

# Generalized Luscher method: an avenue for ab initio calculation of low-energy nuclear scattering and reactions

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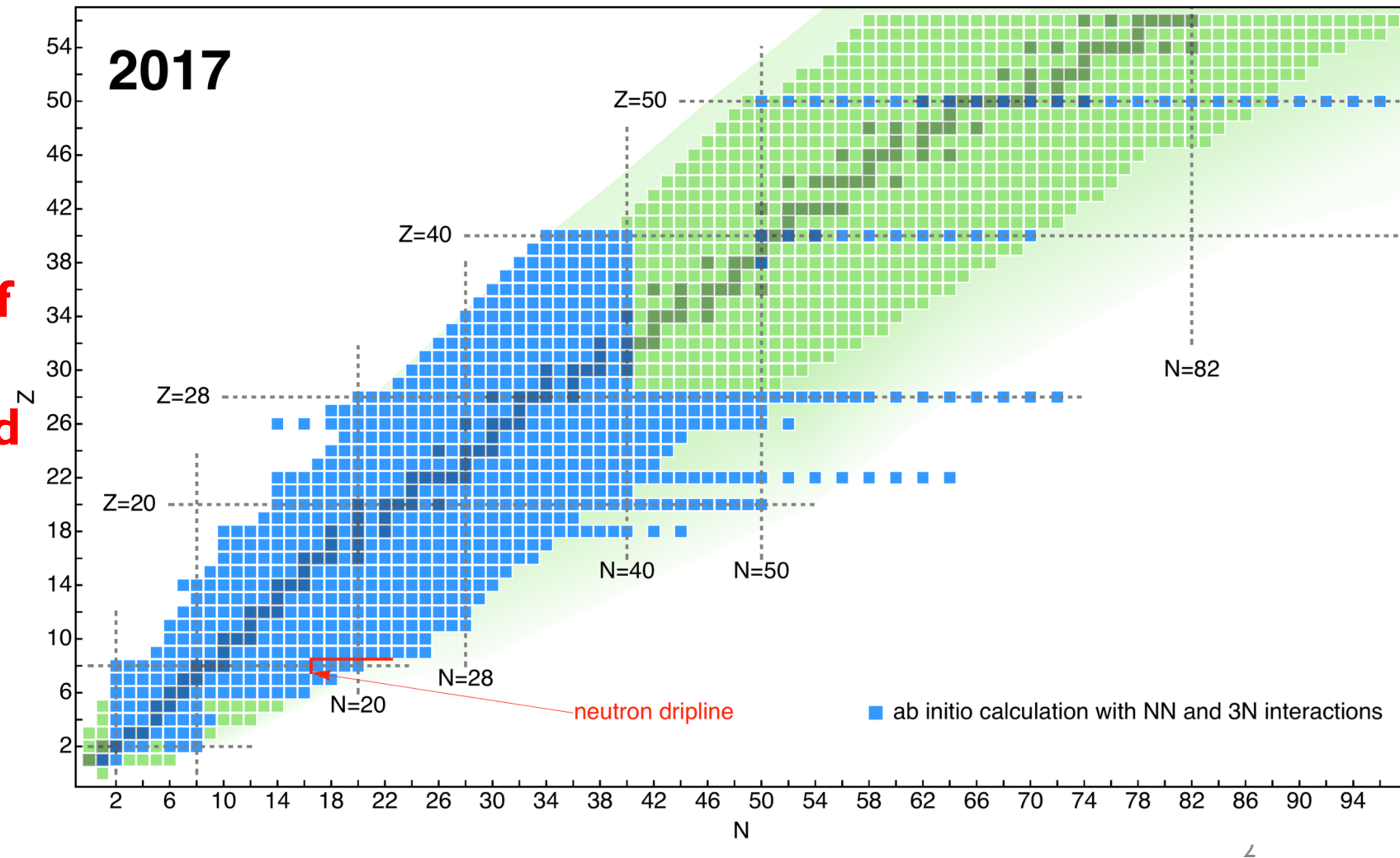
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*In collaboration with Chan Gwak, Ragnar Stroberg, Petr Navratil, and Jason Holt;  
In collaboration with J. Melendez, and R. Furnstahl*


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# The question

- Ab initio structure calculations have made amazing progress, while scattering/reactions are limited
- **Can we take advantage of this bound-state-calculation's progress and compute scattering/reactions?**
- Drip-line nuclei needs consistent treatment of structure and scattering/reactions



# Key idea

$$\sigma_l(E) = \frac{4\pi}{p^2} \times (2l + 1) \sin^2 \delta_l(E)$$


- **Eigen-energies** of trapped projectile-target system **computed by ab initio energy calculation** → **scattering** info (**phase-shift**) at **those energies**
- Interpolating between eigen-energies gives the scattering phase-shift as function of scattering energy
- **The method should work with nuclear systems (e.g., neutron-nucleus) that can be computed by the ab initio structure methods.**

# Outline

- Our method generalizes the Luscher method used in Lattice QCD
  - Perfect **computer** experiment
  - Imperfect computer experiment
- **Benchmark** our scattering calculations against the existing ab initio calculation ( $n - \alpha$ )
- Show it also works for **heavier** system ( $n -^{24} O$ , as recently measured)
- Modeling of errors in the imperfect computer experiment
- Comparison to other methods
- Summary and Outlook: inelastic reactions, and role of R-matrix/potential model

# Luscher's method in Lattice QCD

Discrete **eigen-energies** for pi-pi in a  
finite volume gives  
the **phase shift at  
those energies**

$$(E, V_{Lattice}) \rightarrow \delta_l(E)$$

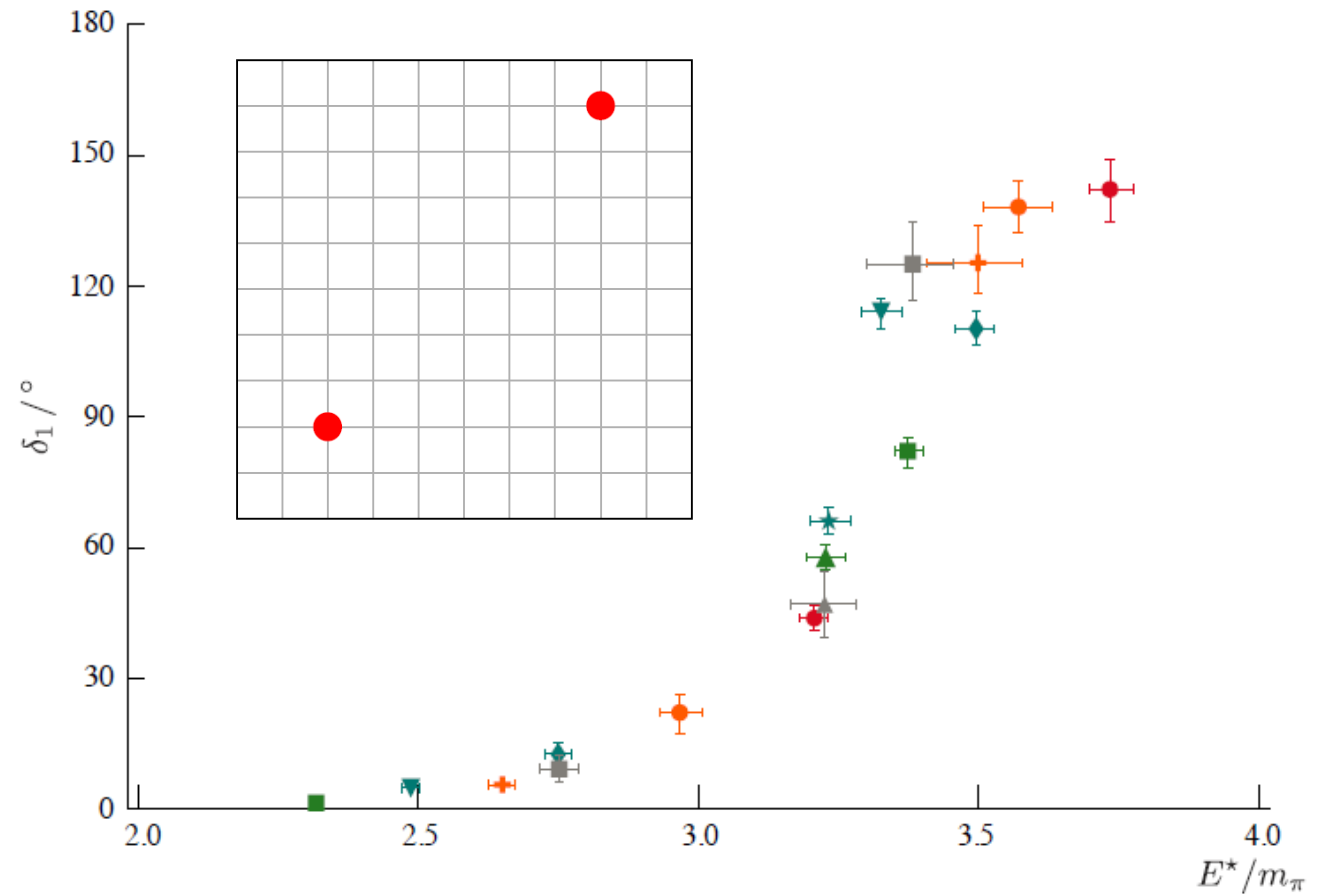


FIG. 14 Elastic  $I = 1$   $\pi\pi$  scattering phase-shifts in  $P$ -wave determined from finite-volume spectra computed the same  $m_\pi \sim 236$  MeV configurations as used in the calculation

Briceno et.al., RMP.90.025002 (2018)

However trapping nucleons in harmonic potential well is better suited for harmonic-oscillator-basis calculations

- Reduces degrees of freedom (DOF)  $\rightarrow$  make ab initio calculations feasible
- The center of mass (CM) and internal DOF are **decoupled**
- Lattice regulator breaks rotational invariance

There is a “universal” formula for two-cluster system at low energy  $\rightarrow$  BERW (Busch) formula

$$(E, \omega_T) \rightarrow \delta_l(E)$$

# BERW formula: intuition

Suppose we know the eigen-energy (**E**) of the system in a trap for s-wave.

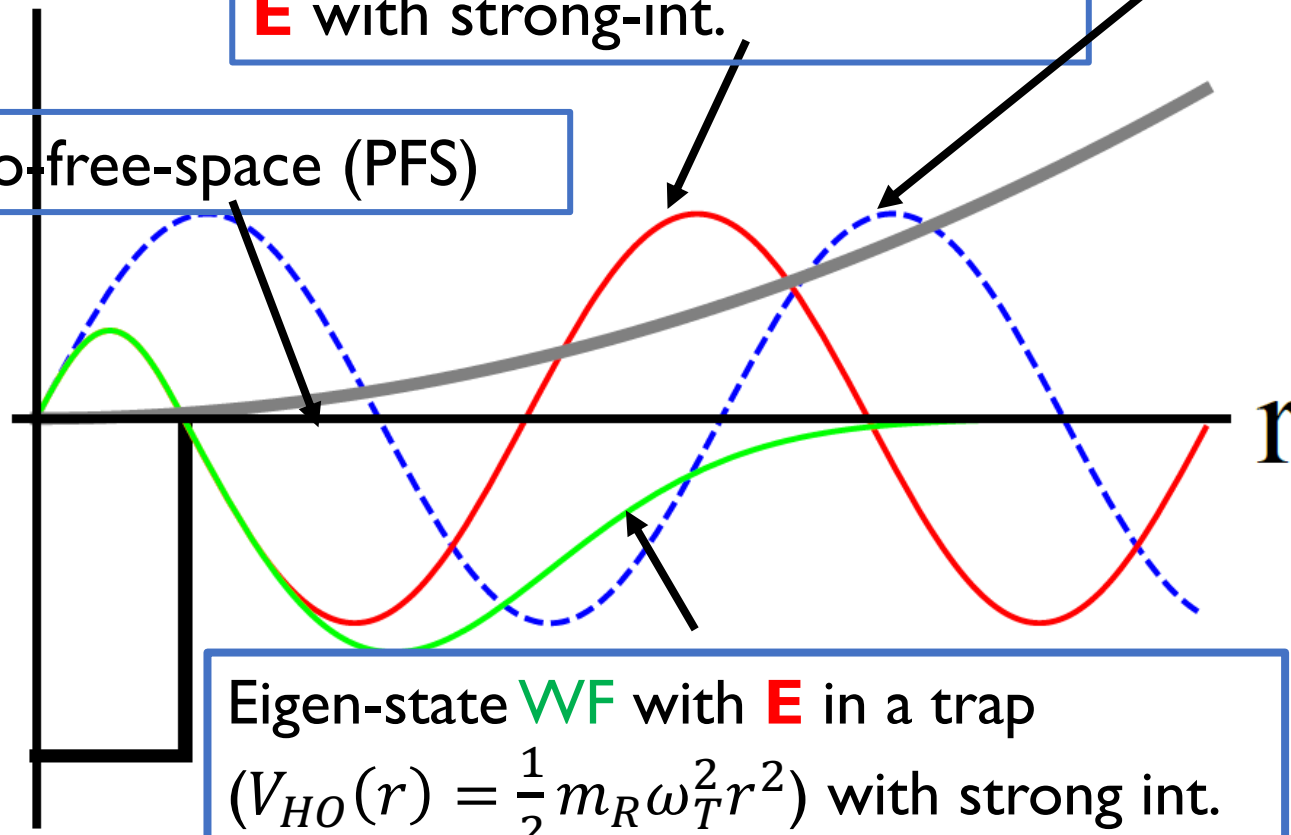
$$p^{2l+1} \cot \delta_l = (-1)^{l+1} (4 M_R \omega_T)^{l+1/2} \frac{\Gamma(\frac{3}{4} + \frac{l}{2} - \frac{E}{2\omega_T})}{\Gamma(\frac{1}{4} - \frac{l}{2} - \frac{E}{2\omega_T})}$$

$V$  or  $\psi$

Scattering **WF** in free space at **E** with strong-int.

Scattering wave function (**WF**) in free space at **E**

Pseudo-free-space (PFS)

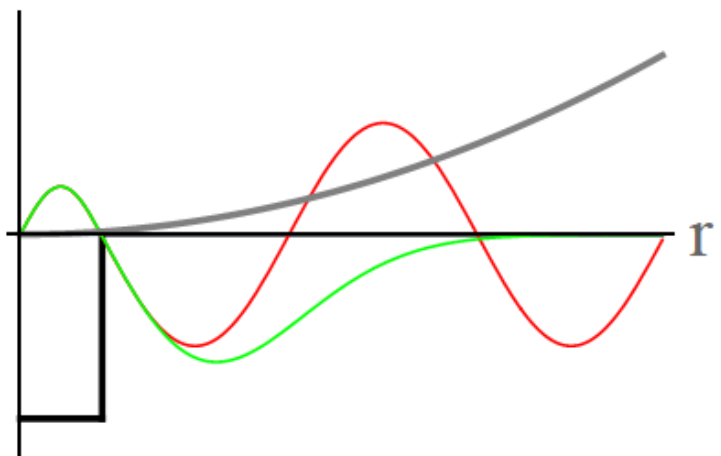


Eigen-state **WF** with **E** in a trap ( $V_{HO}(r) = \frac{1}{2} m_R \omega_T^2 r^2$ ) with strong int.

