

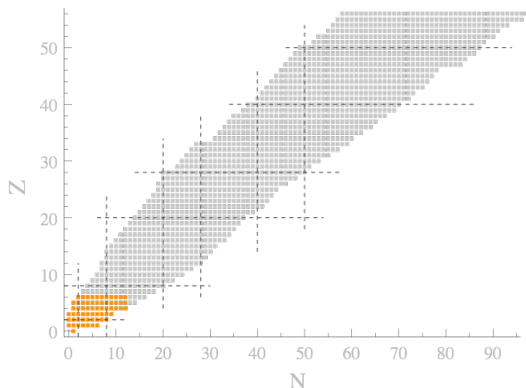
Self-Consistent Green's Functions Calculations of Charge Radii and Distribution up to $A=140$

Pierre Arthuis
University of Surrey, UK

Progress in *Ab Initio* Techniques in Nuclear Physics
TRIUMF, Vancouver, BC – 4 March 2020



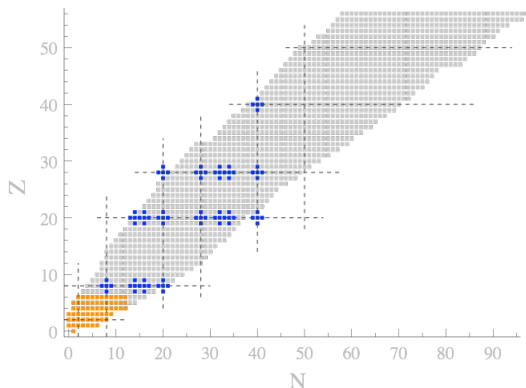
Ab initio methods for the heavy sector



"Exact" methods (80's)

- GFMC, NCSM, FY, HH...

Courtesy of V. Somà, T. Duguet



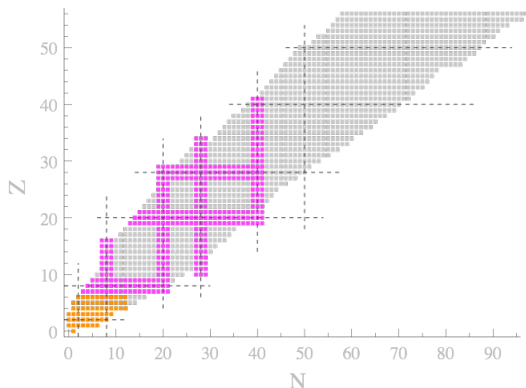
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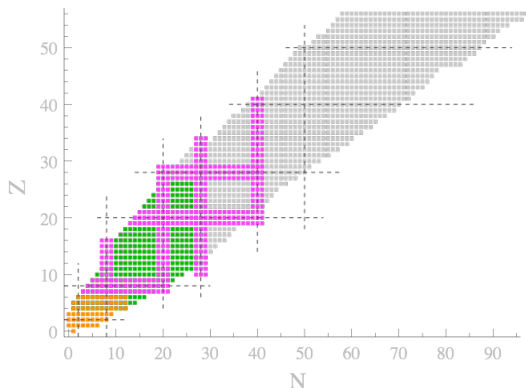
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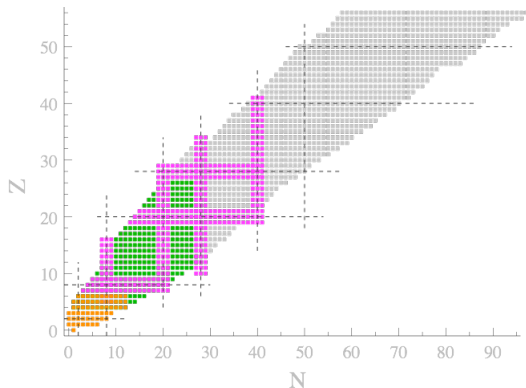
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Ab initio shell model (2014)

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What does prevent us from going further up in mass?

