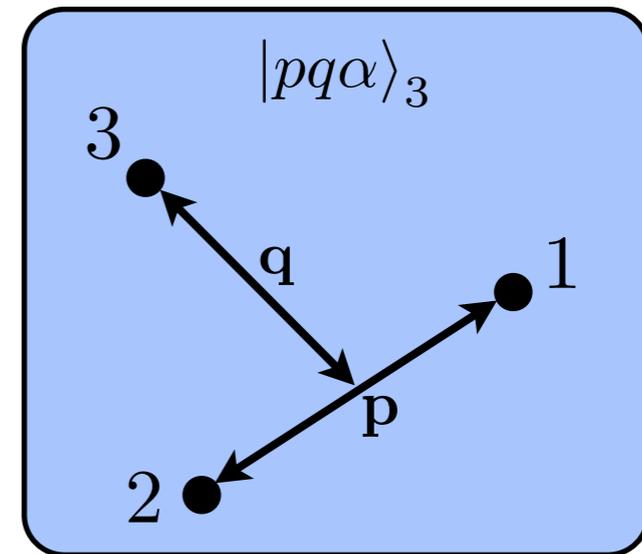
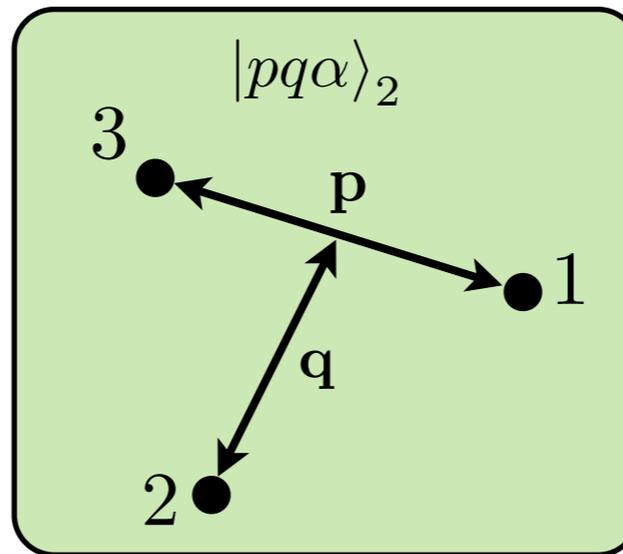
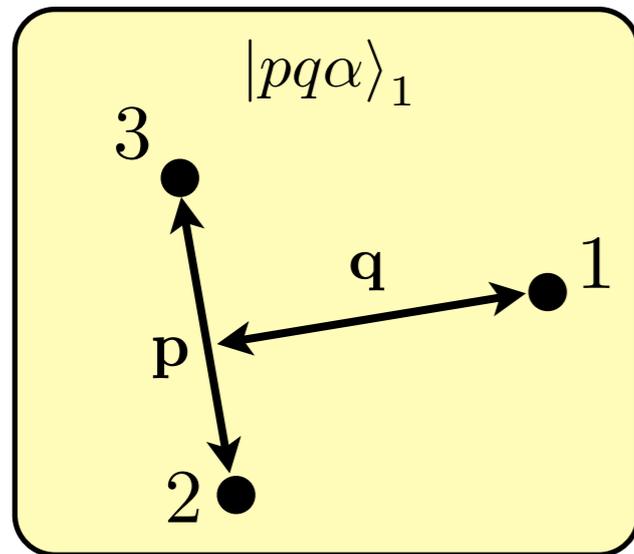
- calculated unregularized partial-wave matrix elements of chiral 3NF up to N3LO
- tested partial wave convergence in nuclear matter (Hartree-Fock appr.)
- contributions to ${}^3\text{H}$ of individual 3NF topologies scheme dependent
- transformation of matrix elements to HO basis straightforward
→ calculations of finite nuclei (shell model, coupled cluster, IM-SRG, SCGF...)