- The **full WF** dies off in a gaussian form. We can integrate Schrodinger equation from large distance inward, and get **WF** at PSF (i.e, WF determined by **E** and  $\omega_T$ ):  $\left[ -\frac{\partial_R^2}{2m_R} + V_{HO}(r) \right] \psi(r) = E\psi(r)$
- At PFS, at lowest order (or shallow trap limit,  $\omega_T \rightarrow 0$ ), the full WF is close to the scattering WF with strong interaction:  $\cos \delta j_0 + \sin \delta n_0$
- Matching them gives BERW formula
- BERW: left side  $\rightarrow$  strong int.; right side  $\rightarrow$  (bound. cond.) long-dis. physics**

# BERW formula: the UV-defect

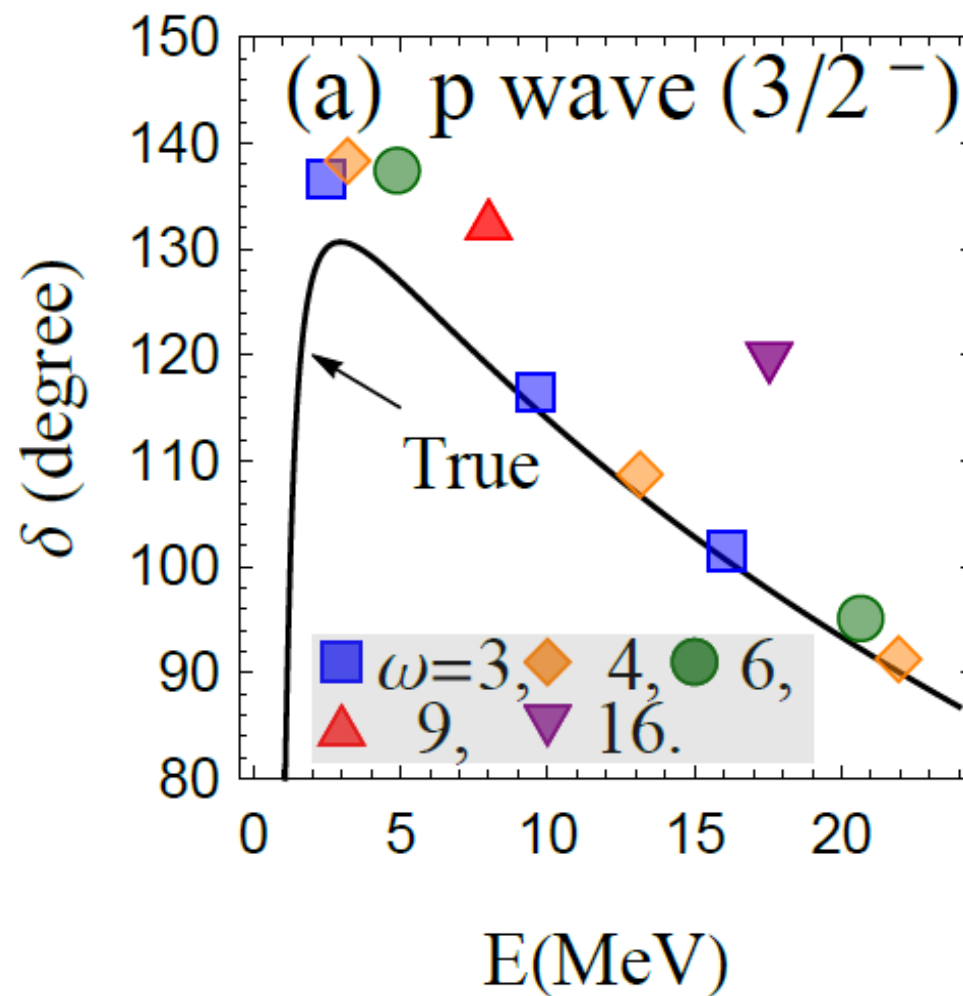
$V$  or  $\psi$



- However, PFS is not real free space.
- And how to separate the IR-physics (boundary condition) and the strong-interaction physics?
- Effective field theory (EFT) separates IR and UV in scattering amplitude

XZ, 1905.05275.

$n - \alpha$  two-body potential model





# Cure the BERW formula with the UV-defect

- T-matrix in free space  $S_l = e^{2i\delta_l} \rightarrow T_l \sim \frac{S_l - 1}{2i}$   

$$T(E) \sim \frac{1}{p^{2l+1} \cot \delta_l - U(E)}$$
- Non-analyticity from IR physics:  $U(E) = ip^{2l+1}$  (note  $ip = \sqrt{-E}$ )
- Analytical expansion from UV physics, i.e., effective range expansion (ERE):

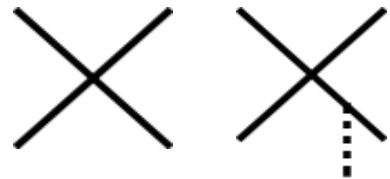
$$p^{2l+1} \cot \delta_l = \sum C_{i=0,j} p^{2j}$$

- EFT derivation :



# Cure the BERW formula with the UV-defect

- In a harmonic trap:

 + loop diags :

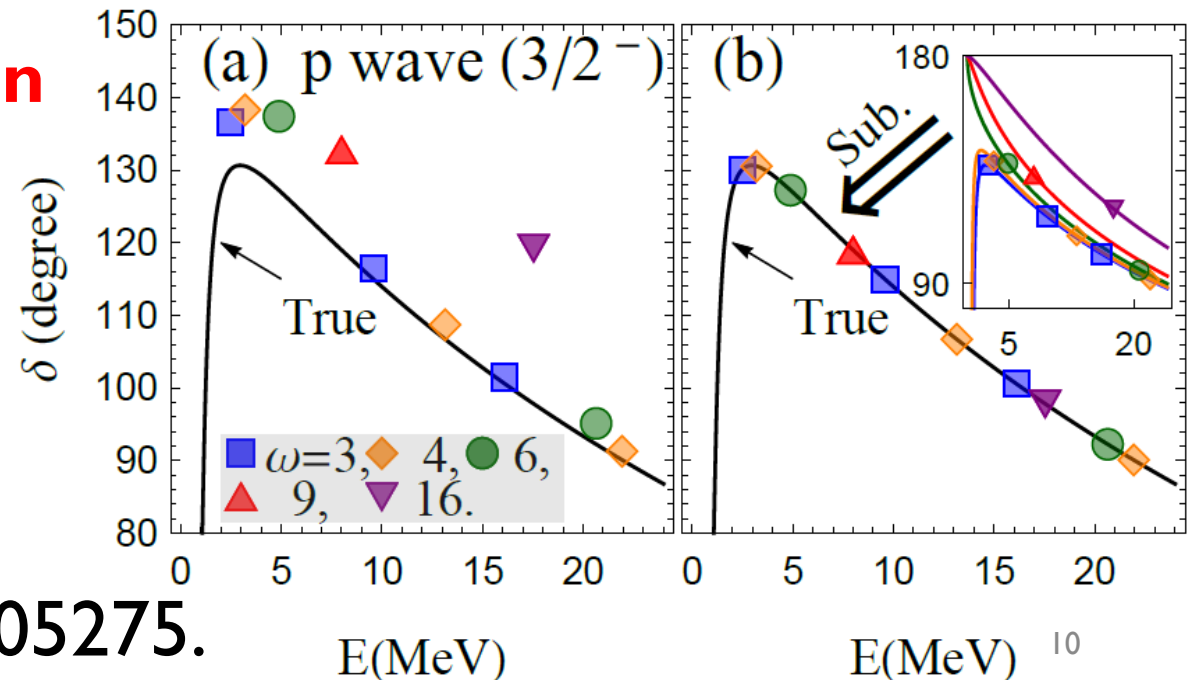
$$U(E) = ip^{2l+1}$$

$$U(E) = (-1)^{l+1} (4 M_R \omega_T)^{l+\frac{1}{2}} \frac{\Gamma\left(\frac{3}{4} + \frac{l}{2} - \frac{E}{2\omega_T}\right)}{\Gamma\left(\frac{1}{4} - \frac{l}{2} - \frac{E}{2\omega_T}\right)}$$

- Generalized ERE and quantization conditions**

$$\sum_{i,j=0}^{+\infty} C_{i,j} (M_R \omega_T)^{2i} p^{2j} = U(E)$$

$$p^{2l+1} \cot \delta_l = \sum_{j=0}^{+\infty} C_{i=0,j} p^{2j}$$



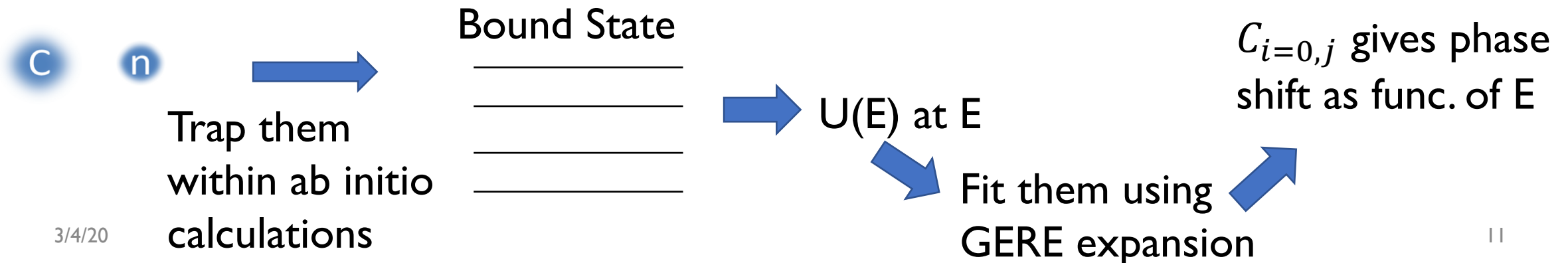
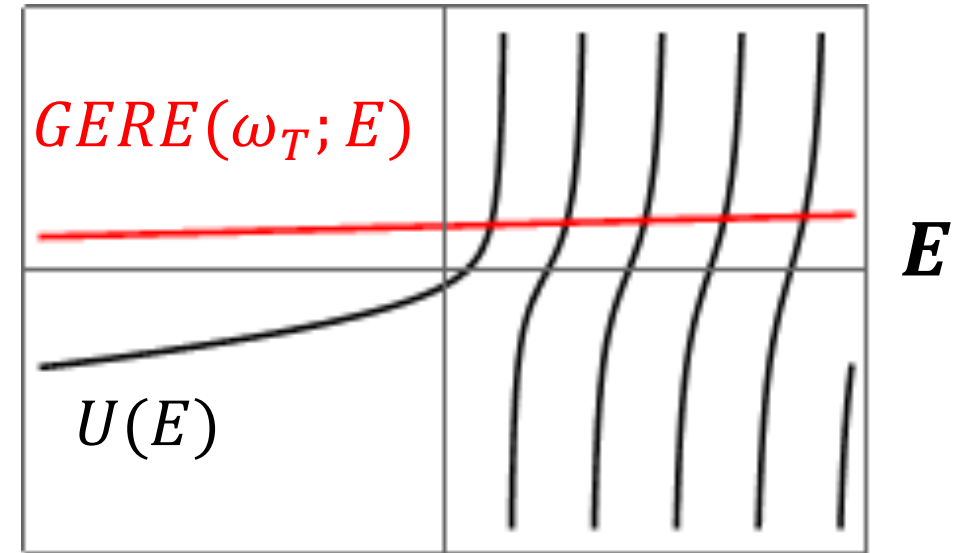
# The perfect computer expt. (no errors!)

Quantization conditions:

$$\sum_{i,j=0}^{+\infty} C_{i,j} (M_R \omega_T)^{2i} p^{2j} = U(E)$$

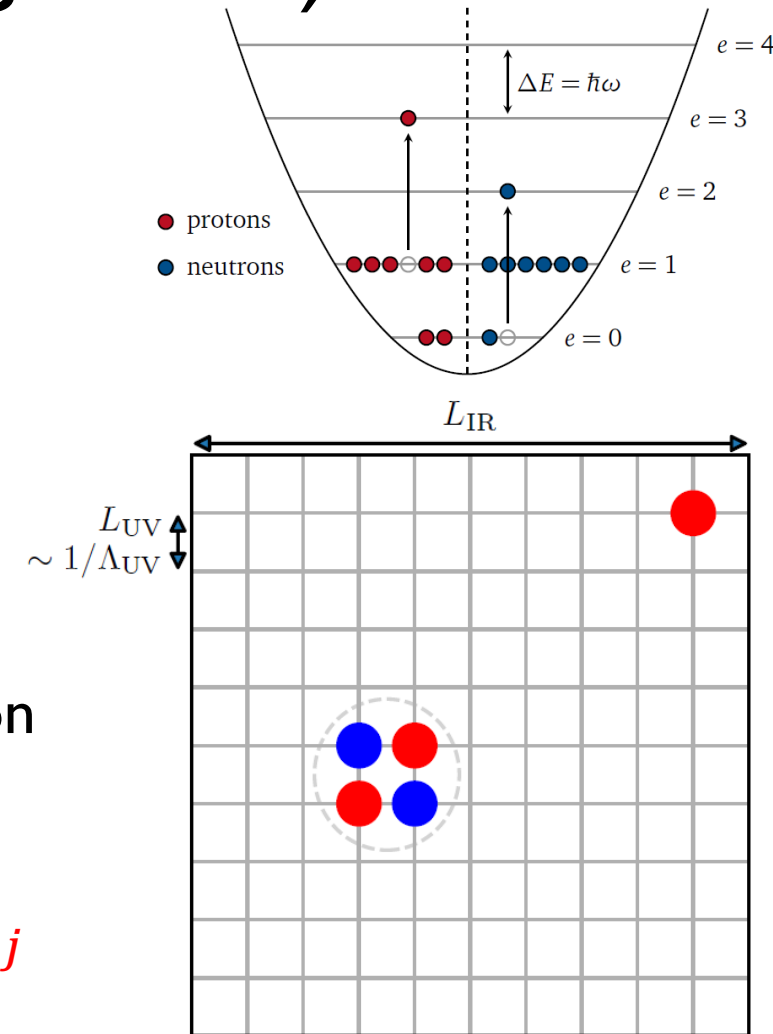
$$U(E) = (-1)^{l+1} (4 M_R \omega_T)^{l+\frac{1}{2}} \frac{\Gamma\left(\frac{3}{4} + \frac{l}{2} - \frac{E}{2\omega_T}\right)}{\Gamma\left(\frac{1}{4} - \frac{l}{2} - \frac{E}{2\omega_T}\right)}$$

$$p^{2l+1} \cot \delta_l = \sum_{j=0}^{+\infty} C_{i=0,j} p^{2j}$$



# Imperfect expts: ab-initio calculations have truncations on Hilbert-space (regulator)

- Working with ab initio groups using harmonic-oscillator-wave-function as basis: NCSM and IMSRG
  - NCSM requires the total energy below  $N_{\max} \omega$  ( $\omega$  as basis frequency)
  - IMSRG requires single particle energy below  $e_{\max} \omega$
- Regulators modify both IR and UV physics  $\rightarrow$  systematic errors
- To model the IR error?  $\rightarrow$  modify U function with truncation on relative motion
- To model the UV impact  $\rightarrow$  the extracted GERE parameters  $C_{i,j}$  becomes  $C_{i,j}(\Lambda_{uv})$ .  $C_{i,j}(\infty)$  (**Cont. limit**), predicts **reality**