Nuclear interaction accuracy

- Chiral potentials tend to largely overbind in mid-mass nuclei
- Uncertainties driven by the interaction

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Model-space-size limitations

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How to circumvent it

- New interactions: $NNLO_{sat}$, $NN+3N_{Inl}$, $1.8/2.0(EM)$, $\Delta NNLO_{GO}$...
- Work on storage cost: CPD, THC, IT, alternative $NO2B$...
- Promising but requires to adapt the existing set-ups, takes time

***Ab initio* method**

- Limit the computational cost: single-reference method
- Fewer and fewer closed-shell nuclei: symmetry-breaking method
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NNLOsat

Concept behind SCGFs

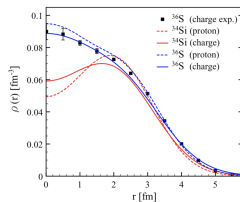
Recast the solution of the SE as a set of one- to A -body Green's functions

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Through the one-body Green's function:

- Ground-state energy
- One-body observables (radii, densities...)
- Spectroscopy of the $A \pm 1$ -body systems
- Elastic nucleon-nucleus scattering



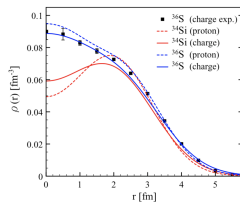
[Duguet et al., 2017]

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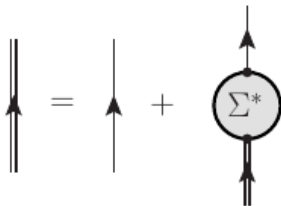


[Duguet et al., 2017]

See Mehdi Drissi's talk tomorrow for neutron matter SCGF!

How to proceed

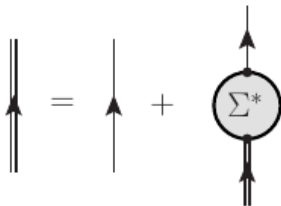
Solve the Dyson equation self-consistently: infinite diagram series resummation



$$g_{\alpha\beta}(\omega) = g_{\alpha\beta}^{(0)}(\omega) + \sum_{\gamma\delta} g_{\alpha\gamma}^{(0)}(\omega) \Sigma_{\gamma\delta}^*(\omega) g_{\delta\beta}(\omega)$$

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Algebraic Diagrammatic Construct at order (n) [ADC(n)]

- Perturbative order + necessary diagrams to preserve analytical properties
- ADC(∞) gives exact result
- In practice numerically: ADC(3) for Dyson, ADC(2) for Gorkov

Our case study: Xenon and Tin isotopes

[Arthuis, Barbieri, Vorabbi, Finelli, arXiv:2002.02214 (2020)]

What are the present-day limits?

- ^{100}Sn , ^{132}Sn with CC: proof-of-principle calculations [Binder *et al.*, 2013]
- ^{100}Sn to ^{111}Sn with VS-IMSRG: experimentally relevant [Morris *et al.*, 2018]

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- Spherical or near-spherical nuclei

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Our goals

- ^{100}Sn , ^{132}Sn : doubly-closed shell milestones, convergence study
- ^{132}Xe , ^{136}Xe and ^{138}Xe : open-shell nuclei of experimental interest
- Use interactions with good accuracy in the mid-mass area
- Look at both ground-state energy and radii

Model space

- Harmonic oscillator basis
- Truncated at $N_{\max} = 13$ ($N_{\max} = 11$ to check convergence)

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- Ideal case: $E_{3\max} = 3 \times N_{\max}$
- Too costly: in practice $E_{3\max} = 16$ (14)