Outlook

- inclusion of non-local regulators trivial, impl. of local regulators in progress
- efficient calculation of nuclear currents?
- chiral forces based on other power counting?
- chiral forces at N4LO?
- delta-full EFT?

Thank you!

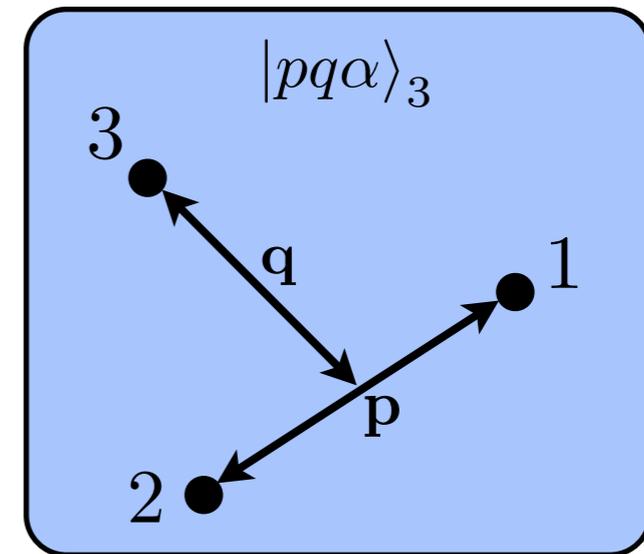
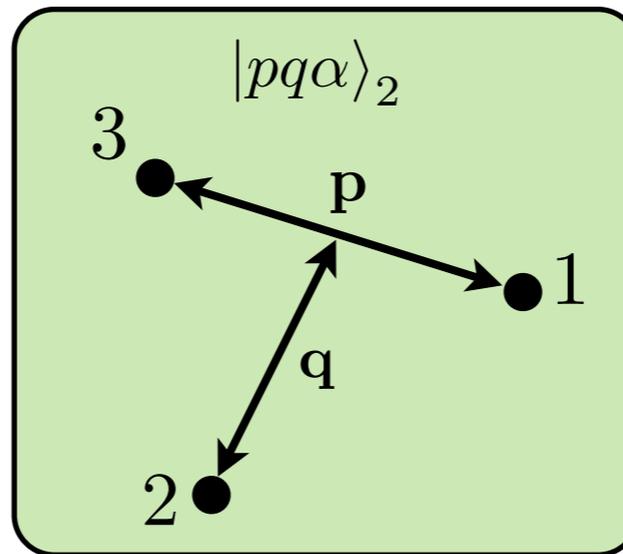
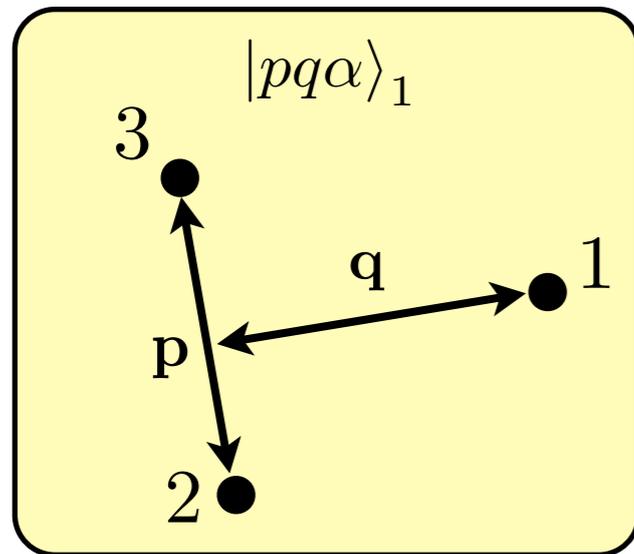
RG evolution of 3N interactions in momentum space