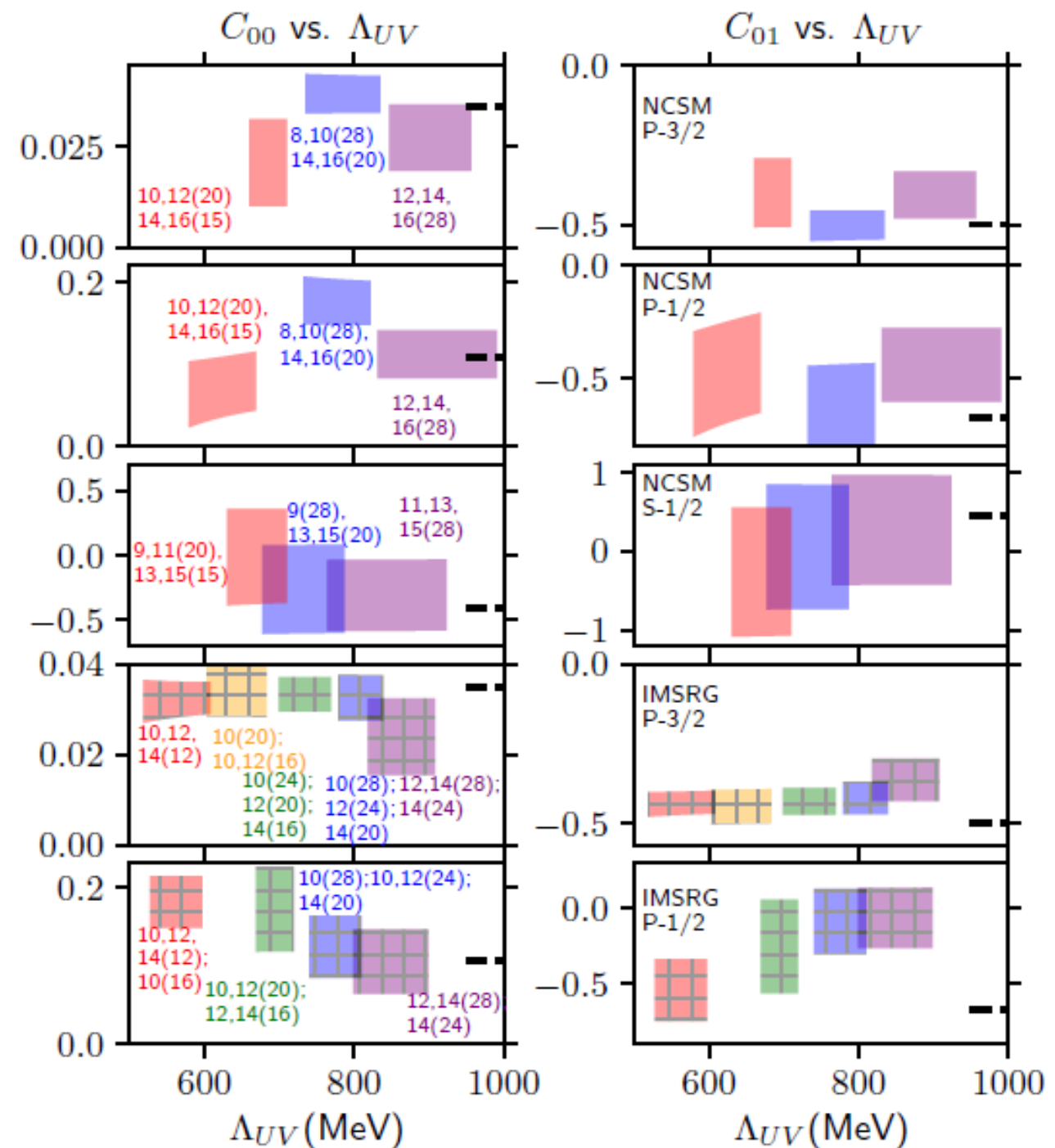
# Before error analysis, let's first see results

Two sets of results based on NCSM and IMSRG output

# n- $\alpha$ scatterings in s and p waves

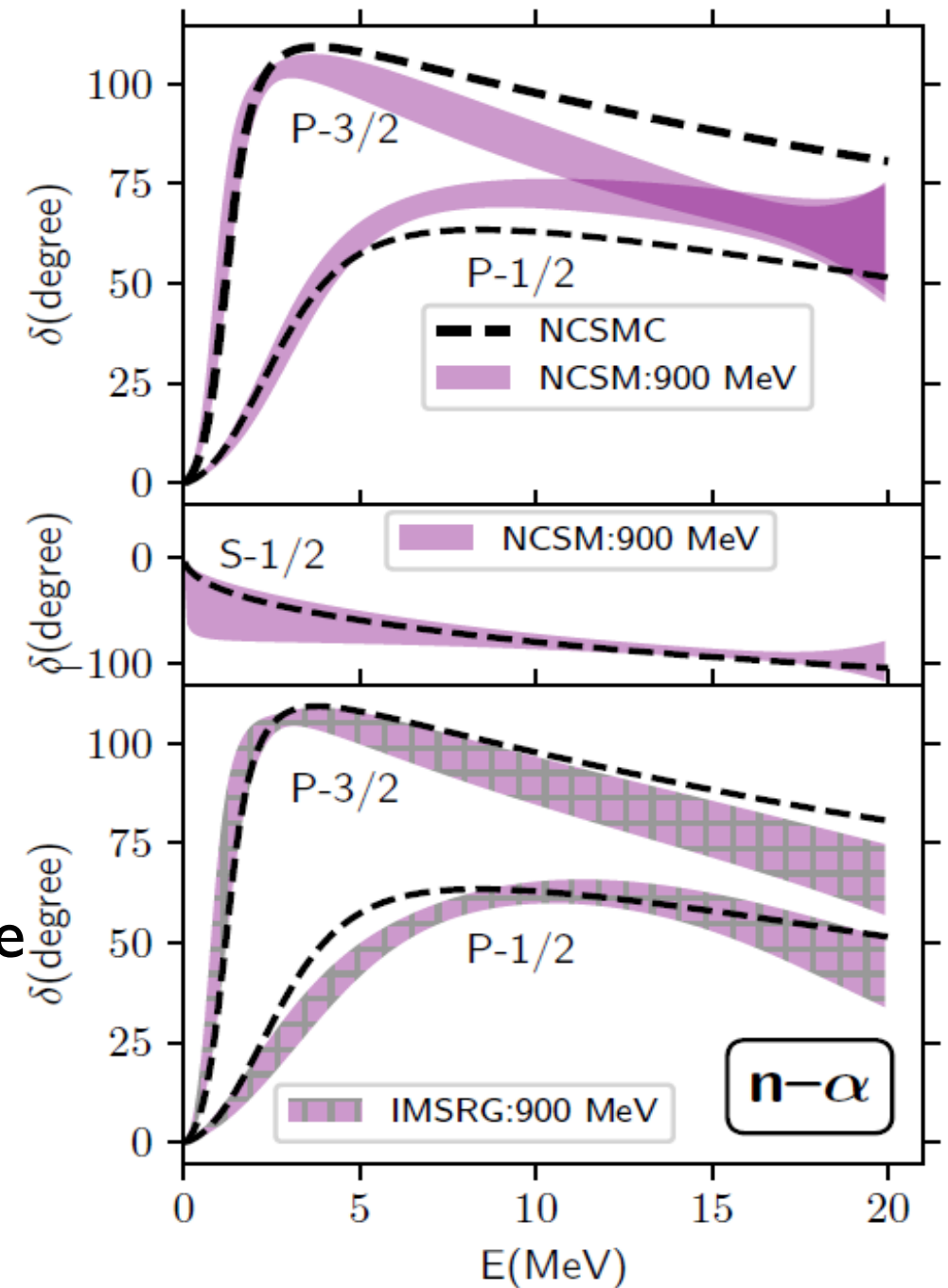
- $C_{i,j}(\Lambda_{uv})$  are the GERE parameters (dimensionless).
- The error band partially comes from IR-error, while the UV-error (not in the band) approaches zero with large  $\Lambda_{uv}$
- Different data sets (using different  $N_{\max}$  and  $\omega$ ) are grouped in different  $\Lambda_{UV}$  bins.
- The parameters are extracted independently among these bins.
- Smooth  $\Lambda_{UV}$ -dependence, a signal that the IR physics is under control

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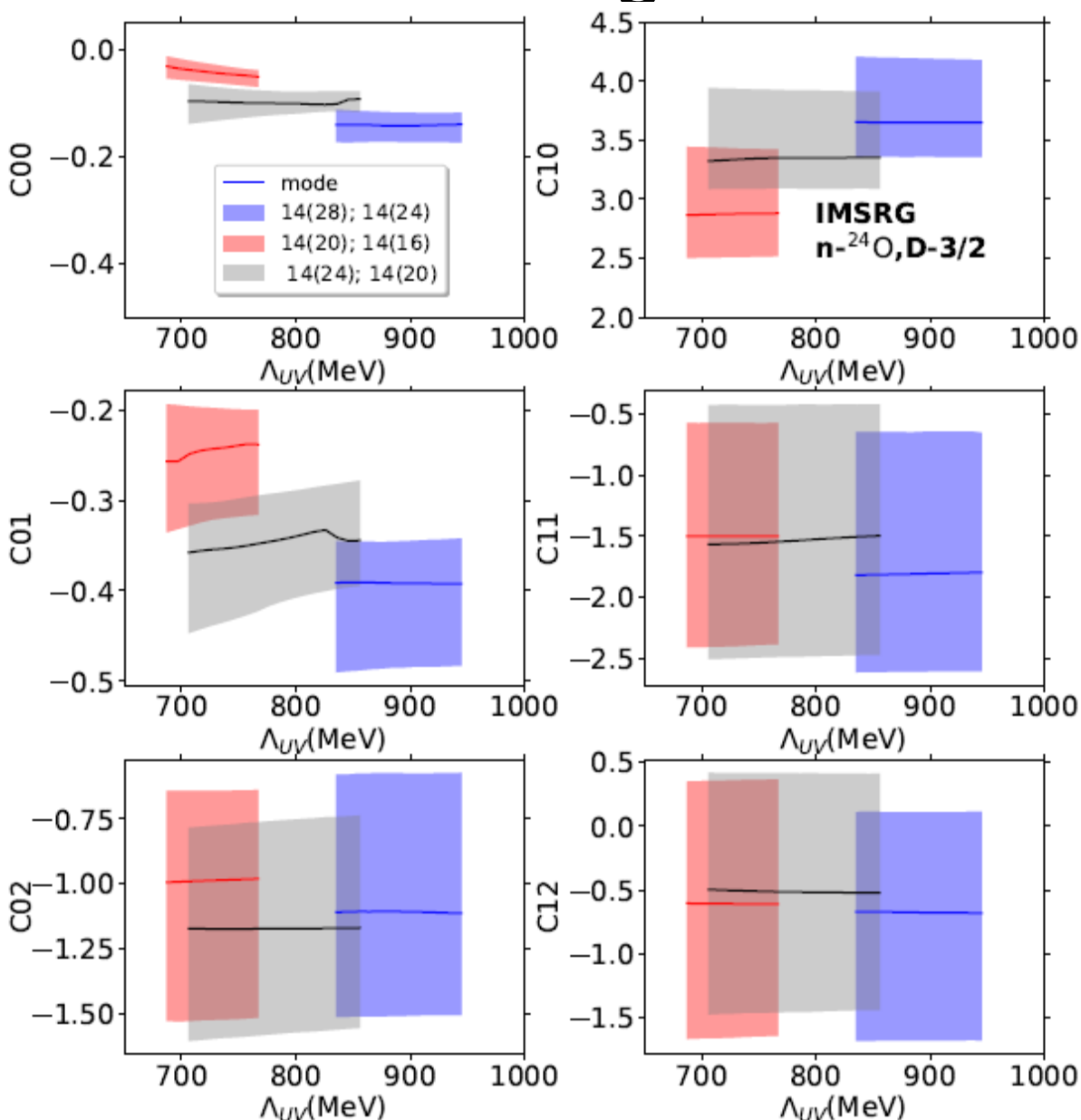


# n- $\alpha$ scatterings in s and p waves

- The NCSM extraction agrees with Petr's direct phase-shift calculation below 5 MeV with **high**  $\Lambda_{UV}$
- The IMSRG p-3/2 agrees with Petr's direct calculation, but p-1/2 is somewhat different
- Since we model both UV and IR physics components, we can use most (Nmax/emax,  $\omega$ ) results and extract phase-shifts. This is like LQCD producing results at different Lattice spacing.



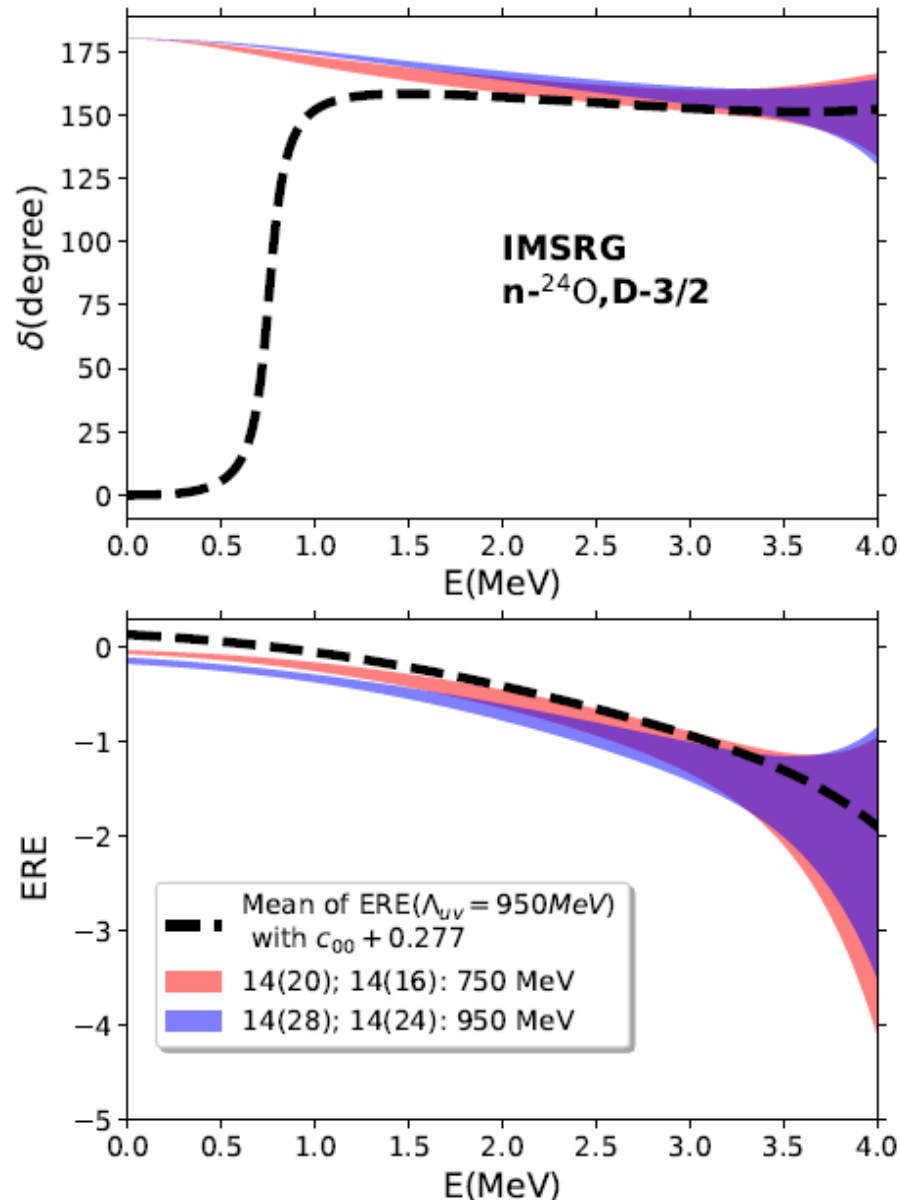
# Analyze O24 and O25 energies from IMSRG and extract n-O24 scattering in d-3/2 channel



- C00 decreases with increasing  $\Lambda_{UV}$   
(smooth evolution: C00~0.1, NOT 10 or 0.001)
- C01 and C10 have also some UV dependence
- Approaching to the continuum limit, no resonance behavior in scattering
- If C00 were positive, the resonance shows up (NOT the prediction of the used nucleon interaction)