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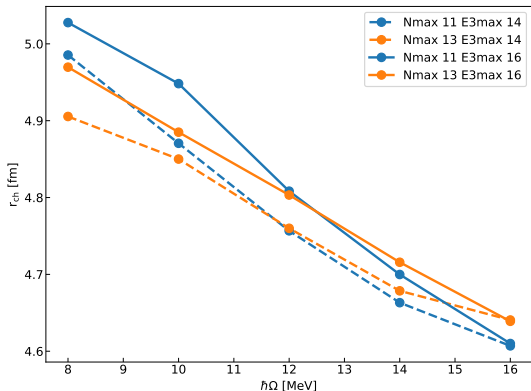
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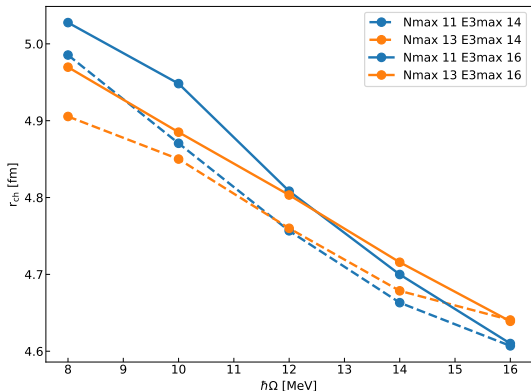
First exploratory calculations

- Variational minimum for $\hbar\Omega$ between 10 and 16 MeV
- Ground-state energy not converged w.r.t. model space size
- Can we say something about the radius?



[Arthuis, Barbieri, Vorabbi, Finelli, arXiv:2002.02214 (2020)]

- Possible to take a conservative estimate from our calculations
- Low influence of the E3max truncation on the expected radius value
- ADC(3) included in the error bars for Sn isotopes



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Possible to extract radii with NNLOsat despite small model space

	SCGF	Exp.
^{100}Sn	4.525 – 4.707	
^{132}Sn	4.725 – 4.956	4.7093
^{132}Xe	4.700 – 4.948	4.7859
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Table: Charge radii in fm obtained from SCGF calculations.

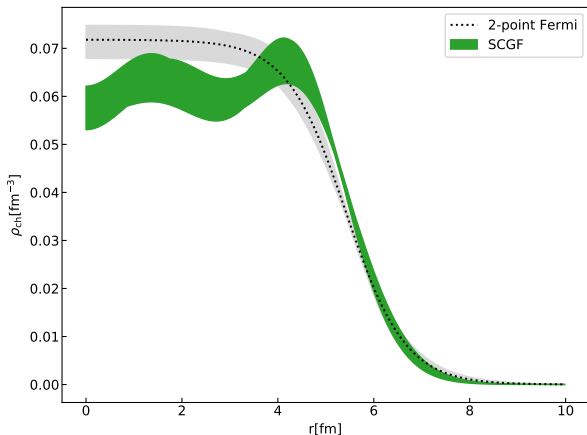
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NNLOsat confirmed to work in the heavy mass regime



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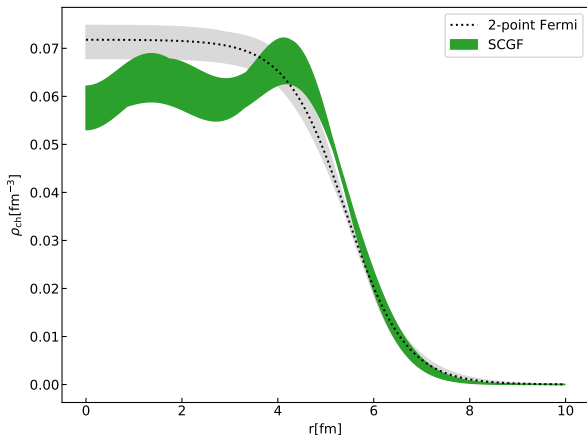
Hamiltonian

- Chiral EFT
Bare NNLO_{sat}
- Effective 3NF

Basis parameters

- $\hbar\Omega = 10$ to 14 MeV
- $N_{max} = 11 - 13$
- $E_{3max} = 16$

- Radius very close to the experiment: 4.82 ± 0.12 fm vs. $4.79^{+0.11}_{-0.08}$ fm
- Two-point Fermi distribution insufficient to describe expected behaviour



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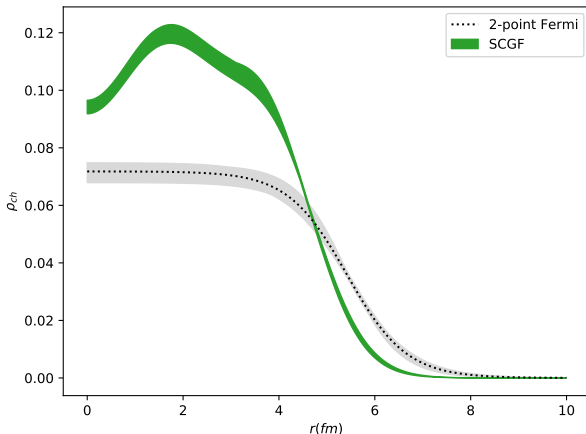
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Need more precise experiment to determine the charge distribution