- represent interaction in basis $|pq\alpha\rangle_i \equiv |p_i q_i; [(LS)J(l s_i)j] \mathcal{J} \mathcal{J}_z (T t_i) \mathcal{T} \mathcal{T}_z\rangle$
- explicit equations for NN and 3N flow equations

$$\begin{aligned} \frac{dV_{ij}}{ds} &= [[T_{ij}, V_{ij}], T_{ij} + V_{ij}], \\ \frac{dV_{123}}{ds} &= [[T_{12}, V_{12}], V_{13} + V_{23} + V_{123}] \\ &\quad + [[T_{13}, V_{13}], V_{12} + V_{23} + V_{123}] \\ &\quad + [[T_{23}, V_{23}], V_{12} + V_{13} + V_{123}] \\ &\quad + [[T_{\text{rel}}, V_{123}], H_s] \end{aligned}$$

RG evolution of 3N interactions in momentum space



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$$\frac{dV_{ij}}{ds} = [[T_{ij}, V_{ij}], T_{ij} + V_{ij}],$$

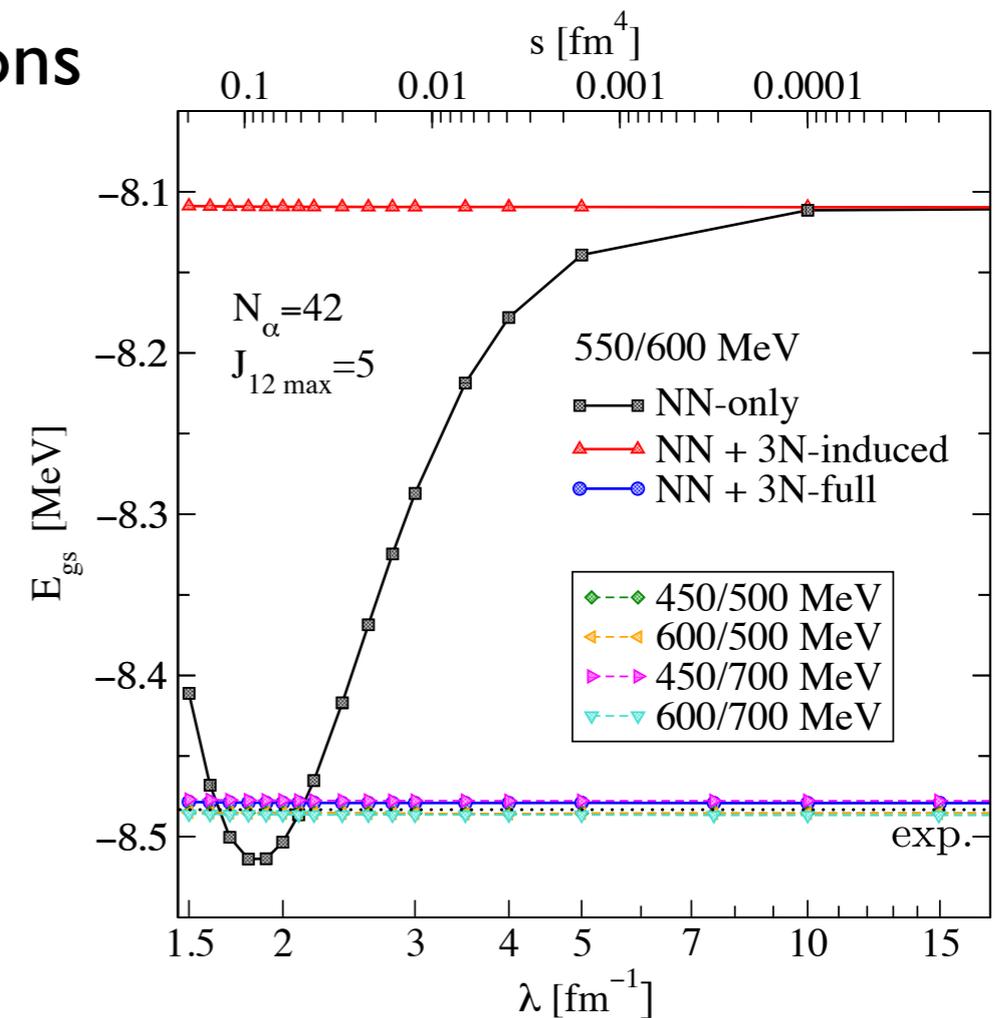
$$\frac{dV_{123}}{ds} = [[T_{12}, V_{12}], V_{13} + V_{23} + V_{123}]$$