# Analyze O24 and O25 energies from IMSRG and extract n-O24 scattering in d-3/2 channel

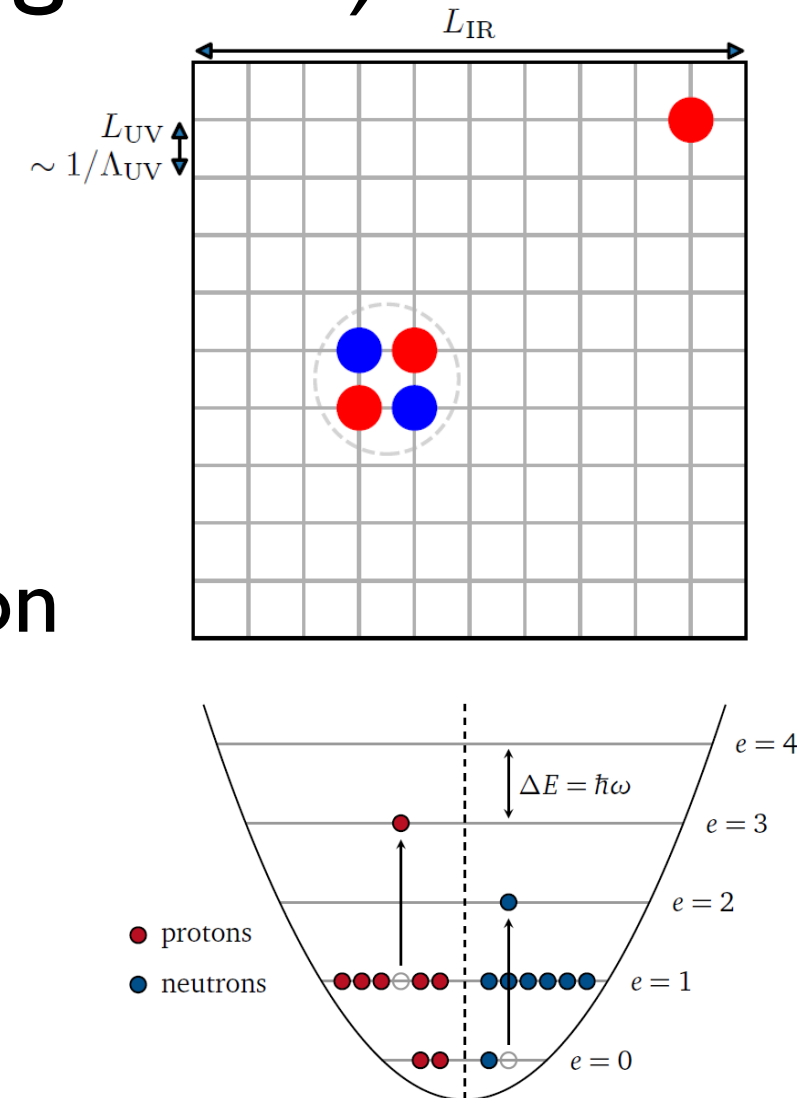


- Studying ERE at  **$E < 0$  region** predicts a low-energy bound state with 75% prob., the BE is  $-1.4 \pm 0.54 \text{ MeV}$  (**the property of the NN force**)
- The method should also be used to compute BE of shallow bound state (drip-line nuclei),
- Use the mean value of Cs(900 MeV) and increase  $C_{00}$  by about 0.28  $\rightarrow$  a resonance (dashed curve) at 0.75 MeV and width of 135 keV (similar to the expt. info): **tuning** nucleon interaction could improve the prediction

# Here comes the messy part

# Imperfect expts: ab-initio calculations have truncations on Hilbert-space (regulator)

- They modify both IR and UV physics  $\rightarrow$  systematic errors (tiny stochastic errors)
- To model the IR error?  $\rightarrow$  modify U function with truncation on relative motion
- To model the UV impact  $\rightarrow$  the extracted GERE parameters  $C_{i,j}$  becomes  $C_{i,j}(\Lambda_{uv})$



# Model the IR-error

With  $n_\Lambda$  truncating number of excitation quanta in terms of  $\omega$ :

$$U = (-1)^{l+1} \left( \frac{2}{xb_T} \right)^{2l+1} \frac{\Gamma(l + 3/2)}{\Gamma(\frac{1}{2} - l)} \frac{\Gamma(n_\Lambda + 2)}{\Gamma(n_\Lambda + l + 5/2)} \frac{{}_2F_1\left(n_E + 1, -n_\Lambda - l - \frac{3}{2}, -l + \frac{1}{2}, x^2\right)}{{}_2F_1\left(n_E + l + \frac{3}{2}, -n_\Lambda - 1, l + \frac{3}{2}, x^2\right)}$$

- $n_E = \frac{E}{2\omega_T} - \frac{l}{2} - \frac{3}{4}; b_T = \frac{1}{\sqrt{m_R\omega_T}}, b = \frac{1}{\sqrt{m_R\omega}}, x = \frac{2b_Tb}{b_T^2 + b^2}$
- **The U depends on truncation parameters. It models the IR-error**
- **It approaches the U(E) in the continuum limit (zero IR-error) with  $\omega \rightarrow \omega_T$  or  $n_\Lambda \rightarrow \infty$**

# Access the UV-error (besides series truncation error in GERE)

Use the following as interpolator over a small  $\Lambda_{uv}$  bin (less than 100 MeV).  $C_{i,j,0}$   $C_{i,j,1}$  and  $Q$  are fitting parameters

$$C_{i,j}(\Lambda_{uv}) = C_{i,j,0} + C_{i,j,1} \left( \frac{Q}{\Lambda_{uv}} \right)^2$$

## Another error

For example, the n-alpha data:

He-5 (He-4) mass tabulated in terms of Nmax5 (Nmax4),  $\omega$ , and  $\omega_T$

$$E(N_{\max 5}, \omega, \omega_T) = E_{He5}(N_{\max 5}, \omega, \omega_T) - E_{He4}(N_{\max 4}, \omega, \omega_T)$$

For IR-error:  $n_\Lambda = n_\Lambda(N_{max,5})$

# A digression to Bayesian inference

$$\text{pr}(\mathbf{C}, \mathbf{d} | D, T, I) \propto \text{pr}(D | \mathbf{C}, \mathbf{d}, T, I) \text{pr}(\mathbf{C}, \mathbf{d} | I)$$

↑  
Posterior  
distribution

↑  
Likelihood function

↑  
Prior distribution

**C** are the GERE parameters; **d** are the other parameters describing the IR and UV error

# A few thoughts on the connection to other approaches and its potential application in quantum computing

	Generalized Luscher	Luscher in Lattice calculations	Single-State HORSE (Harmonic Oscillator Representation of Scattering Eqs)
IR regulator	Trap	Large volume	Effective cavity in the Nmax truncations
UV regulator	Nominal $\Lambda_{UV}$	Lattice spacing	$\Lambda_{UV}$ ?
IR-UV explicitly coupled?	Yes, so we need a model to extract them	Decoupled	Yes
IR-error	Use effective Nmax or emax for relative motion to access this error	Imbedded in MC sampling: the calculation can not handle arbitrarily large volume, i.e., there is error related to IR physics in sampling. (what is the signal for that error?)	?
Error types	Mostly systematic, except the rounding errors	Stochastic error in sampling	?

**In our approach, if the base frequency approaches trap frequency, the IR-error would disappear. While doing so, to reduce the UV error, Nmax/emax has to be increased significantly. So perhaps quantum computer is one way to go.**

# Summary

- The Luscher method used in LQCD is generalized to work with ab initio nuclear structure methods (NCSM and IMSRG)
- The extracted  $n$ - $\alpha$  phase-shifts are compatible with the existing ab initio results using the same nucleon interaction
- The  $n$ - $^{24}\text{O}$  scattering calculation demonstrates the method's capability to study systems heavier (it is based on bound-state spectrum calculation)
- The UV and IR physics are correctly modeled  $\rightarrow$  results from different regulators are consistent (smooth UV-scale dependence)
- We now report results as function of UV scale, a practice common in Lattice methods (Lattice QCD, NLEFT), but not in other ab initio calculations



# U(E) in the finite model space (H.O. basis, NCSM)

- In a “cavity” for relative motion with radius L

$$U = p^{2l+1} \frac{y_l(Lp)}{j_l(Lp)} \leftarrow \langle i \left| V_s \frac{1}{E-H_0} V_s \right| i \rangle$$

Approaches approx. to a **fuzzy** cavity limit, with

$$L = \sqrt{2(2n_\Lambda + l + \frac{3}{2} + 2)} b, \text{ in a non-trivial}$$

$$\text{way } (\Lambda = \sqrt{2(2n_\Lambda + l + \frac{3}{2} + 2)} / b)$$

- In Nmax truncation scheme for two-body relative motion (“**U for the un-trapped and the truncated**”):

$$U = \frac{(-1)^{l+1}}{b^{2l+1}} \frac{\Gamma(l+3/2)}{\Gamma(\frac{1}{2}-l)} \frac{\Gamma(n_\Lambda+2)}{\Gamma(n_\Lambda+l+5/2)} \frac{M(-n_\Lambda-l-\frac{3}{2}, -l+\frac{1}{2}, p^2 b^2)}{M(-n_\Lambda-1, l+\frac{3}{2}, p^2 b^2)} \leftarrow \langle i \left| V_s \hat{P} \frac{1}{E-H_0} \hat{P} V_s \right| i \rangle$$

$$H_0 \rightarrow H_{\omega_T}$$

$$\bullet \hat{P} = \hat{P}(N_\Lambda, \omega) = \sum_{n=0}^{n_\Lambda} |HO_n\rangle \langle HO_n| = \sum_{\mu=0}^{n_\Lambda} |DVR_\mu\rangle \langle DVR_\mu|; N_{max} = 2n_\Lambda + l$$

$$\bullet b = 1/\sqrt{m_R \omega} \text{ (}\omega \text{ is base freq.)}$$

Approaches to above when  $\omega_T \rightarrow 0$ , also to the continuum limit properly when  $\omega \rightarrow \omega_T$  or  $n_\Lambda \rightarrow \infty$

- Now with the trapping potential, with  $\omega_T$  frequency (“**U for the trapped and truncated**”)

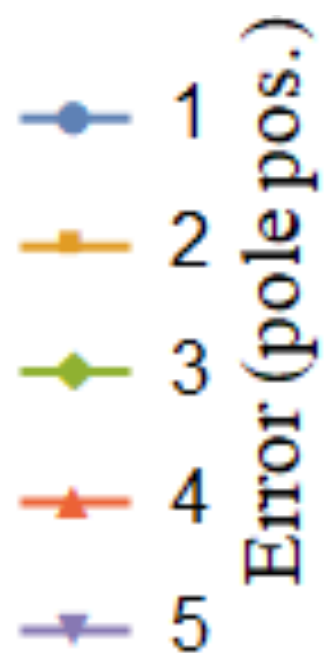
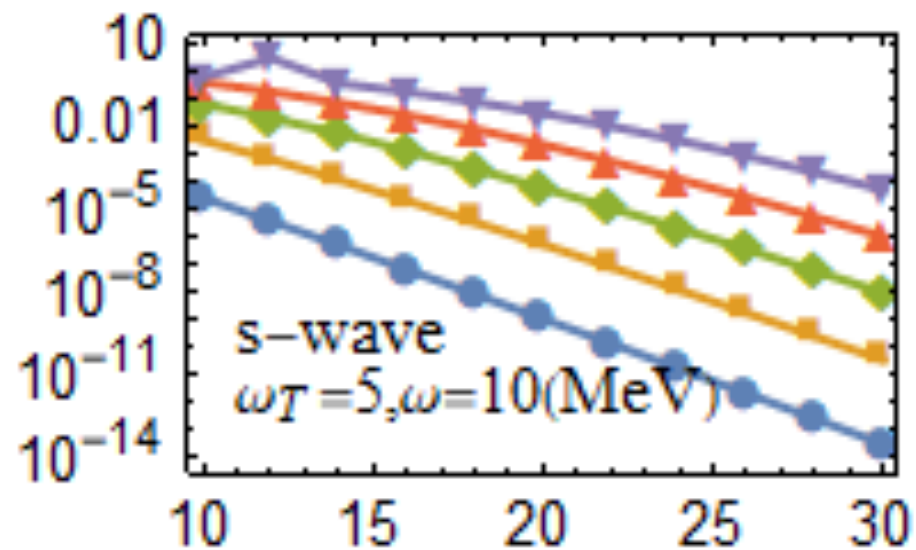
$$U = (-1)^{l+1} \left( \frac{2}{x b_T} \right)^{2l+1} \frac{\Gamma(l+3/2)}{\Gamma(\frac{1}{2}-l)} \frac{\Gamma(n_\Lambda+2)}{\Gamma(n_\Lambda+l+5/2)} \frac{{}_2F_1\left(n_E+1, -n_\Lambda-l-\frac{3}{2}, -l+\frac{1}{2}, x^2\right)}{{}_2F_1\left(n_E+l+\frac{3}{2}, -n_\Lambda-1, l+\frac{3}{2}, x^2\right)}$$

$$\bullet n_E = \frac{E}{2\omega_T} - \frac{l}{2} - \frac{3}{4}; b_T = \frac{1}{\sqrt{m_R \omega_T}}, x = \frac{2 b_T b}{b_T^2 + b^2}$$

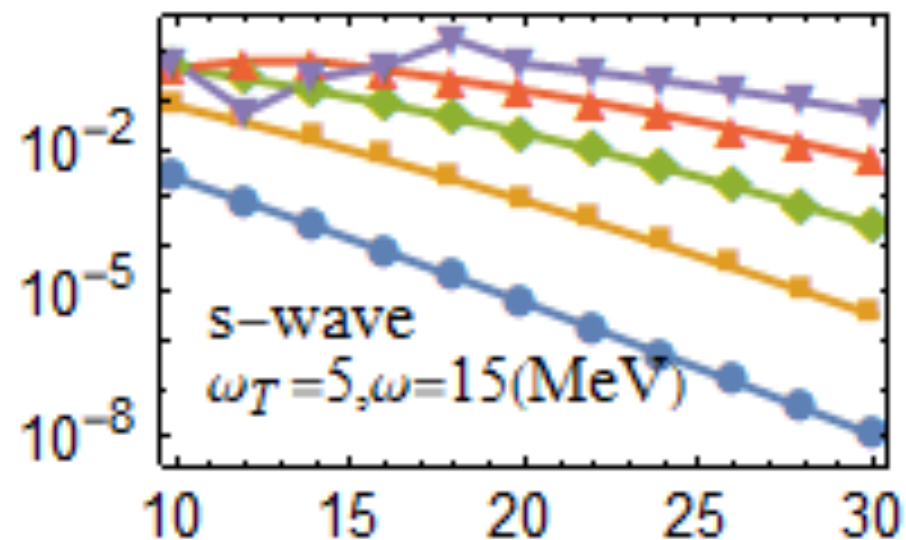
- the low-energy pole position of this function is very interesting:

$$n + O\left((n_\Lambda + 1)^{d+\frac{1}{2}} e^{-\frac{(n_\Lambda+1)}{2}\eta^2}\right), \text{ with } d < 0, \eta \sim -\sqrt{-2 \text{Log}(1-x^2)}$$

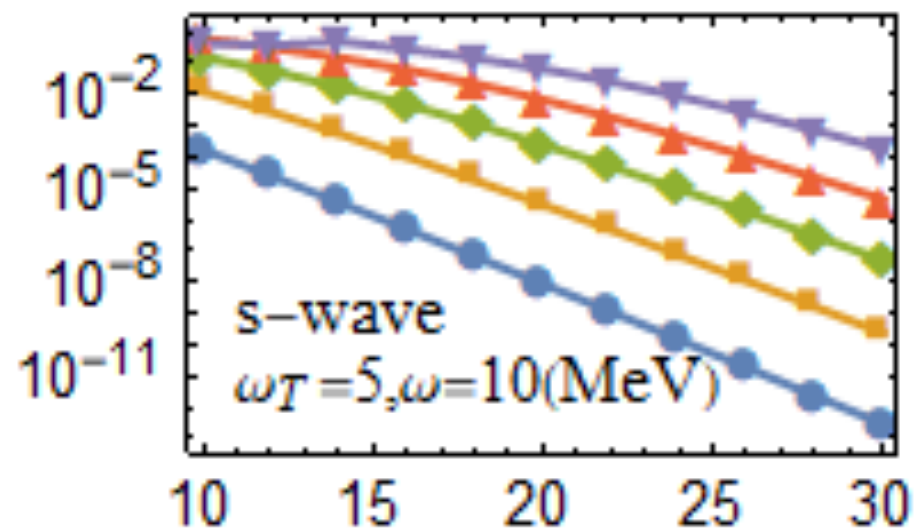
Error (pole pos.)