[Arthuis, Barbieri, Vorabbi, Finelli, arXiv:2002.02214 (2020)]

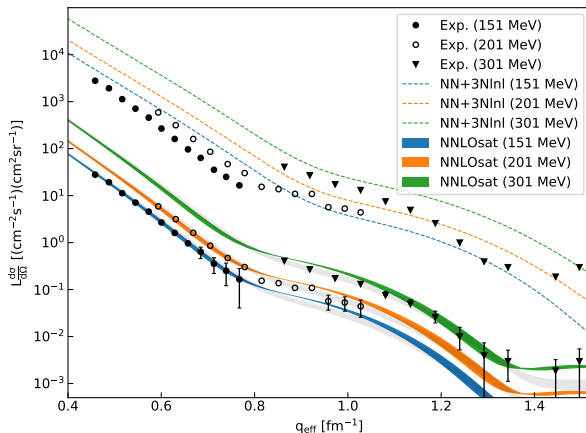
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- Radius largely underestimated: 4.070 ± 0.045 fm
- Inability of NN+3Nlnl to reproduce the experimental properties



[Arthuis, Barbieri, Vorabbi, Finelli, arXiv:2002.02214 (2020)]

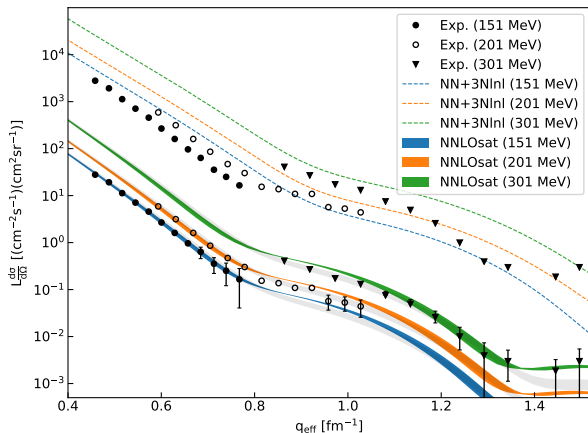
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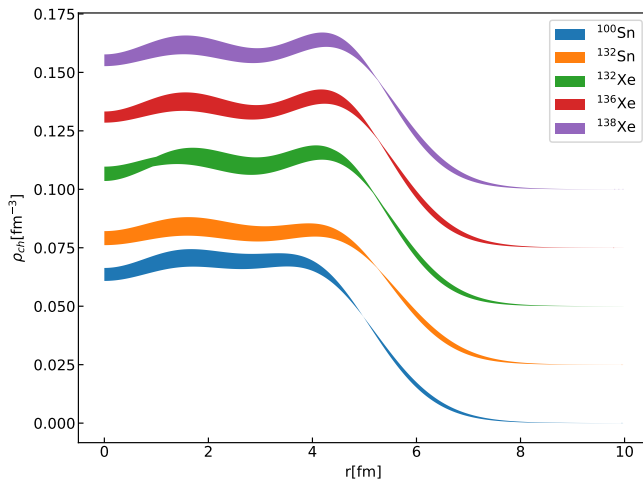
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Theory and experiment can be compared in the heavy mass sector



- Charge distributions consistent over the different isotopes
- Paves the way to density distribution studies of heavy nuclei

First *ab initio* study in the $A \sim 140$ region of the nuclear chart

- Radii and charge densities are already accessible
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The future of *ab initio* methods

- Experiments can already be constrained by *ab initio* calculations
- Every improvement in ME storage is a move forward in mass



C. Barbieri
M. Drissi
J. Keeble
A. Rios



M. Vorabbi



P. Finelli

Thank you for your attention!