$$+ [[T_{13}, V_{13}], V_{12} + V_{23} + V_{123}]$$

$$+ [[T_{23}, V_{23}], V_{12} + V_{13} + V_{123}]$$

$$+ [[T_{\text{rel}}, V_{123}], H_s]$$

Bogner, Furnstahl, Perry PRC 75, 061001(R) (2007)



Hebeler PRC(R) 85, 021002 (2012)

SRG flow equations of NN and 3N forces in momentum basis

$$\frac{dH_s}{ds} = [\eta_s, H_s] \quad \eta_s = [T_{\text{rel}}, H_s]$$

$$H = T + V_{12} + V_{13} + V_{23} + V_{123}$$

- spectators correspond to delta functions, matrix representation of H_s ill-defined
- **solution**: explicit separation of NN and 3N flow equations

$$\begin{aligned} \frac{dV_{ij}}{ds} &= [[T_{ij}, V_{ij}], T_{ij} + V_{ij}], \\ \frac{dV_{123}}{ds} &= [[T_{12}, V_{12}], V_{13} + V_{23} + V_{123}] \\ &\quad + [[T_{13}, V_{13}], V_{12} + V_{23} + V_{123}] \\ &\quad + [[T_{23}, V_{23}], V_{12} + V_{13} + V_{123}] \\ &\quad + [[T_{\text{rel}}, V_{123}], H_s] \end{aligned}$$

- only connected terms remain in $\frac{dV_{123}}{ds}$, 'dangerous' delta functions cancel

SRG evolution in momentum space

- evolve the antisymmetrized 3N interaction

$$\bar{V}_{123} =_i \langle pq\alpha | (1 + P_{123} + P_{132}) V_{123}^{(i)} (1 + P_{123} + P_{132}) | p'q'\alpha' \rangle_i$$

- embed NN interaction in 3N basis:

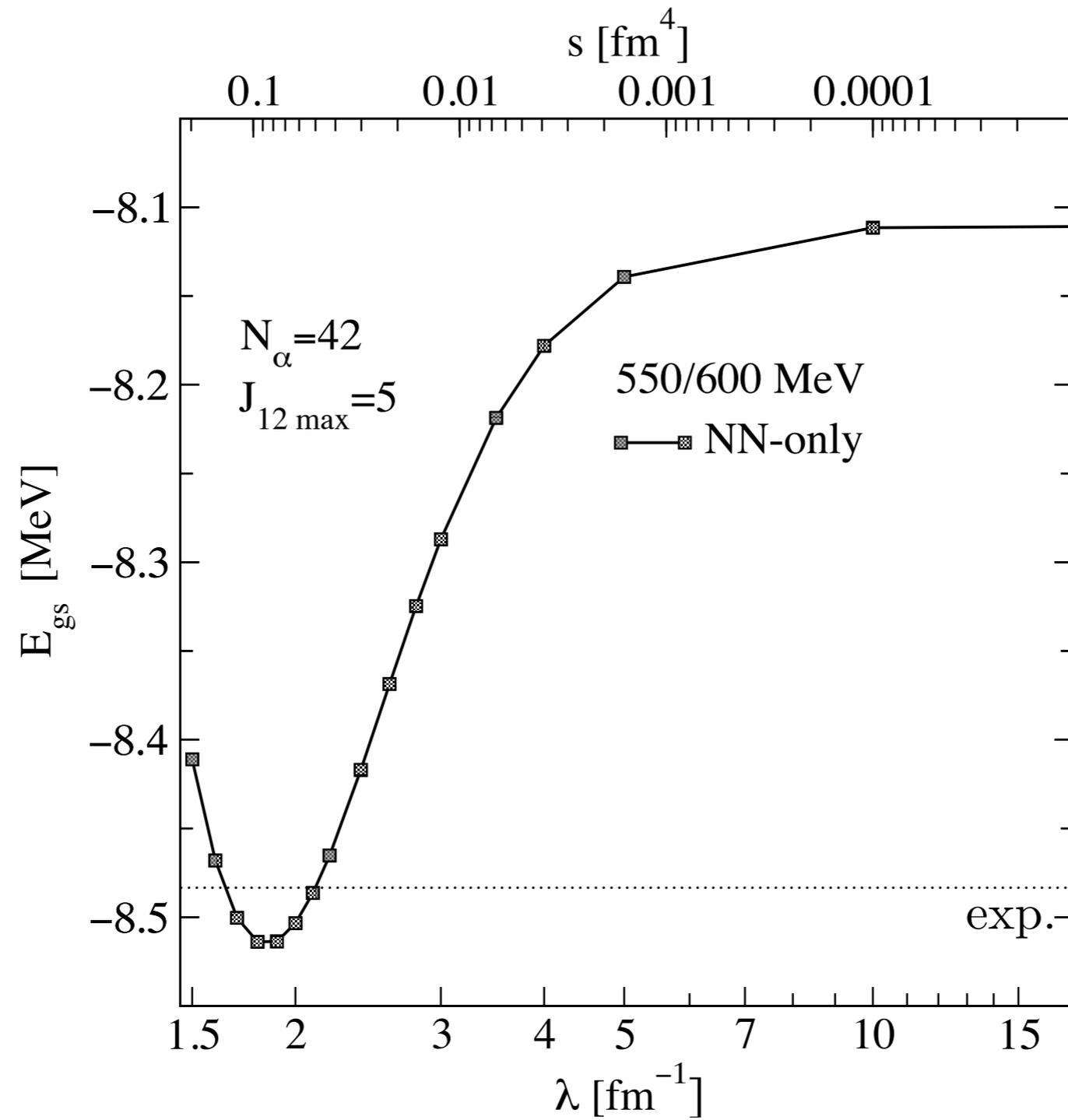
$$V_{13} = P_{123} V_{12} P_{132}, \quad V_{23} = P_{132} V_{12} P_{123}$$

$$\text{with } {}_3 \langle pq\alpha | V_{12} | p'q'\alpha' \rangle_3 = \langle p\tilde{\alpha} | V_{\text{NN}} | p'\tilde{\alpha}' \rangle \delta(q - q') / q^2$$