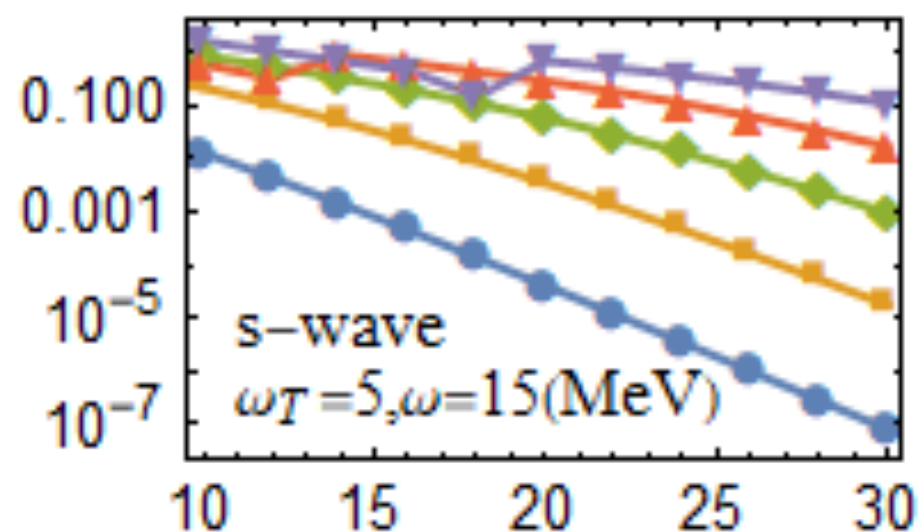
Error (pole pos.)



Per. error (res)



Per. error (res)



$N_{\max}$

The pole structures of the  $U$  function converge to the continuum limit in terms  $\text{Exp}(-C \cdot N_{\max})$

$N_{\max}$

# Outlook

- Need to understand the origin of the parameters modeling errors in our “data” analysis in terms of microscopic picture
- Explore other observables
- To compute **reactions**, need to study coupled-channel problem within traps. EFT provide one way to approach this problem
- From the angle of computer experiment, other existing models used in nuclear experiments (**potential model and R-matrix**), could also be used in data analysis. The important steps involve understanding the theory error of the model and the computer-experiment’s error (UV and IR modification within the ab initio calculation)

# In retrospect

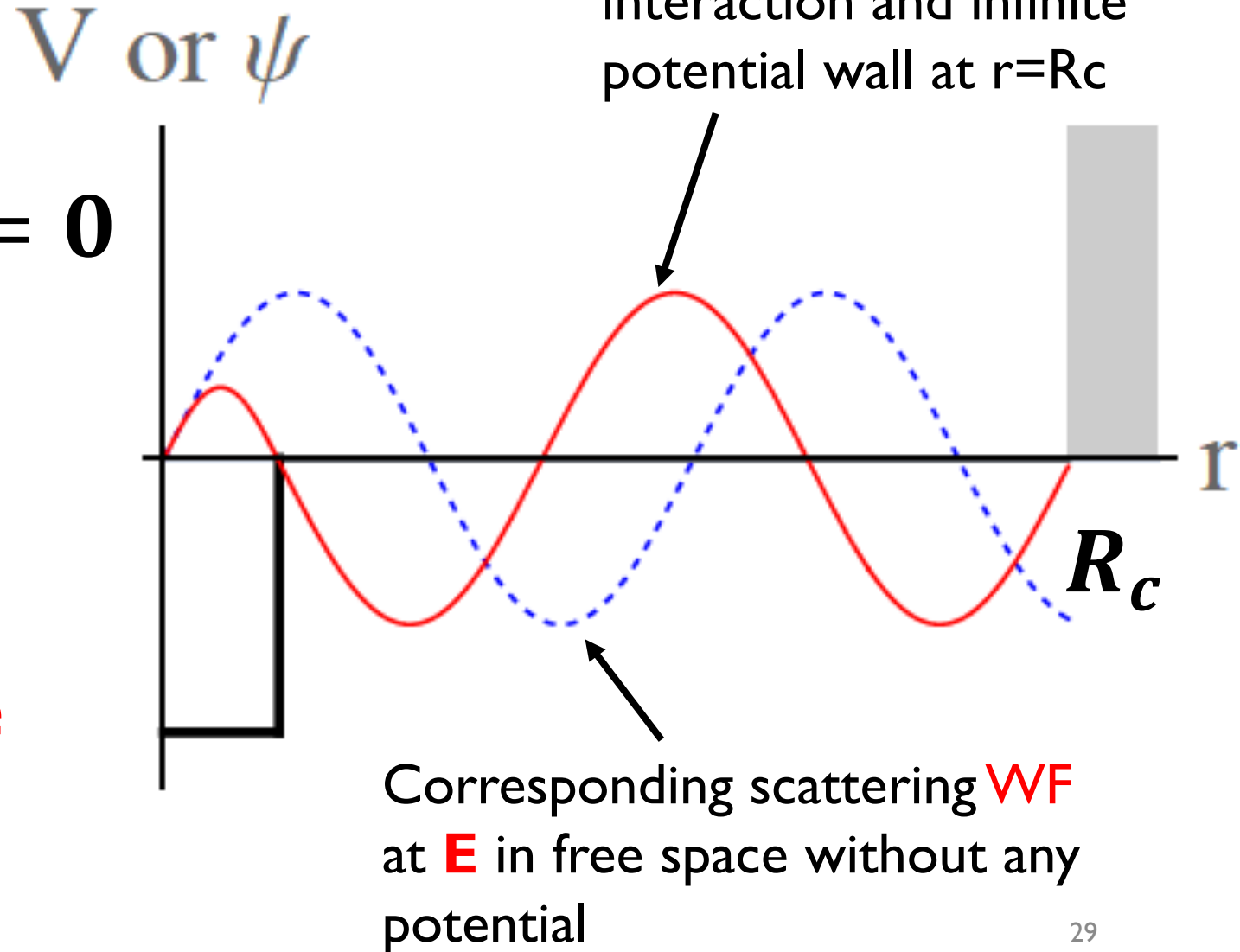
- In the harmonic-oscillator-basis structure calculations, the **clustering physics already exists**, i.e., the basis can handle continuum physics **with the help of a trap**
- The key is to understand properly the IR physics/condition in terms of clusters.
- Our IR is dictated by the trap. Truncation leads to IR error and approaches to zero in the continuum limit

Cavity boundary condition (for relative motion)

$$\tan \delta_l \times n_l(kR_c) + j_l(kR_c) = 0$$

$$\text{i.e., } (E, R_c) \rightarrow \delta_l(E)$$

Discrete **eigen-energies** for system in a **cavity** gives the **phase shift at those energies**

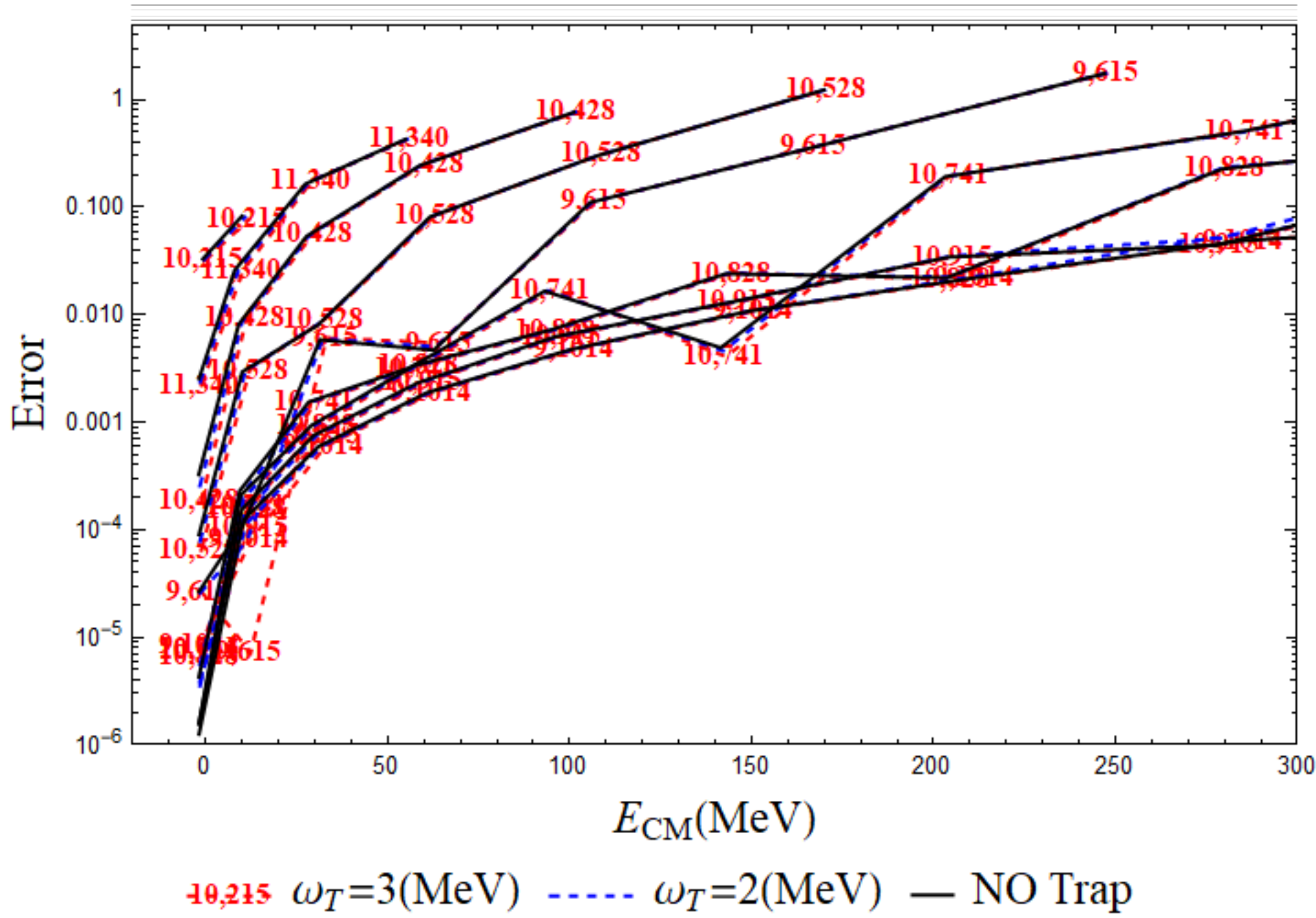


However trapping nucleons in harmonic potential well is better suited for harmonic-oscillator-basis calculations

- Reduces degrees of freedom (DOF)  $\rightarrow$  make ab initio calculations feasible
- The center of mass (CM) and internal DOF are **decoupled**
- Cavity boundary condition requires high energy modes (CM-internal-decoupling is violated in cavity)
- Lattice regulator breaks rotational invariance

There is a “universal” formula for two-cluster system at low energy  $\rightarrow$  BERW (Busch) formula  $(E, \omega_T) \rightarrow \delta_l(E)$

- NN model:(G)ERE extractions for the s-wave, for both trapped and non-trapped system.
- The markers are **L** and  **$\Lambda$**  as defined before
- UV error scales as  $1/L\Lambda^3$  at leading order
- Different IR modification is disentangled from UV error
- The UV error can be reduced by improving potential's UV behavior
- This also suggests that if NN is optimized, adding further trap doesn't require further optimization





# U(E) in the finite model space (H.O. basis, NCSM)

- In a “cavity” for relative motion with radius L

$$U = p^{2l+1} \frac{y_l(Lp)}{j_l(Lp)} \leftarrow \langle i \left| V_s \frac{1}{E-H_0} V_s \right| i \rangle$$

Approaches approx. to a **fuzzy** cavity limit, with

$$L = \sqrt{2(2n_\Lambda + l + \frac{3}{2} + 2)} b, \text{ in a non-trivial}$$

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- In Nmax truncation scheme for two-body relative motion (“**U for the un-trapped and the truncated**”):

$$U = \frac{(-1)^{l+1}}{b^{2l+1}} \frac{\Gamma(l+3/2)}{\Gamma(\frac{1}{2}-l)} \frac{\Gamma(n_\Lambda+2)}{\Gamma(n_\Lambda+l+5/2)} \frac{M(-n_\Lambda-l-\frac{3}{2}, -l+\frac{1}{2}, p^2 b^2)}{M(-n_\Lambda-1, l+\frac{3}{2}, p^2 b^2)} \leftarrow \langle i \left| V_s \hat{P} \frac{1}{E-H_0} \hat{P} V_s \right| i \rangle$$

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- Now with the trapping potential, with  $\omega_T$  frequency (“**U for the trapped and truncated**”)

$$U = (-1)^{l+1} \left( \frac{2}{x b_T} \right)^{2l+1} \frac{\Gamma(l+3/2)}{\Gamma(\frac{1}{2}-l)} \frac{\Gamma(n_\Lambda+2)}{\Gamma(n_\Lambda+l+5/2)} \frac{{}_2F_1\left(n_E+1, -n_\Lambda-l-\frac{3}{2}, -l+\frac{1}{2}, x^2\right)}{{}_2F_1\left(n_E+l+\frac{3}{2}, -n_\Lambda-1, l+\frac{3}{2}, x^2\right)}$$

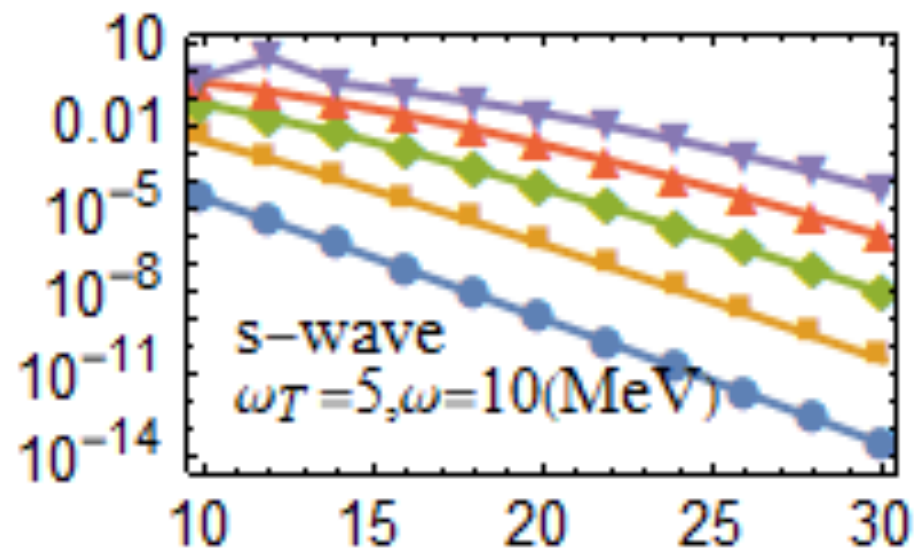
$$\bullet n_E = \frac{E}{2\omega_T} - \frac{l}{2} - \frac{3}{4}; b_T = \frac{1}{\sqrt{m_R \omega_T}}, x = \frac{2 b_T b}{b_T^2 + b^2}$$

- the low-energy pole position of this function is very interesting:

$$n + O\left((n_\Lambda + 1)^{d+\frac{1}{2}} e^{-\frac{(n_\Lambda+1)}{2}\eta^2}\right), \text{ with } d < 0, \eta \sim -\sqrt{-2 \text{Log}(1-x^2)}$$

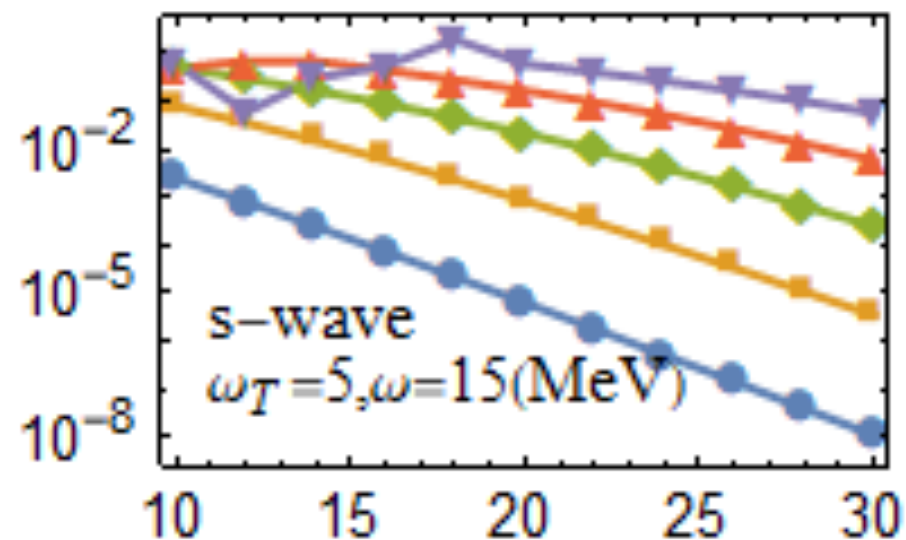


Error (pole pos.)

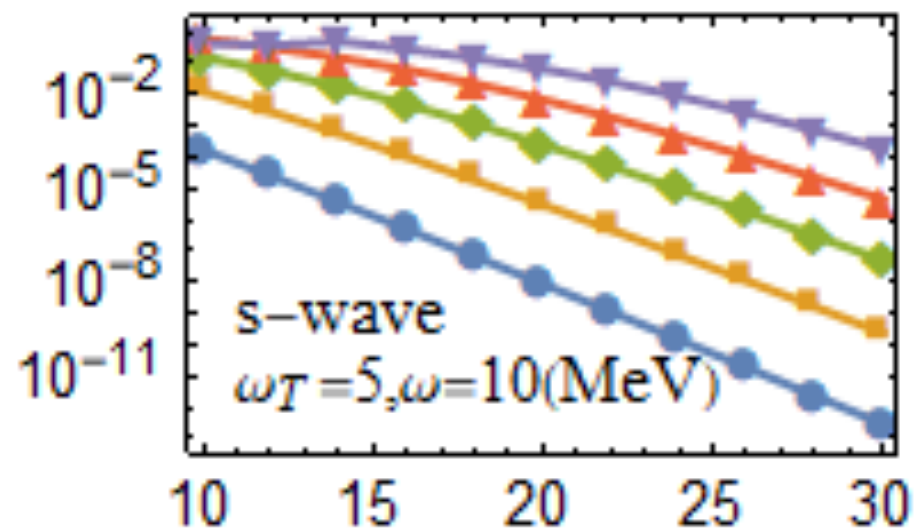


1  
2  
3  
4  
5

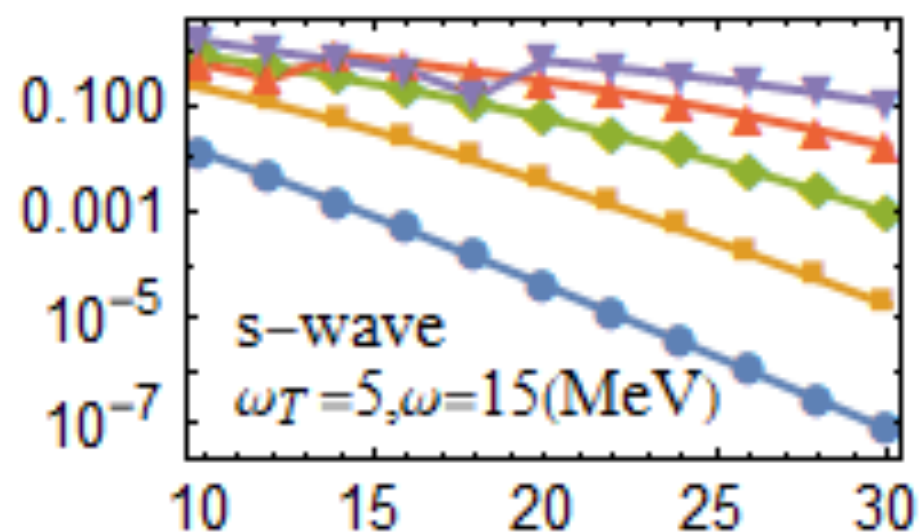
Error (pole pos.)



Per. error (res)



Per. error (res)



Nmax

The pole structures of the U function converge to the continuum limit in terms  $\text{Exp}(-C \cdot N_{\text{max}})$

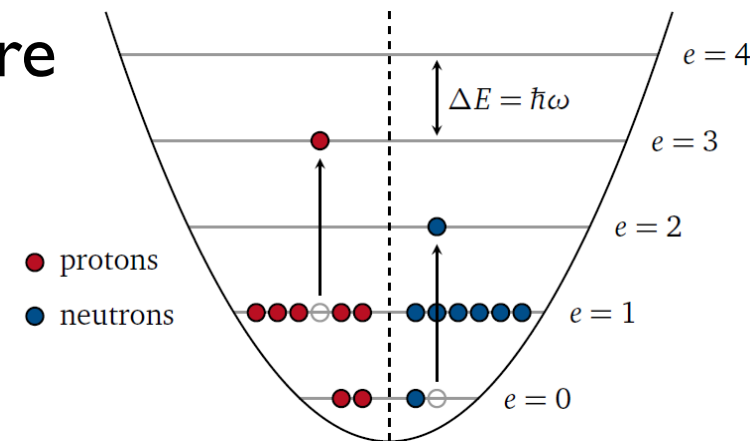
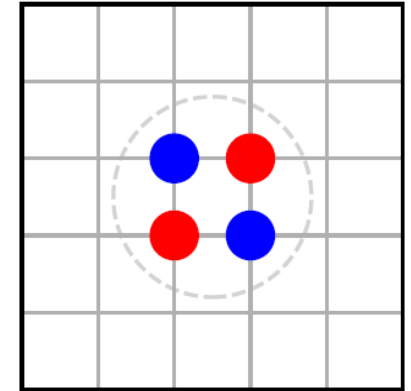
Nmax

# Useful references:

- **Cavity boundary condition:** E.g., R. J. Furnstahl, S. N. More, and T. Papenbrock (2014)
- **Discrete Variable Representation (DVR):** S. Binder, A. Ekstrom, G. Hagen, T. Papenbrock, and K.A. Wendt (2016)
- **Many-body extrapolation:** K.A. Wendt, C. Forssen, T. Papenbrock, and D. Saaf (2015)
- **J-matrix derivation (known as HORSE or SS-HORSE in nuclear physics for extracting phase-shift from energy spectrum):** A. M. Shirokov, A. I. Mazur, I.A. Mazur, and J. P.Vary (2016); A. M. Shirokov, A. I. Mazur, I.A. Mazur, E.A. Mazur, I. J. Shin, Y. Kim, L. D. Blokhintsev, and J. P.Vary (2018)

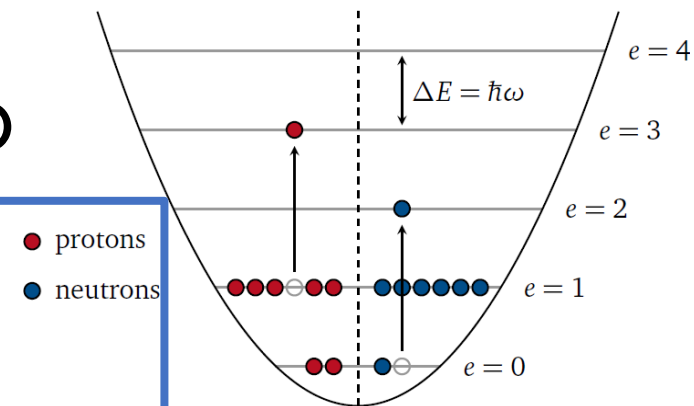
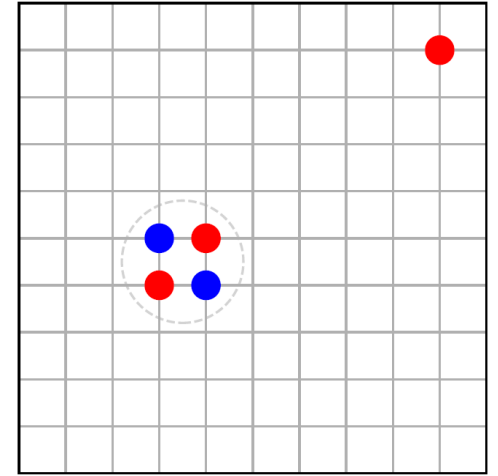
# Ab initio calculations of bound nuclei

- Count degrees of freedom (DOF): nucleons' space locations (& internal DOF). Treating differently:
  - Green's function MC, nuclear lattice effective field theory (NLEFT), as well as lattice QCD (LQCD)
  - Basis method, e.g. Hamiltonian diagonalization in no-core shell model (**NCSM**) and in-medium similarity renormalization group (**IMSRG**), and coupled-cluster

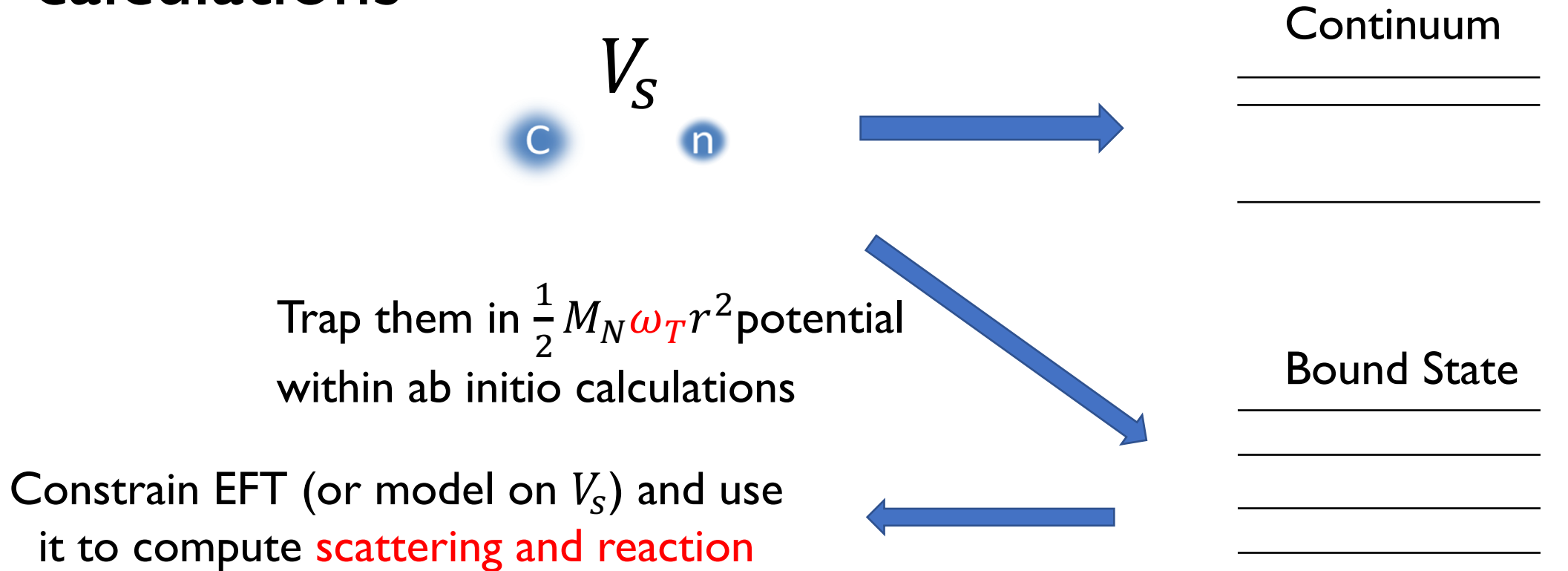


# Ab initio calculations of scattering/reactions

- For scattering/reactions, all the nucleons must be treated in the same way  $\rightarrow$  too many DOF. Then options are
  - NCSM+continuum
  - Gamow shell-model
  - Ab initio optical potential
  - Compute energies by MC-sampling of important configurations **at finite volume**: NLEFT, GFMC, LQCD
- **Finite volume reduces DOF. So is trapping!**
- **And eigen-energies give phase shift.**
- Regulators also in other calculations (e.g., Geant simulation)