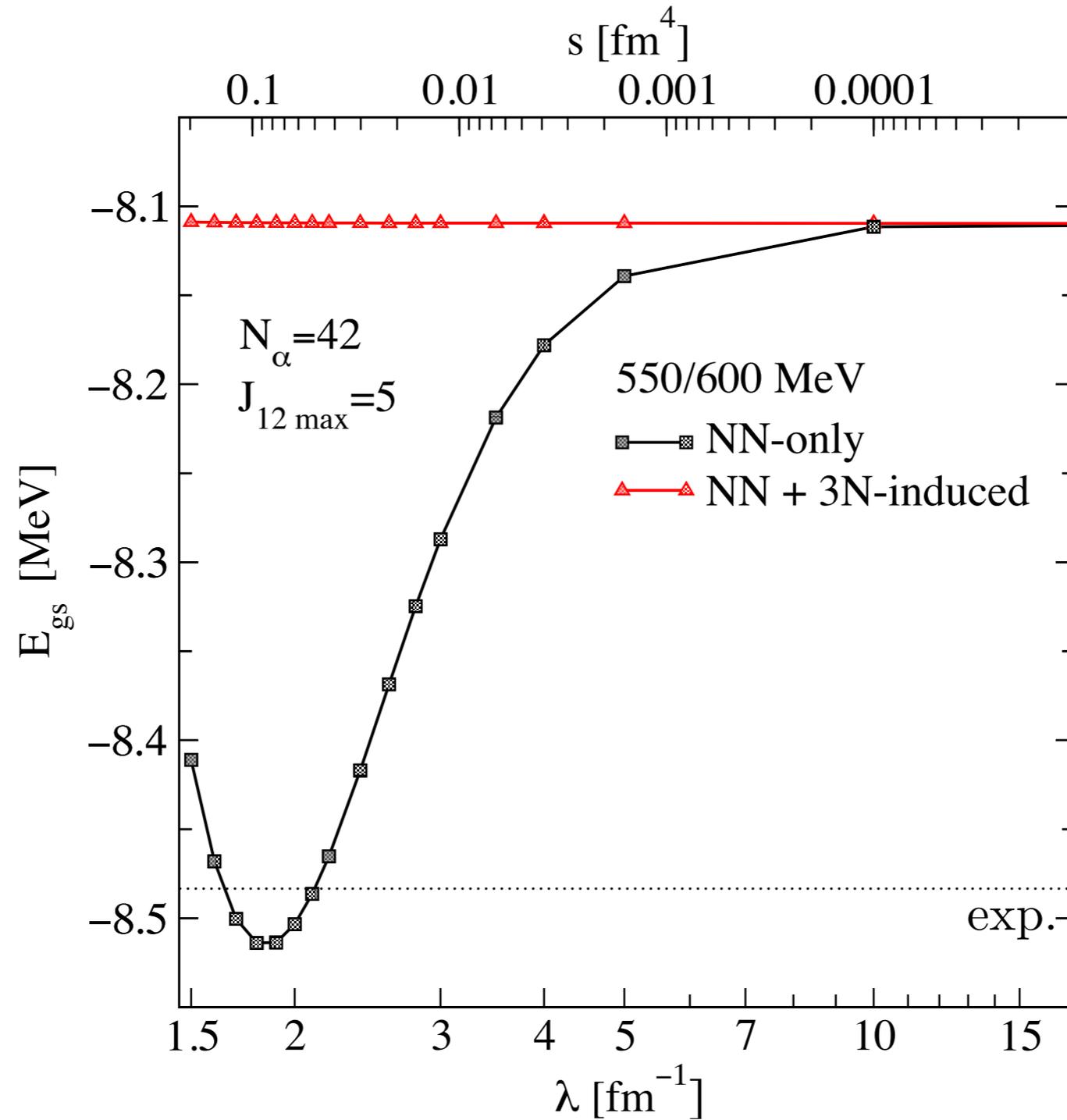
- use $P_{123} \bar{V}_{123} = P_{132} \bar{V}_{123} = \bar{V}_{123}$

$$\begin{aligned} \Rightarrow \quad d\bar{V}_{123}/ds &= C_1(s, T, V_{\text{NN}}, P) \\ &+ C_2(s, T, V_{\text{NN}}, \bar{V}_{123}, P) \\ &+ C_3(s, T, \bar{V}_{123}) \end{aligned}$$