# Trapping nucleons in ab-initio spectrum calculations

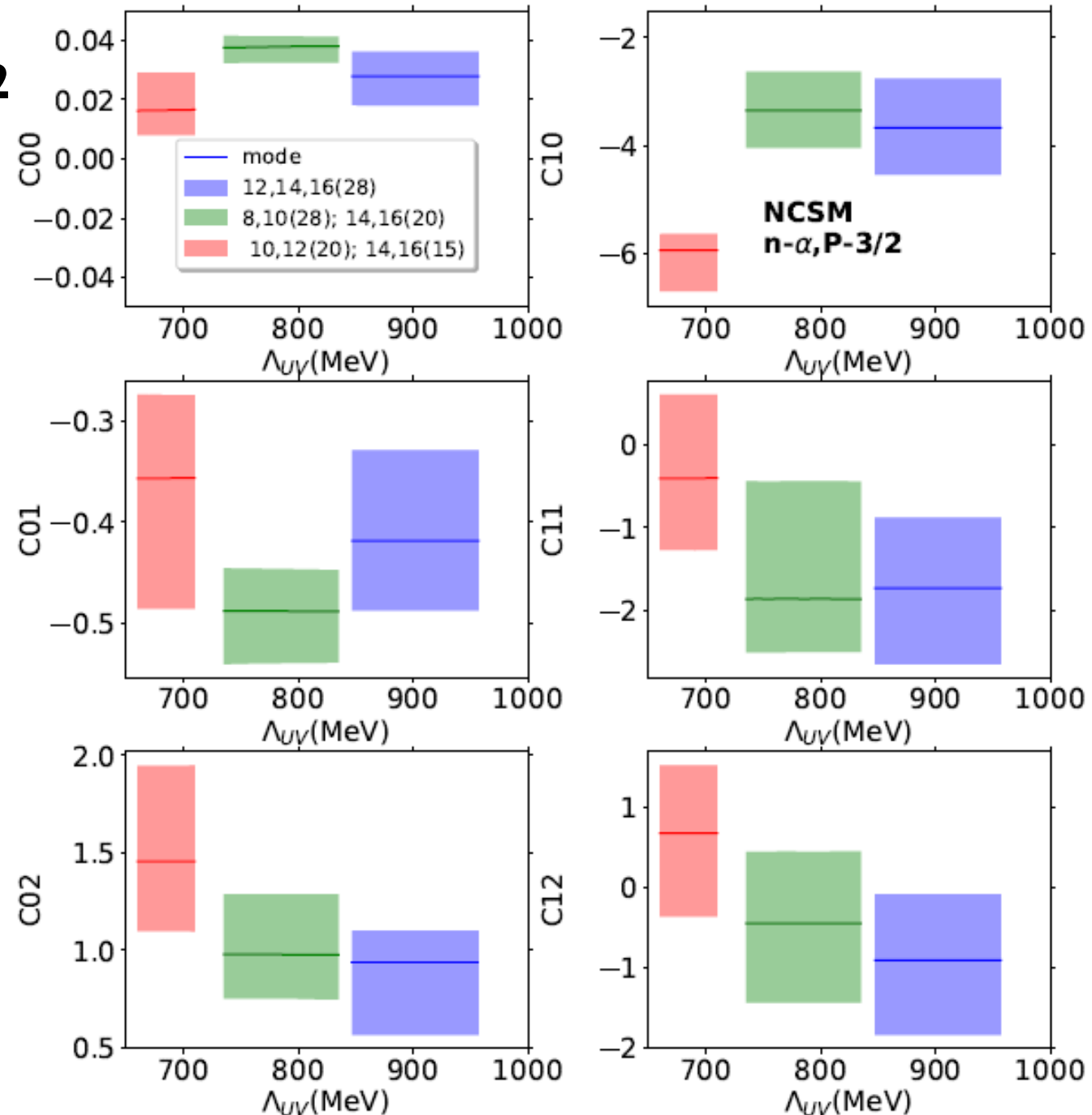


**There is a “universal” formula for two-cluster system at low energy  $\rightarrow$  BERW (Busch) formula**

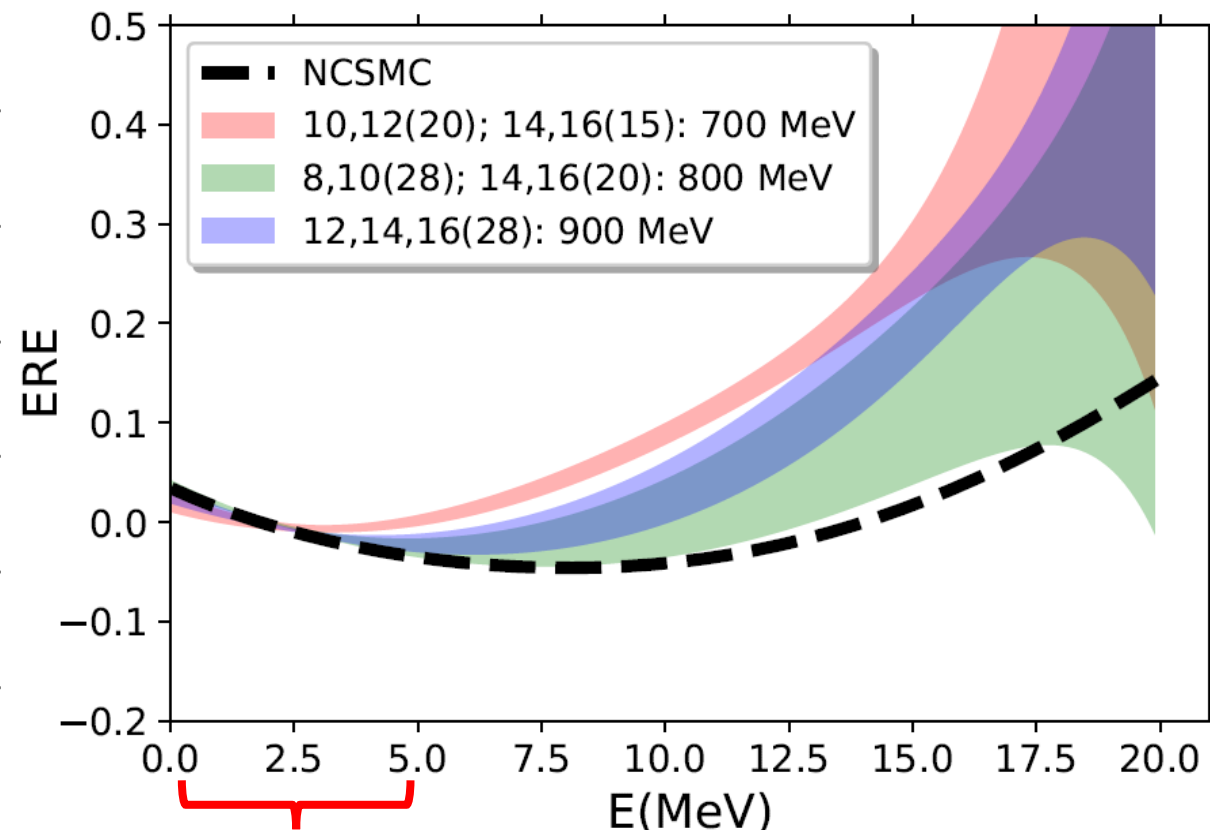
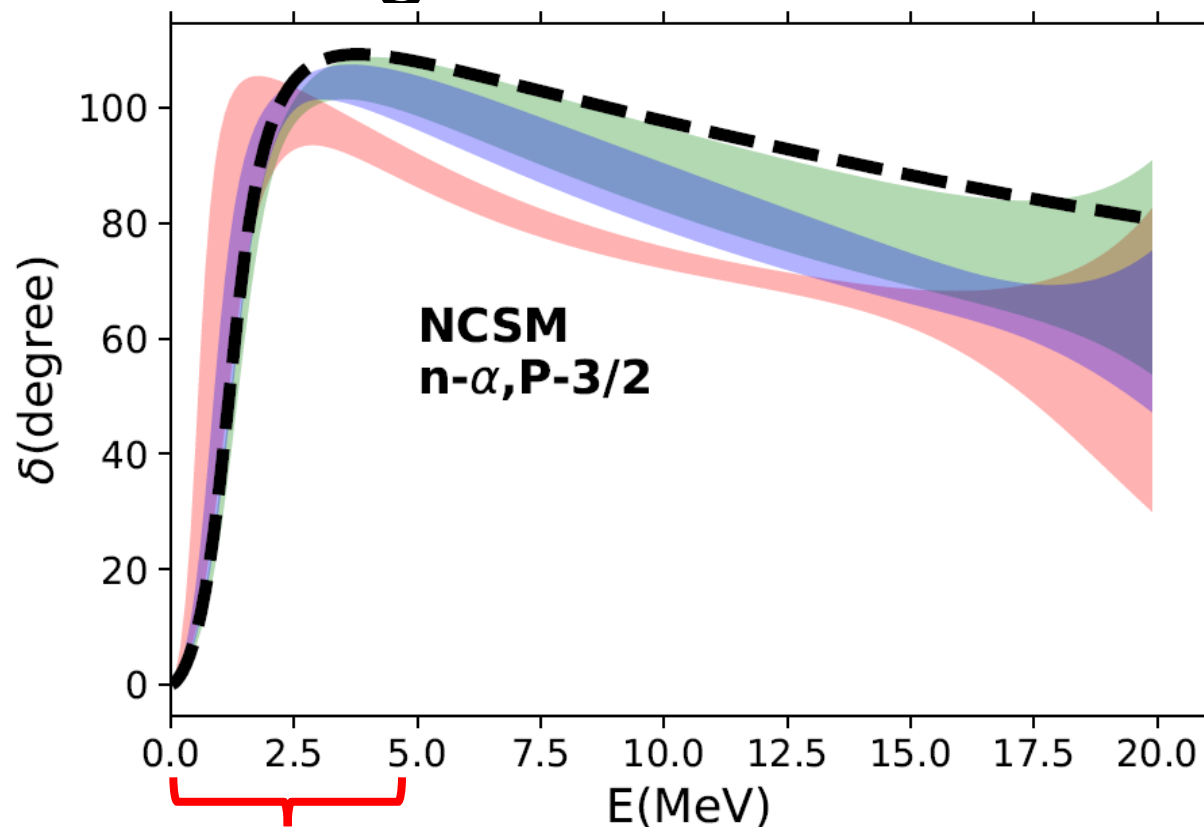
$$(E, \omega_T) \rightarrow \delta_l(E)$$

# Analyze He4 and He5 energies from NCSM and extract n- $\alpha$ scattering in P-3/2 channel

- $C_{i,j}(\Lambda_{uv})$  are the GERE parameters (dimensionless).
- **The error band partially comes from IR-error, while the UV-error (not in the band) approaches zero with large  $\Lambda_{uv}$**
- Different data sets (using different  $N_{\max}$  and  $\omega$ ) are grouped in different  $\Lambda_{UV}$  bins.
- The parameters are extracted independently among these bins.
- Smooth  $\Lambda_{UV}$ -dependence, a signal that the IR physics is under control

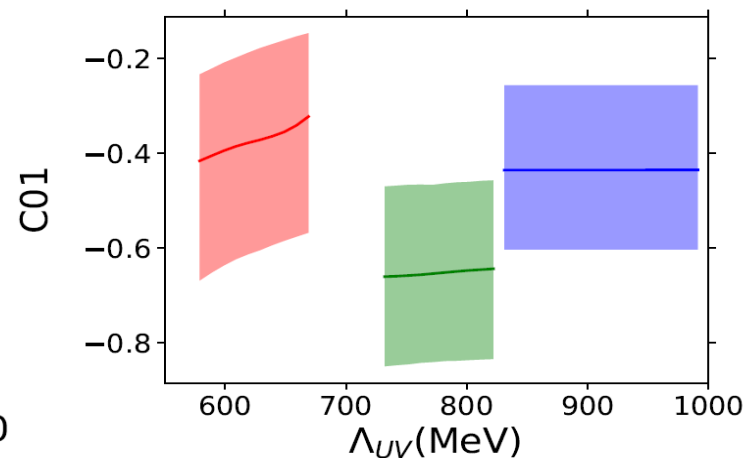
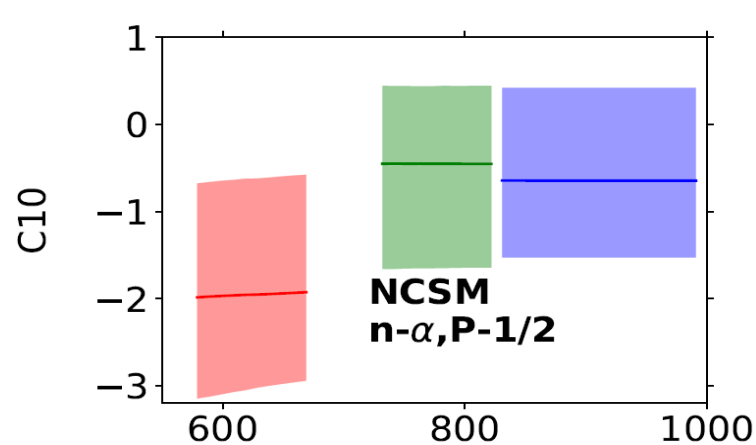
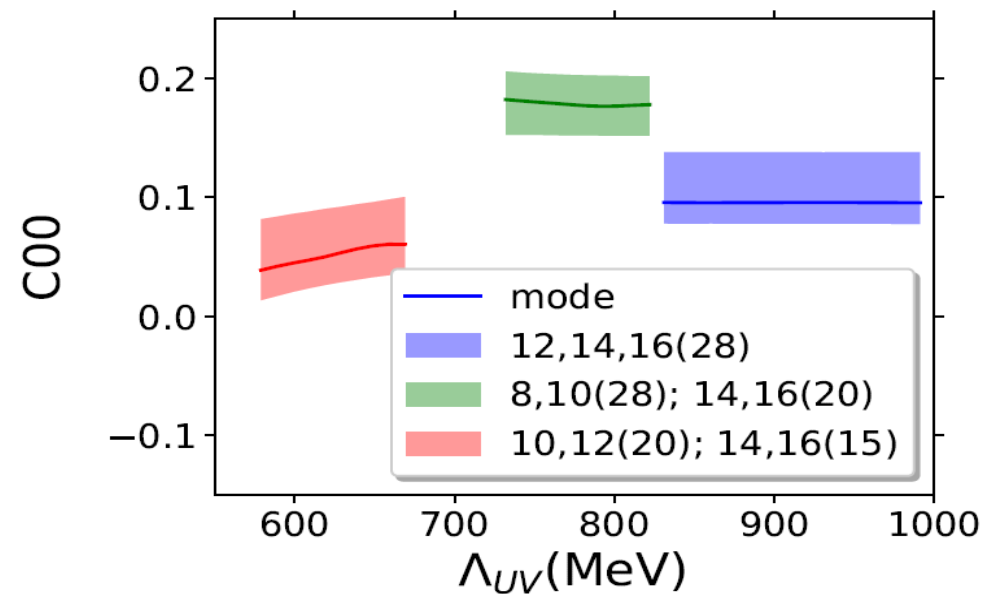
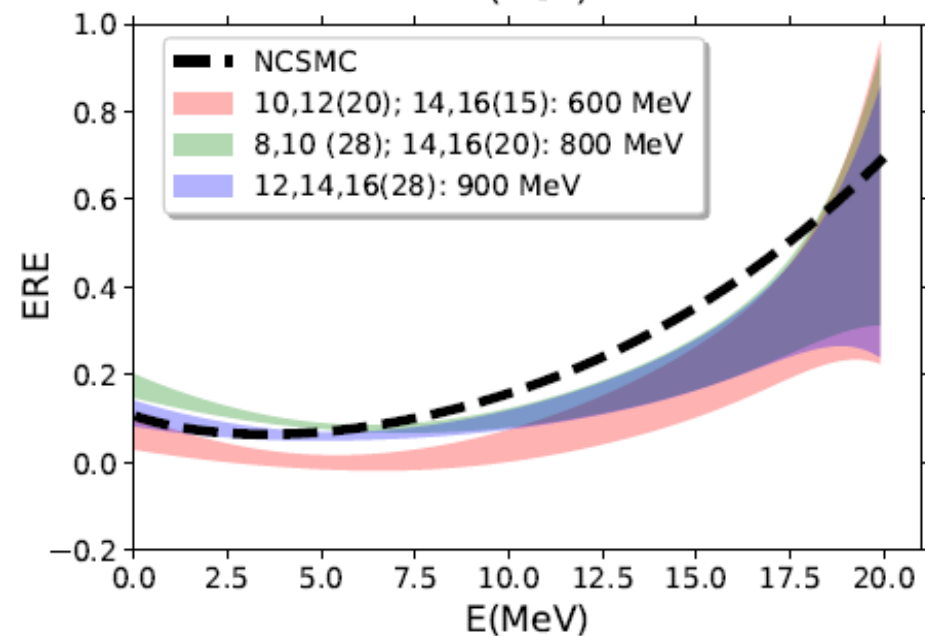
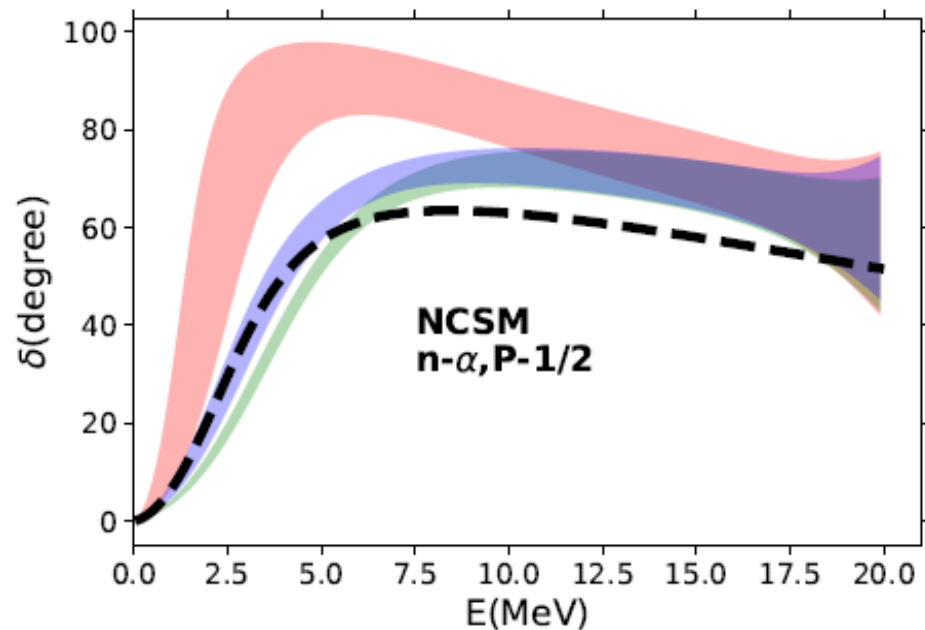


# Analyze He4 and He5 energies from NCSM and extract n- $\alpha$ scattering in P-3/2 channel



- Lowering  $\Lambda_{UV}$  has a trend to turn a resonance to bound state (also seen in p- $1/2$  channel)
- The extraction agrees with Petr's direct phase-shift calculation below 5 MeV with high  $\Lambda_{UV}$
- Since we model both UV and IR physics components, we can use most (Nmax,  $\omega$ ) results and extract phase-shifts. This is like LQCD producing results at different Lattice spacing.