SRG evolution of 3N interactions in momentum space: Results for the Triton



SRG evolution of 3N interactions in momentum space: Results for the Triton

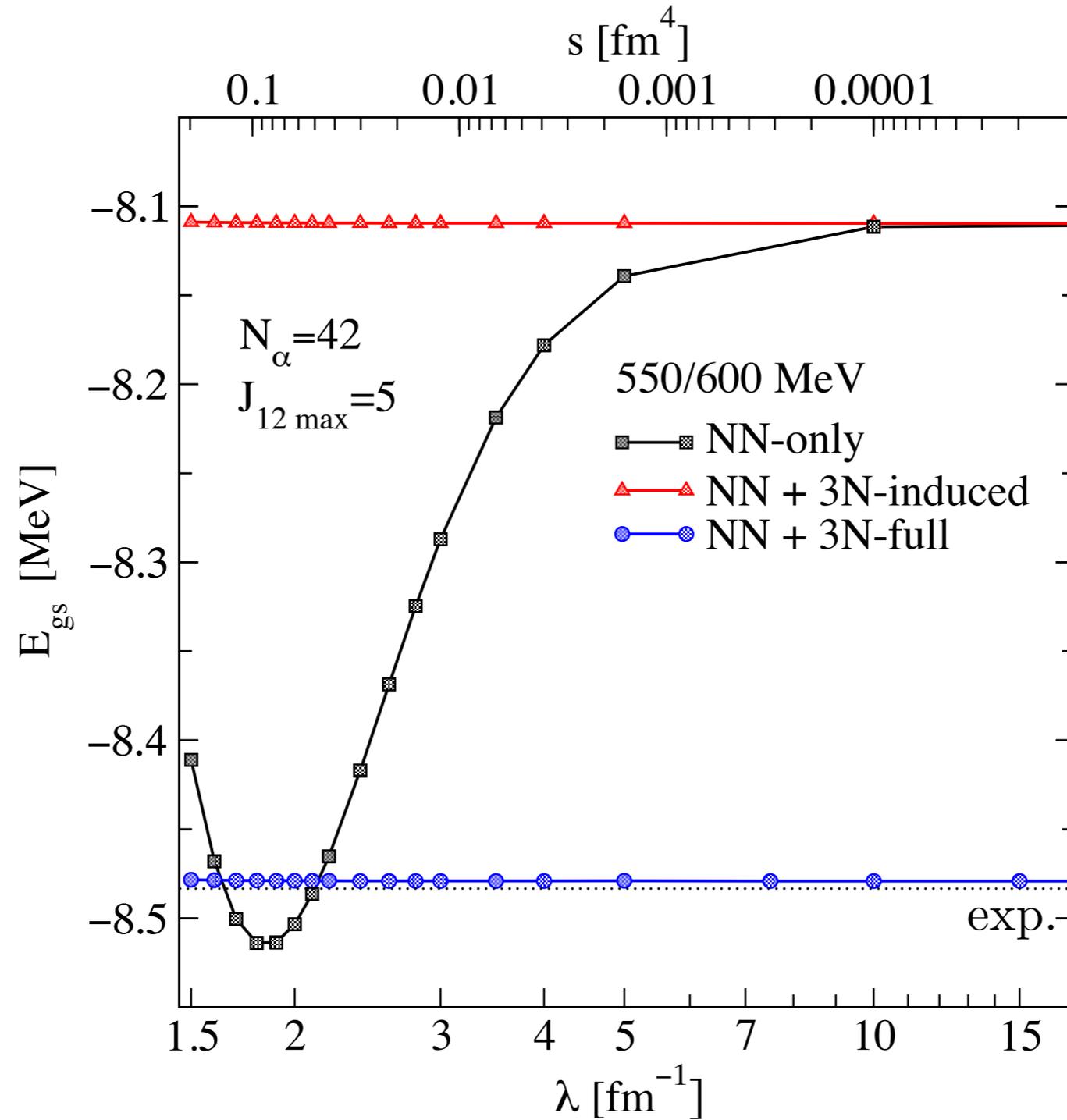


Hebeler PRC(R) 85, 021002 (2012)

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Invariance of E_{gs}^{3H} within $\leq 1 \text{ eV}$ for consistent chiral interactions at $N^2\text{LO}$