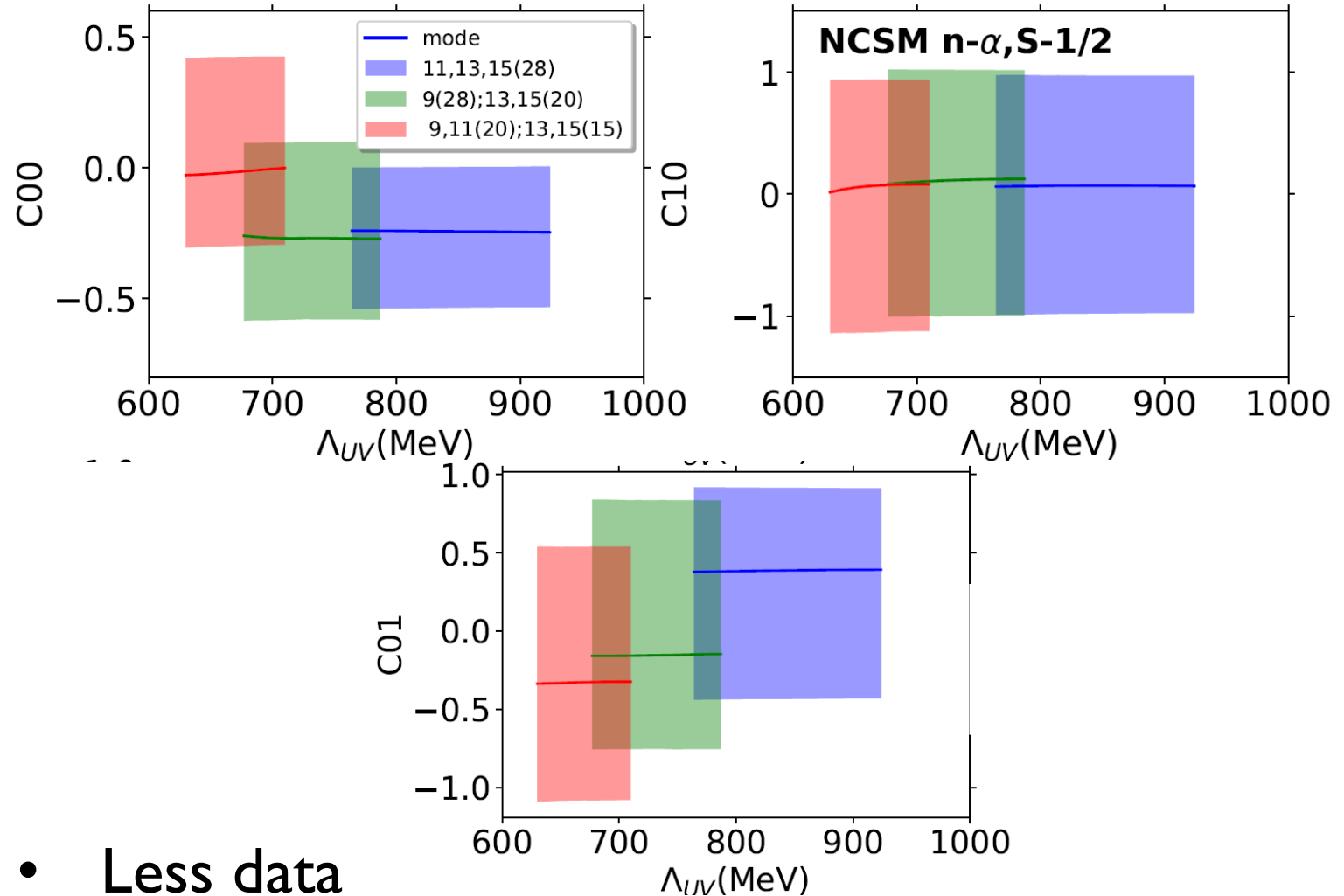
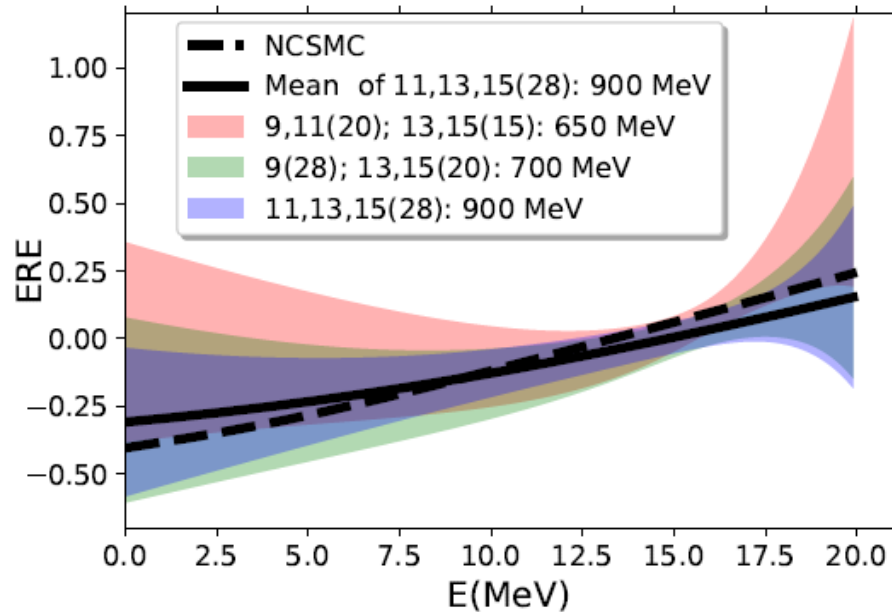
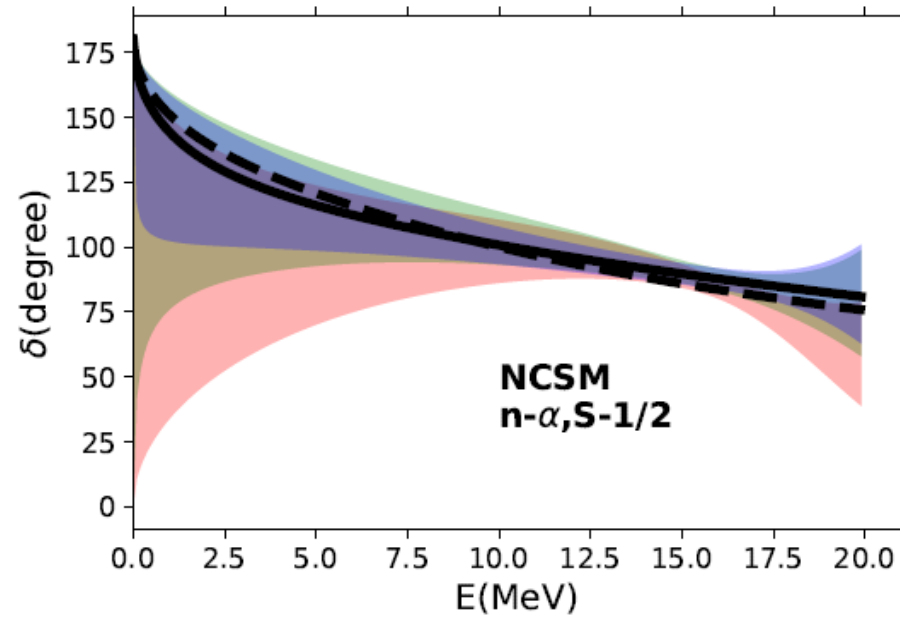
# n- $\alpha$ in P-1/2 channel from NCSM



Similar conclusions can be drawn as in P-3/2

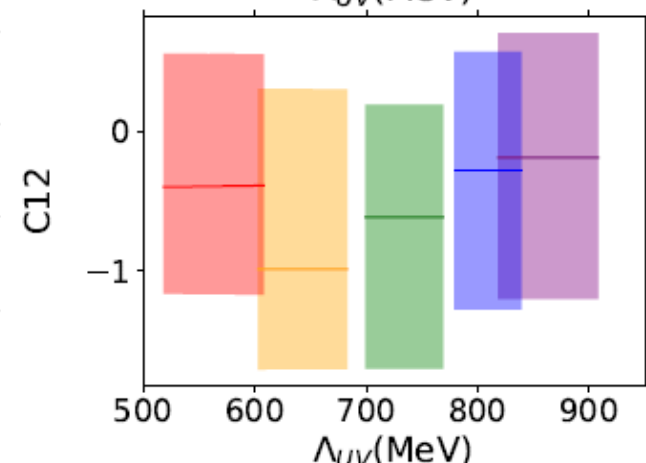
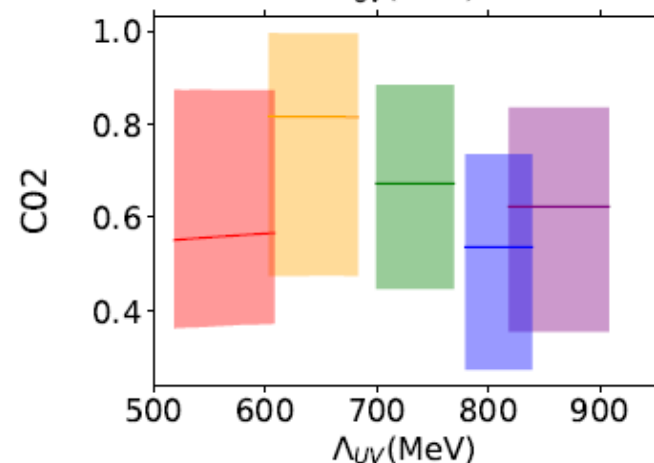
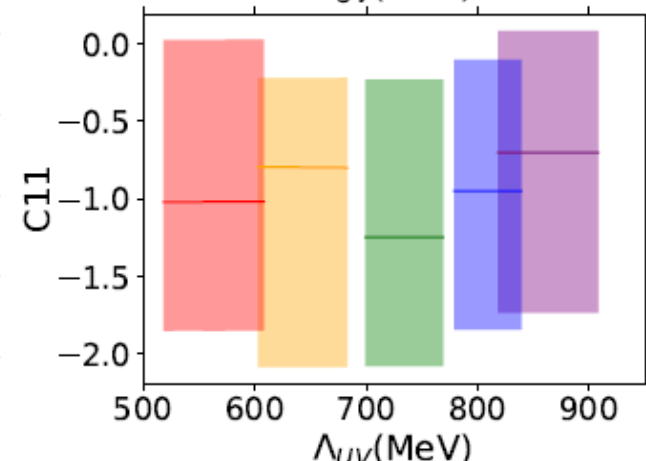
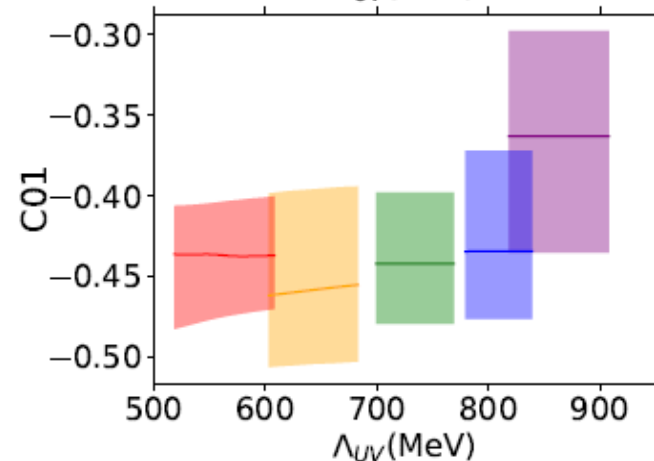
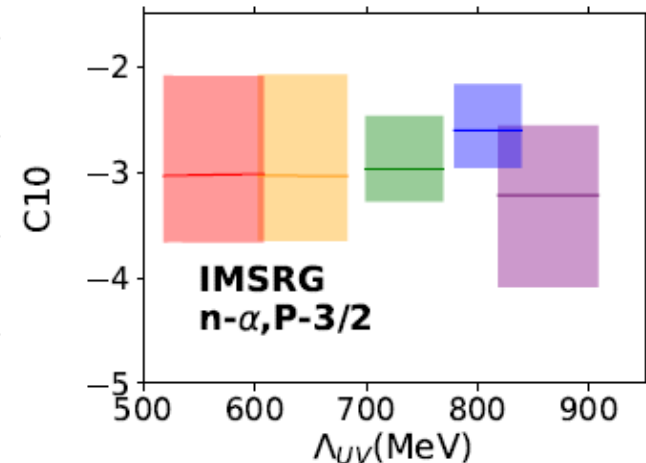
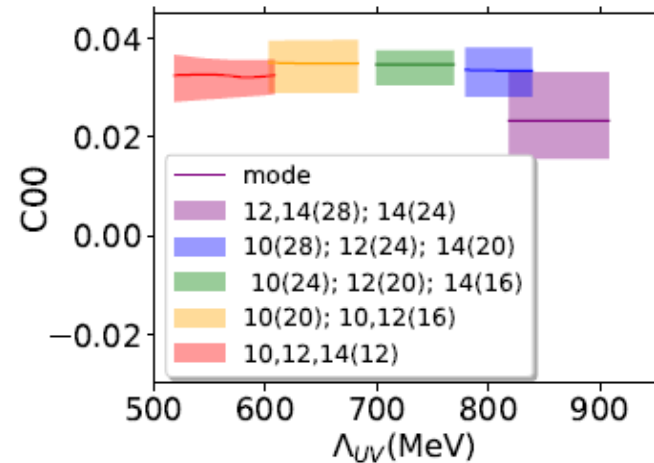
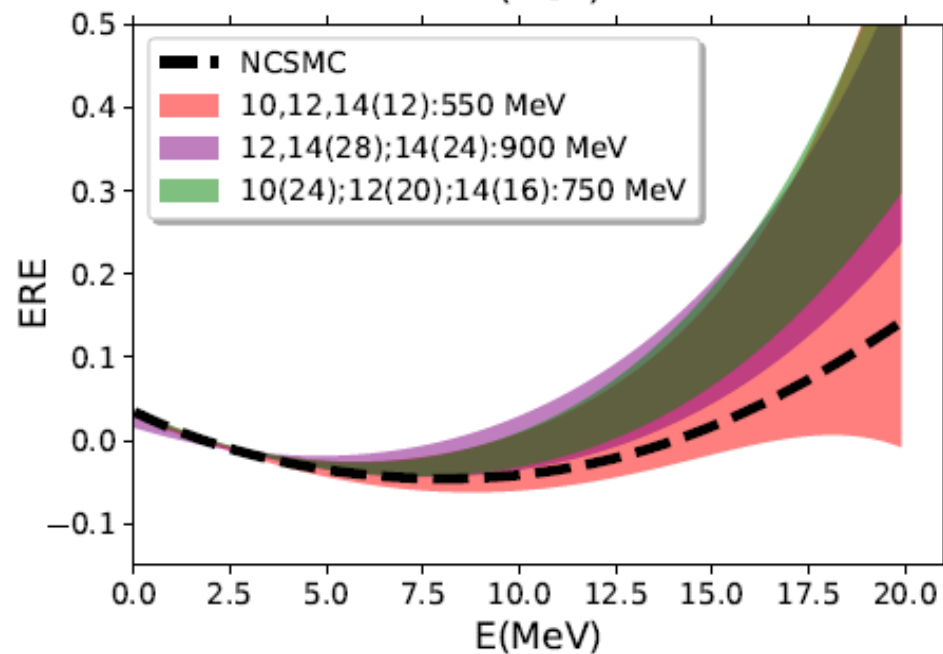
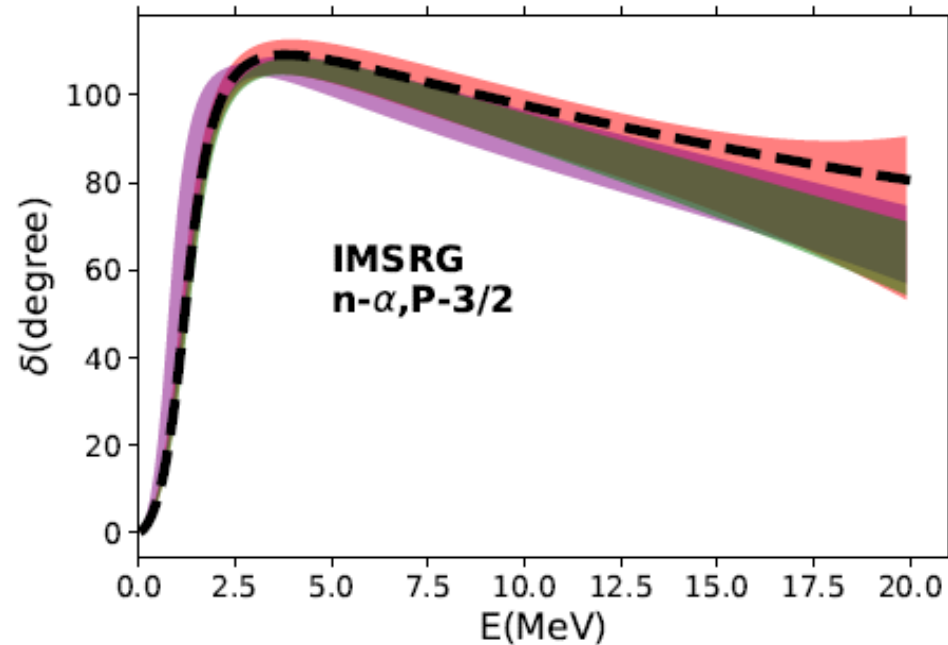


# n- $\alpha$ in S-1/2 channel from NCSM



- Less data
- Mean values agree with NCSMC quite well

# n- $\alpha$ in P-3/2 channel from IMSRG



# n- $\alpha$ in P-1/2 channel from IMSRG