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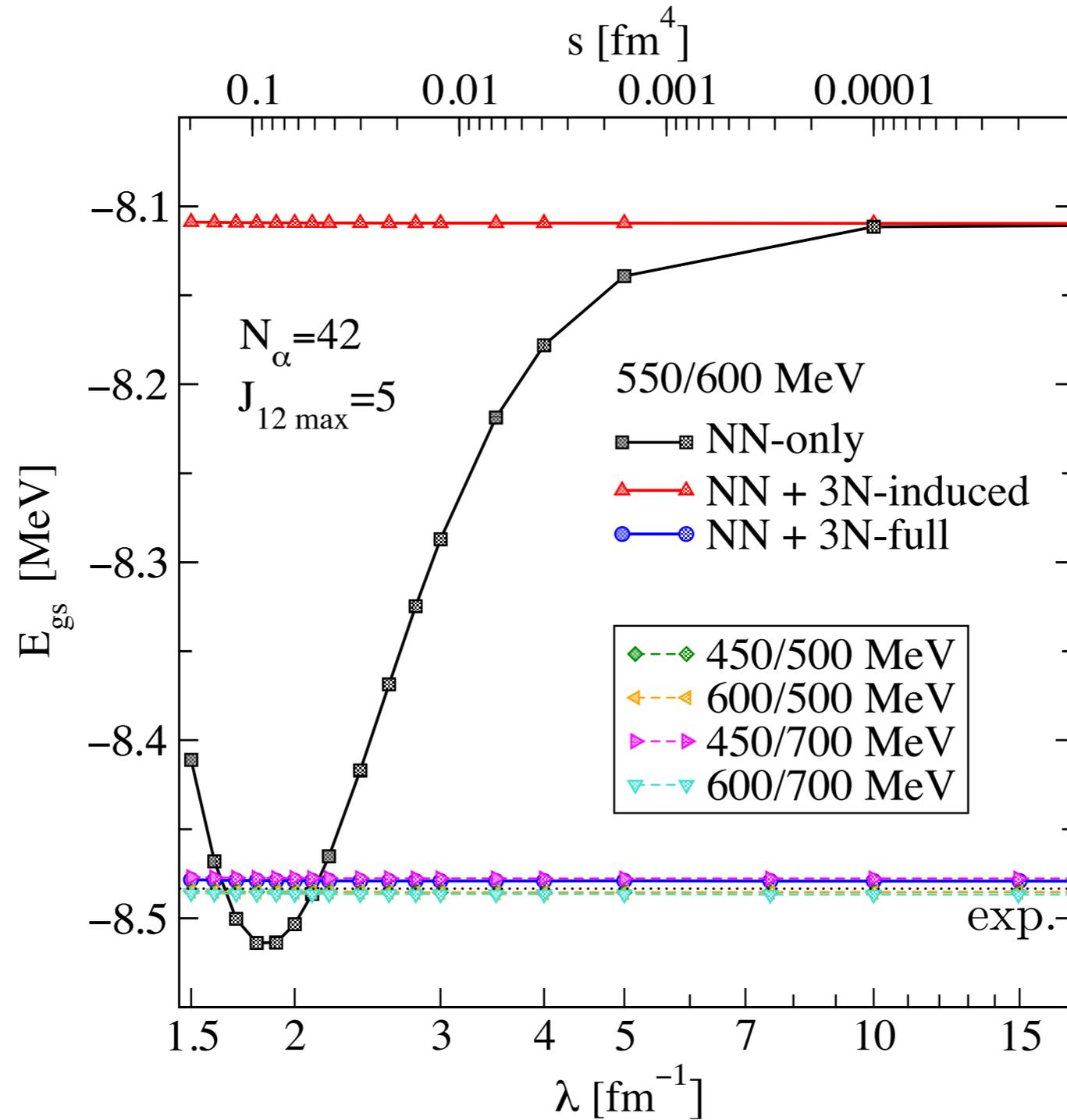


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